

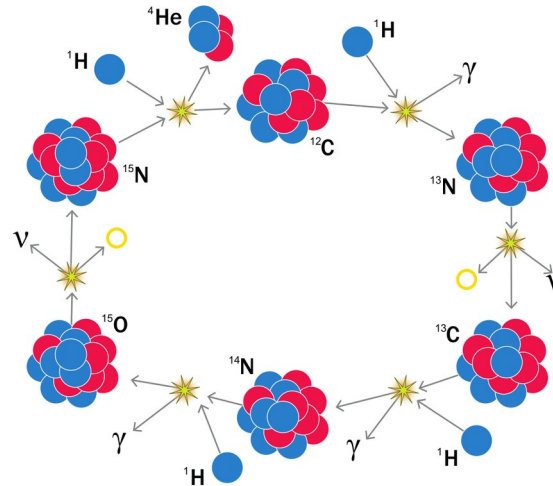
Experimental detection of the CNO cycle

Nicola Rossi

Laboratori Nazionali del Gran Sasso (INFN)

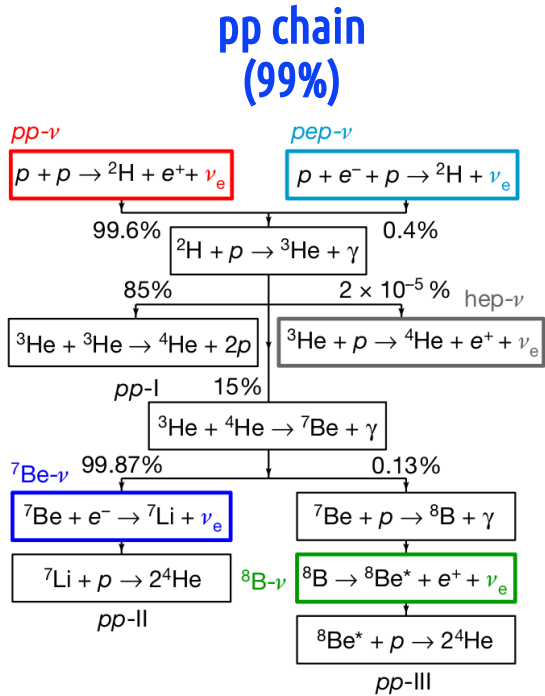


Istituto Nazionale di Fisica Nucleare
Laboratori Nazionali del Gran Sasso

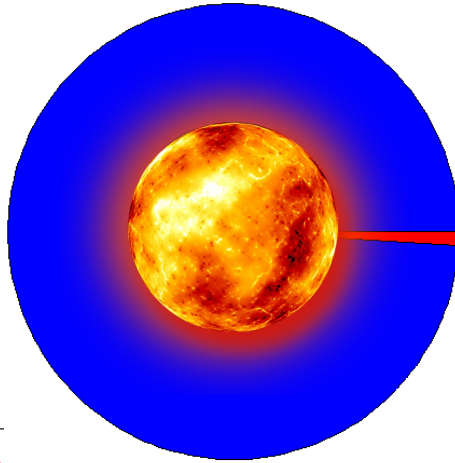


MG16  5-10 JULY 2021
SIXTEENTH MARCEL GROSSMANN MEETING
ON RECENT DEVELOPMENTS IN THEORETICAL AND EXPERIMENTAL GENERAL RELATIVITY, ASTROPHYSICS AND RELATIVISTIC FIELD THEORIES

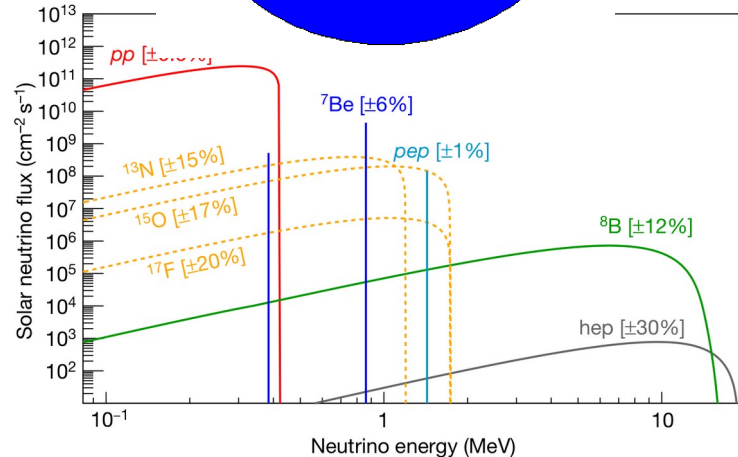
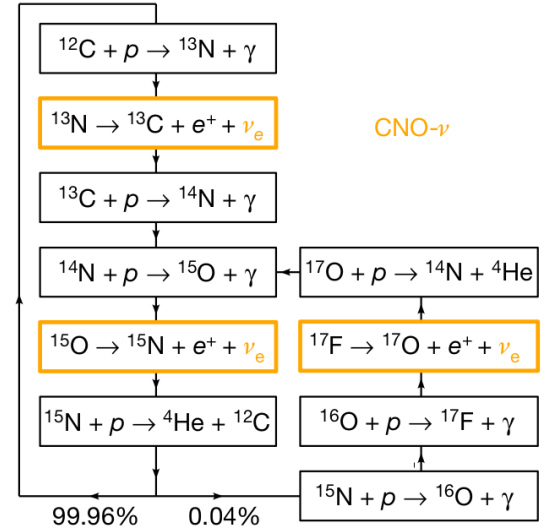
pp vs CNO Competition



Dominant in the sun
 $T_{\text{core}} \sim 15 \times 10^6 \text{ K}$



CNO cycle (1%)



Solar neutrinos: astro & particle physics

- dominant in stars 1.3 heavier than sun
 - crucial for the solar *metallicity problem*

Neutrinos from the sun

Solar luminosity

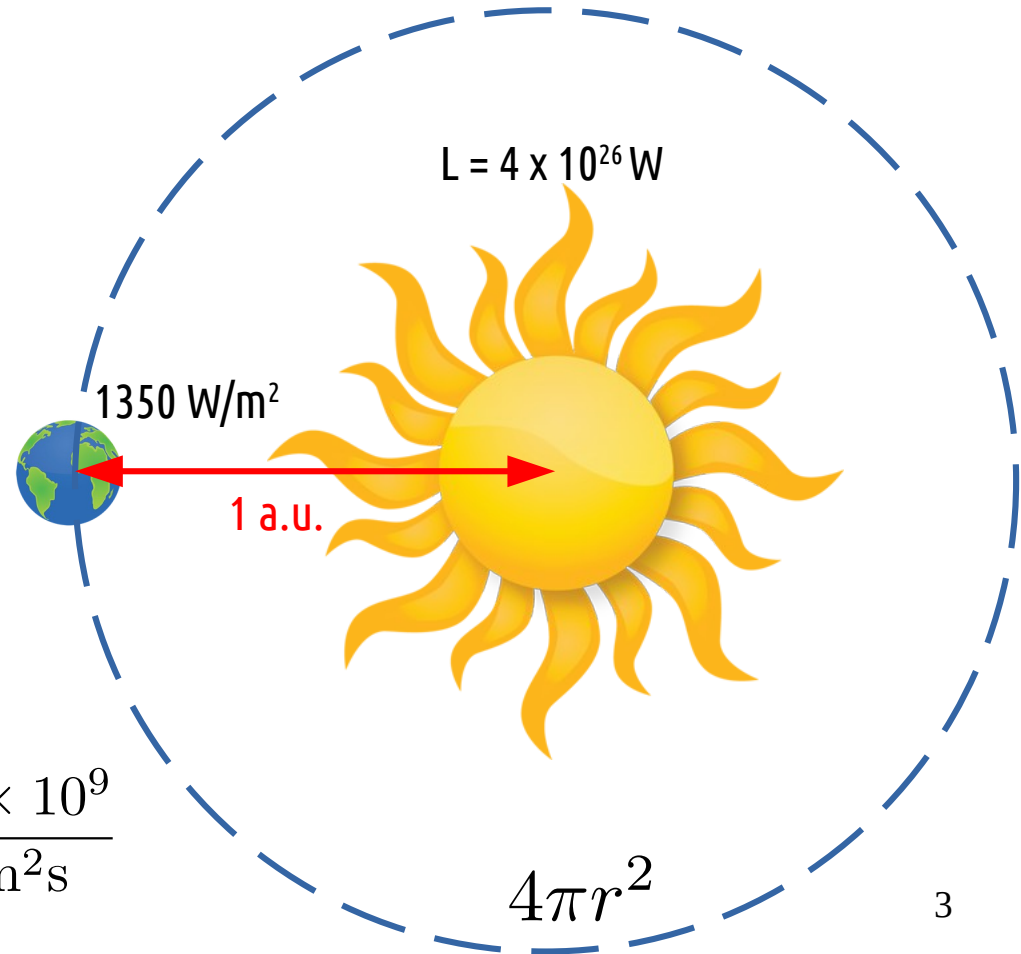
$$L_{\odot} \approx 4 \times 10^{26} \text{ W}$$

Reaction rate

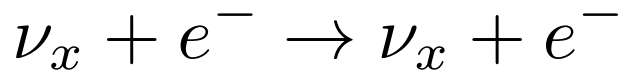
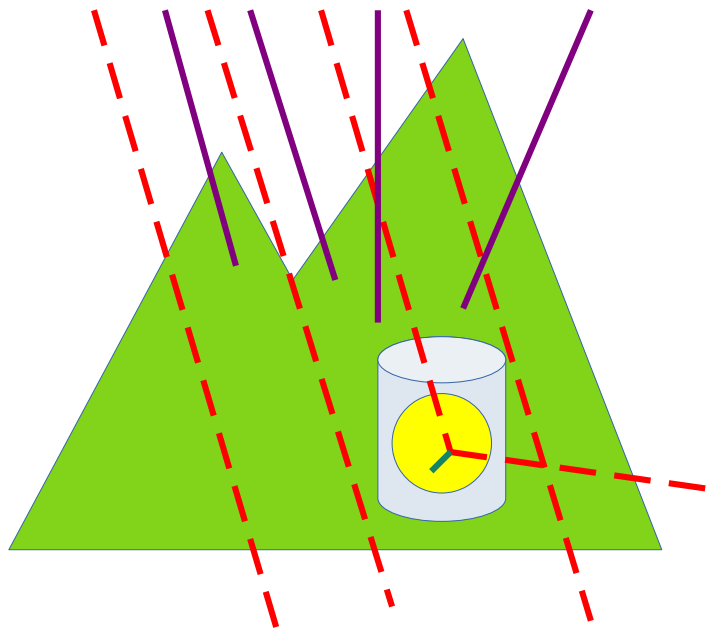
$$R = \frac{L_{\odot}}{\Delta M c^2} \approx \frac{2.4 \times 10^{39} \text{ MeV/s}}{26 \text{ MeV}} \approx \frac{10^{38}}{\text{s}}$$

Neutrino flux on the earth

$$\Phi_{\nu}(\text{Earth}) = 2 \frac{R}{4\pi d^2} \approx 2 \frac{10^{38}}{4\pi (\text{a.u.})^2} \approx \frac{65 \times 10^9}{\text{cm}^2 \text{ s}}$$



Building a solar neutrino detector



Ingredients:

- Underground
- High radio-purity
- Large mass

An example

$$\Phi_\nu(^7\text{Be}) \approx 0.5 \times 10^9 \text{ cm}^{-2} \text{ s}^{-1}$$

Targets in 100 ton of C_9H_{12}

$$\mathcal{N}_e \approx 3.3 \times 10^{31}$$

Average cross section

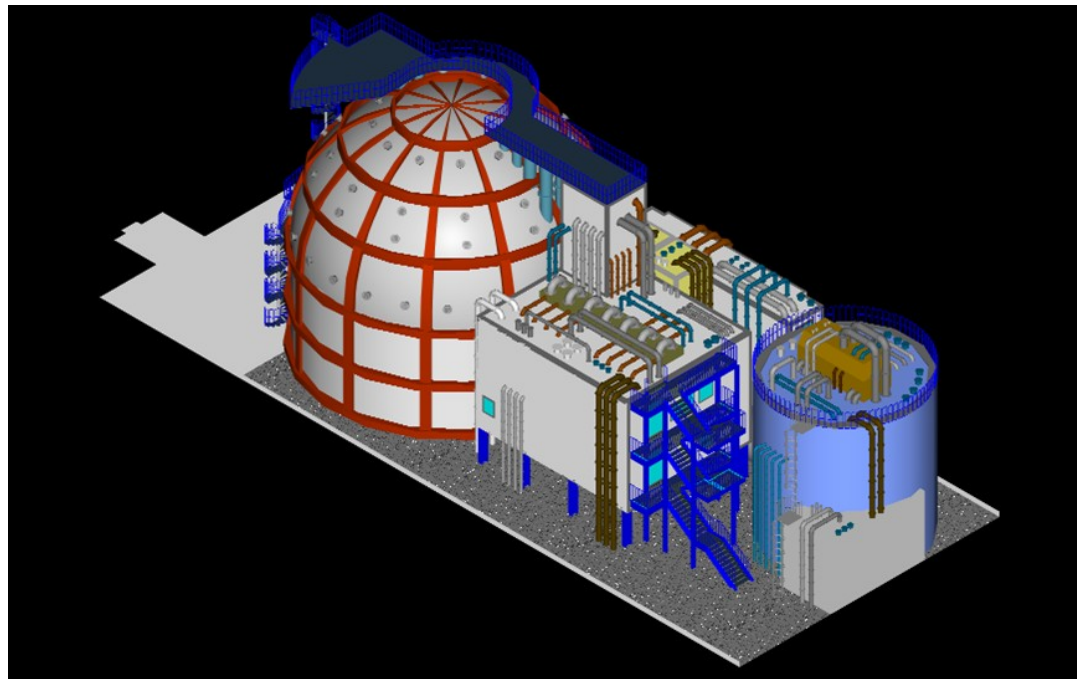
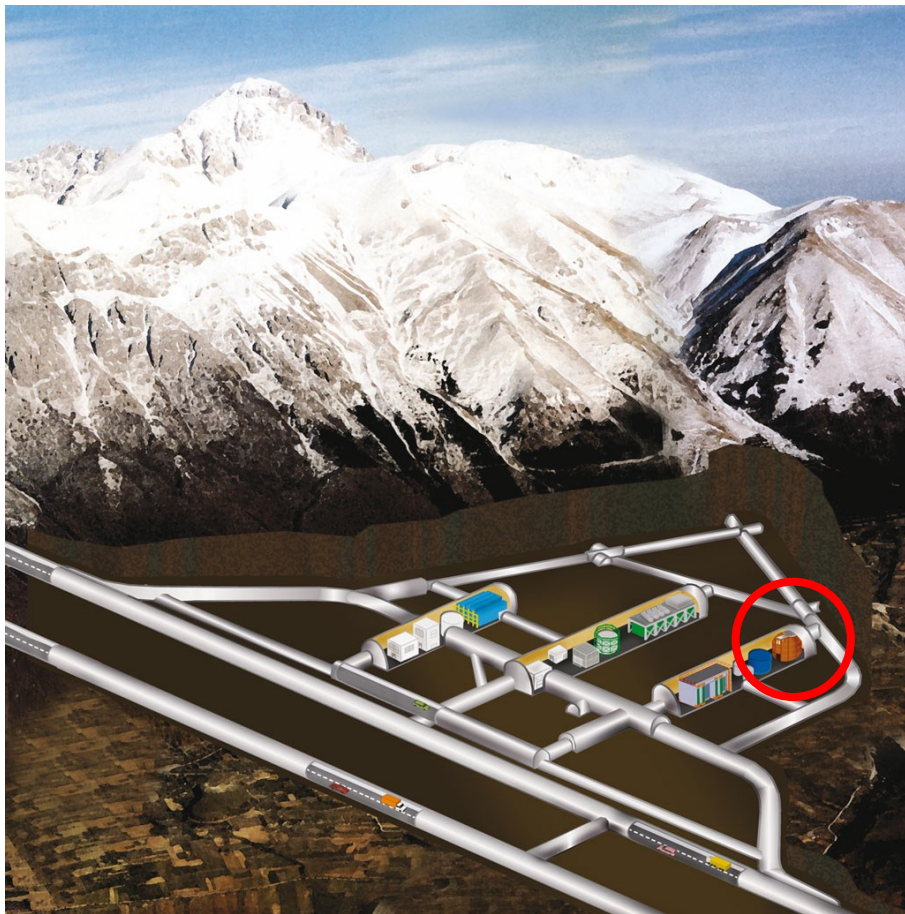
$$\langle \sigma_{\nu e} \rangle \approx 3.4 \times 10^{-45} \text{ cm}^2$$

Interaction rate

$$\mathcal{R} = \mathcal{N}_e \Phi_\nu \langle \sigma_{\nu e} \rangle \approx 48 \text{ cpd}/100\text{t}$$

1 cpd/100t ~ 0.1 nBq/kg !!!

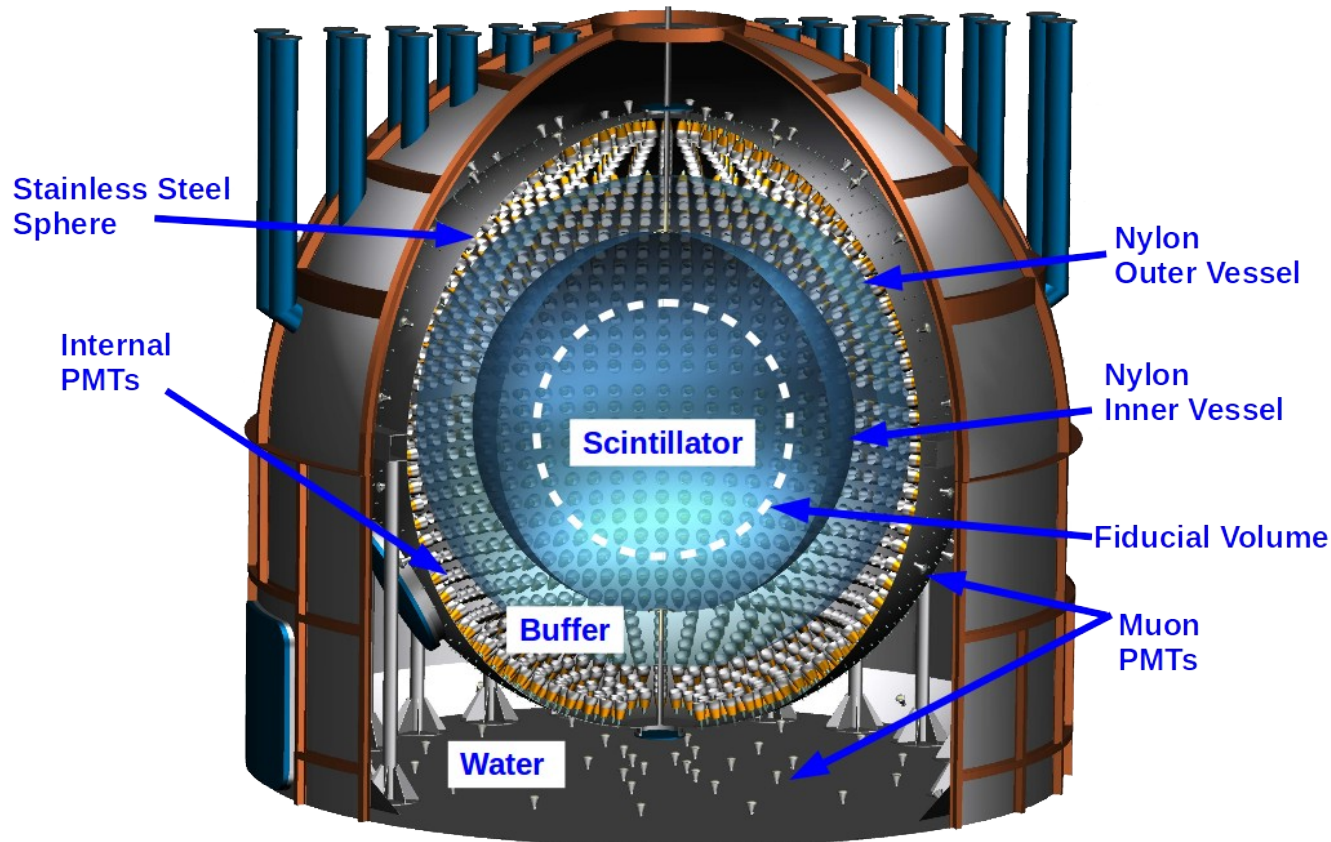
The BOREXINO detector



Laboratori Nazionali del Gran Sasso – INFN (Hall C)

Rock: 3.800 m w.e. – muon flux $\sim 1 \text{ m}^{-2}\text{h}^{-1}$

The Borexino detector



Sphere of diameter
13.7 m ($\sim 1300 \text{ m}^3$)

Target (Inner vessel): ~ 300 ton
of liquid scintillator

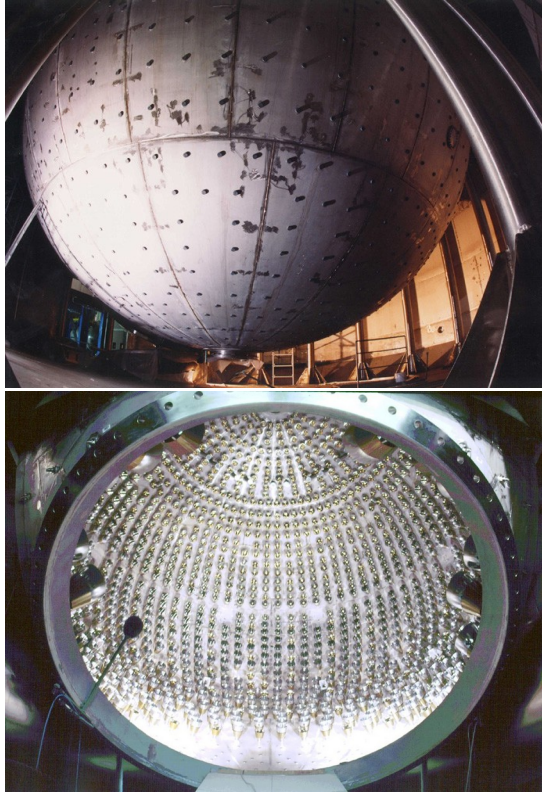
Nylon balloon $125 \mu\text{m}$

2212 PMTs (nominal)

$\sim 35\%$ optical coverage

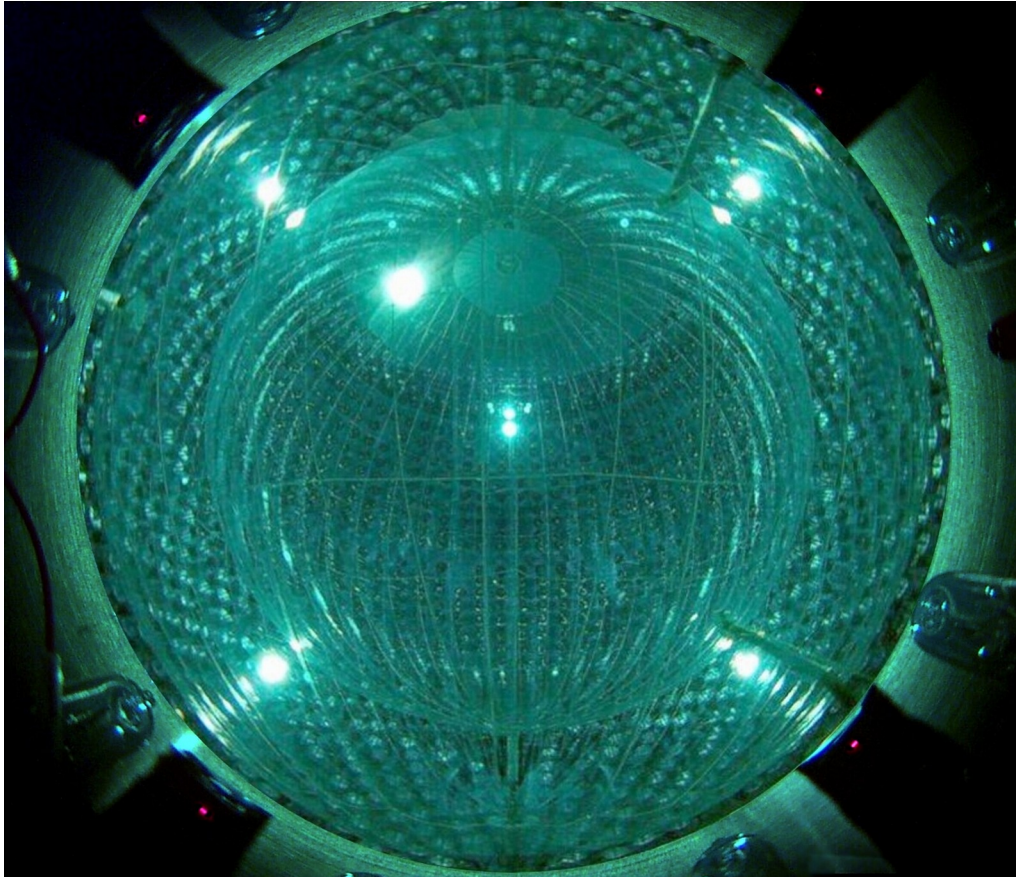
Water Cherenkov Veto
($d = 18 \text{ m}$) 2000 m^3 (208 PMT)

Borexino's pictures



Under construction, before the filling

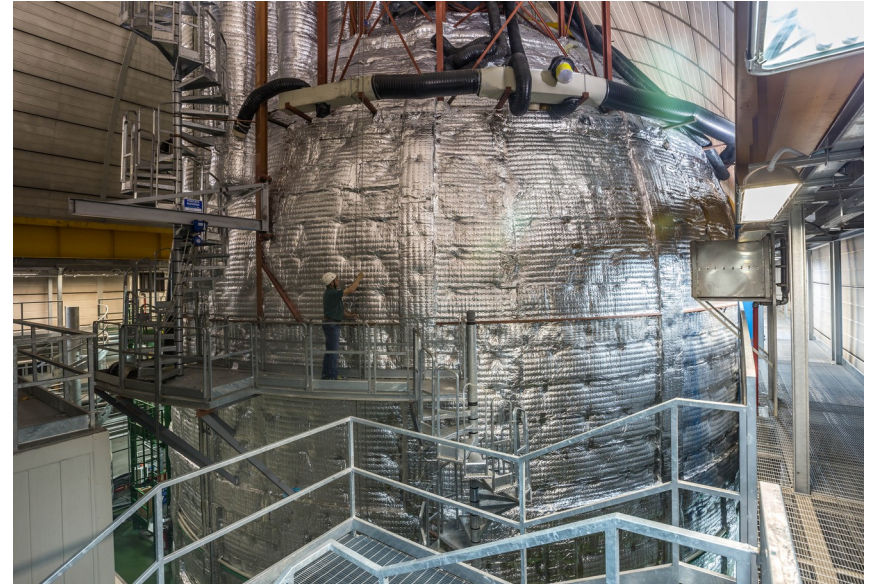
Borexino today



From monitoring
camera

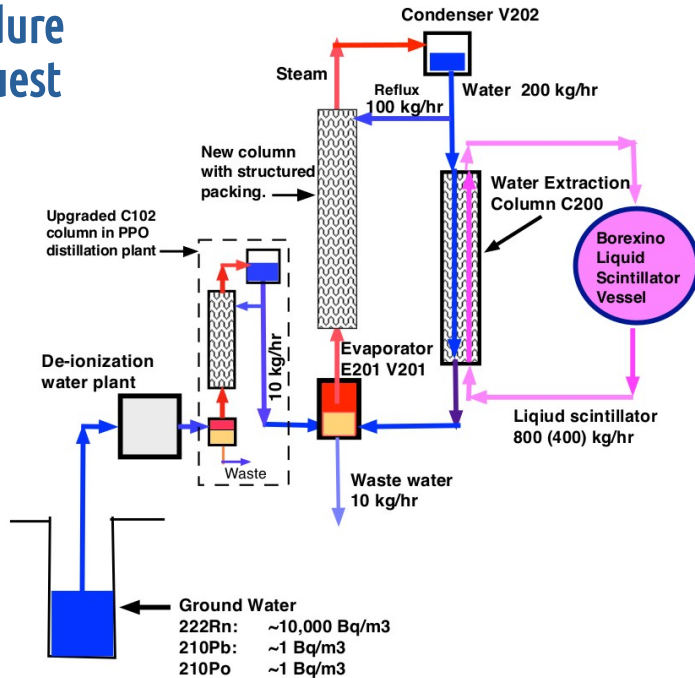
Software reconstruction
of the inner vessel

Hall C



The water extraction campaign

Key procedure
for CNO quest



Six cycles of water extraction
from mid-2010 to mid-2011

[1 cpd/100t \sim 0.1 nBq/kg]

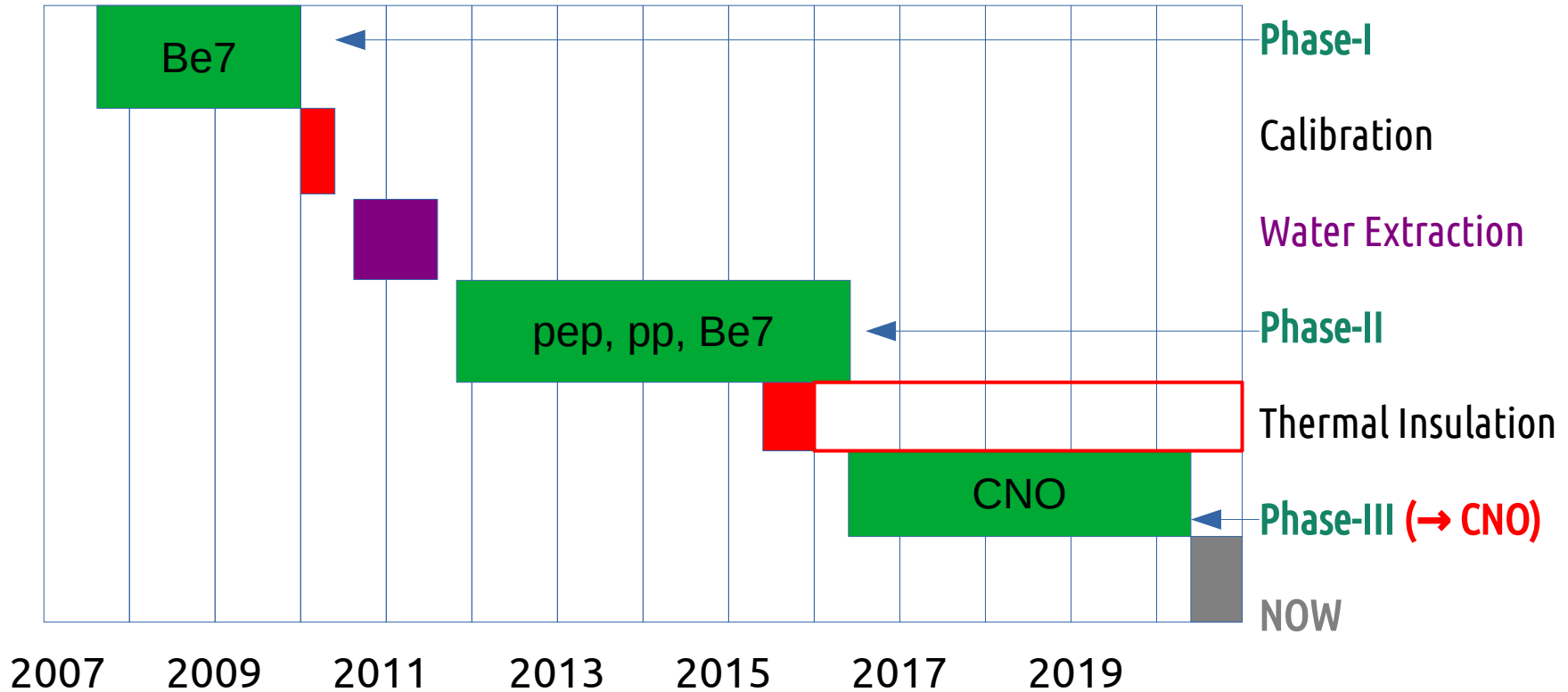
	Before [cpd/100t]	After [cpd/100t]
^{210}Bi	~ 40	~ 10
^{85}Kr	~ 30	~ 5
^{210}Po	>2000	<30 (decay)

^{238}U , $^{232}\text{Th} < 10^{-19}$ g/g

^{39}Ar , $^{40}\text{K} \ll 1$ cpd/100t

Expected CNO ~ 5 cpd/100t

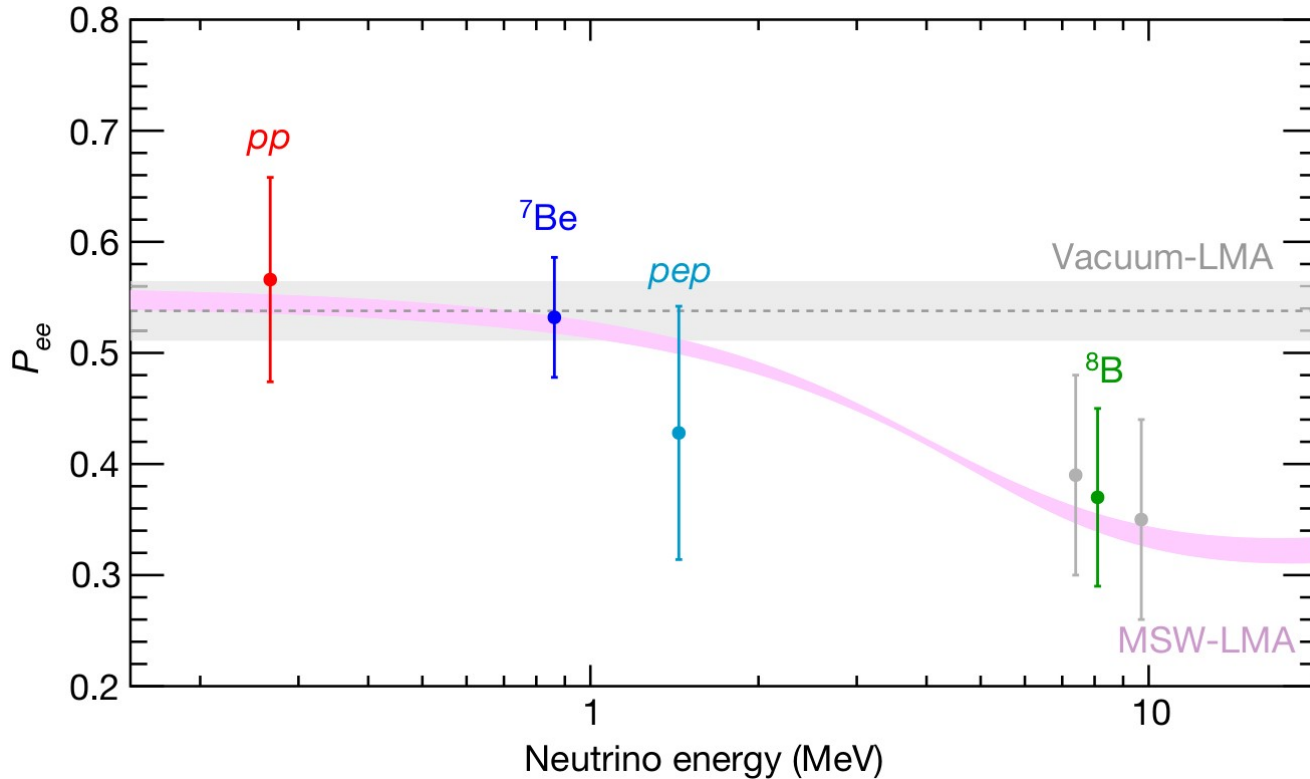
The three phases of Borexino



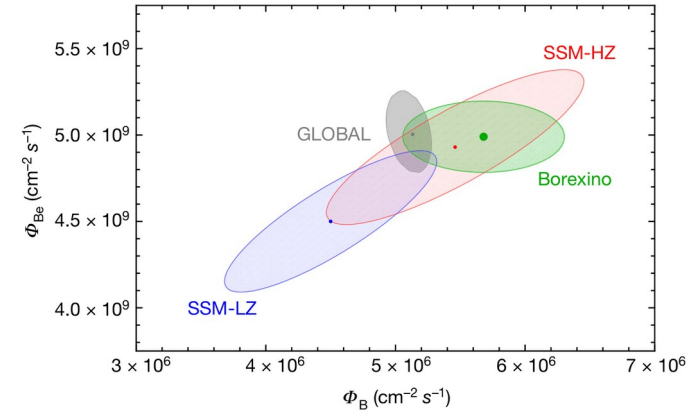
Results from Borexino (2018)

Neutrinos	References	Rate [cpd/100t]	Flux [cm ⁻² s ⁻¹]
pp	Nature 2014, Nature 2018, PRD 2019	$(134 \pm 10)_{-10}^{+6}$	$(6.1 \pm 0.5)_{-0.5}^{+0.3} \times 10^{10}$
⁷ Be	PLB 2008, PRL 2011, Nature 2018, PRD 2019	$(48.3 \pm 1.1)_{-0.7}^{+0.4}$	$(4.99 \pm 0.11)_{-0.08}^{+0.06} \times 10^9$
pep	PRL 2012, Nature 2018 PRD 2019	$(2.7 \pm 0.4)_{-0.2}^{+0.1}$	$(1.3 \pm 0.3)_{-0.1}^{+0.1} \times 10^8$
⁸ B	PRD 2010, Nature 2018, PRD 2020	$0.223_{-0.022}^{+0.021}$	$5.68_{-0.41-0.03}^{+0.39+0.03} \times 10^6$
hep	Nature 2018, PRD 2020	<0.002 (90% CL)	<1.8x10 ⁵ (90% CL)
CNO	PRL 2010, Nature 2018 (upper limit)	<8.1 (95% CL)	<7.9x10 ⁸ (95%CL)

Implications of Borexino results



P_{ee} survival probability with Borexino data only!

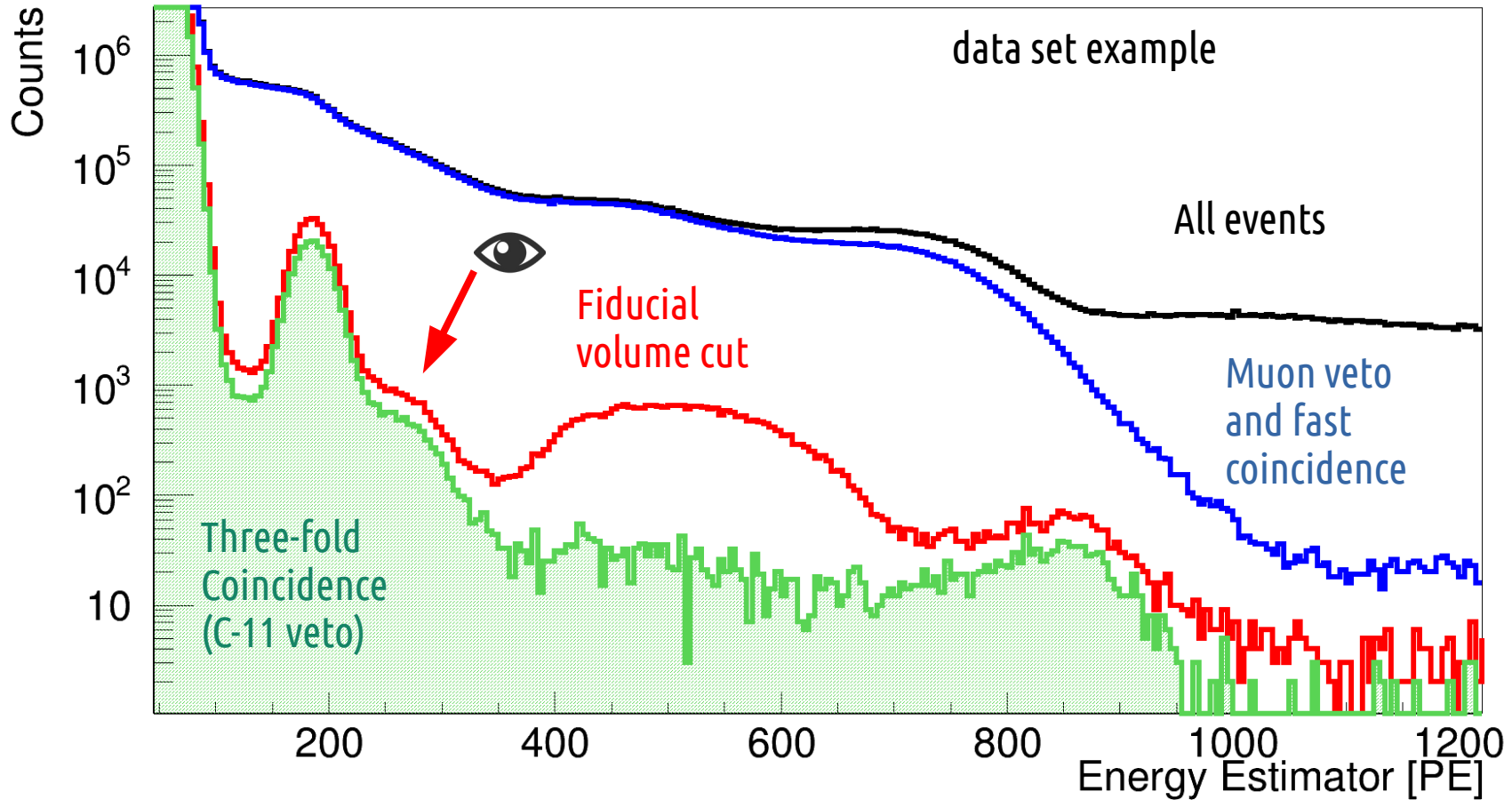


Precise measurements era:

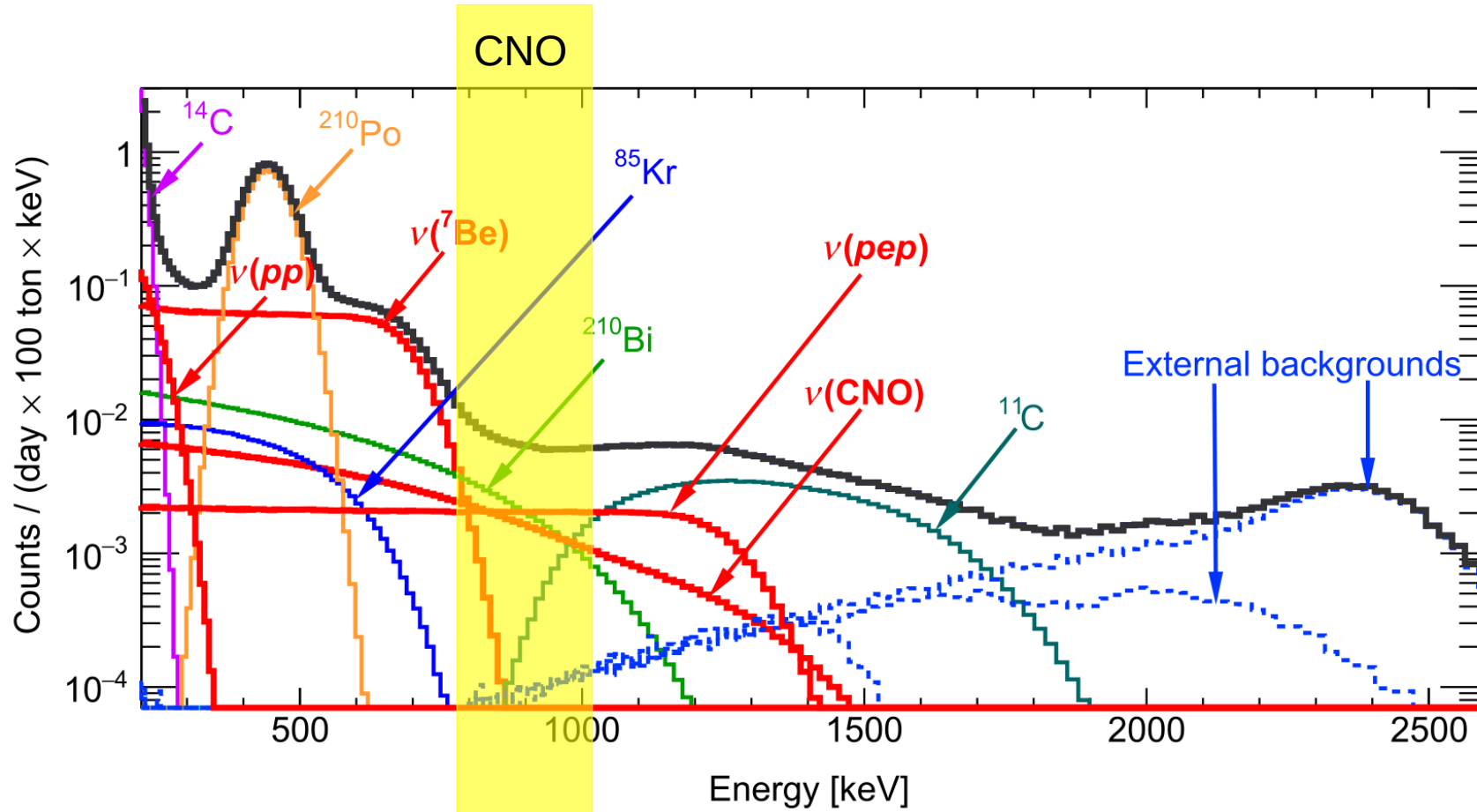
Probing the **MSW-LMA** scenario

Low metallicity disfavored at **1.8 σ**

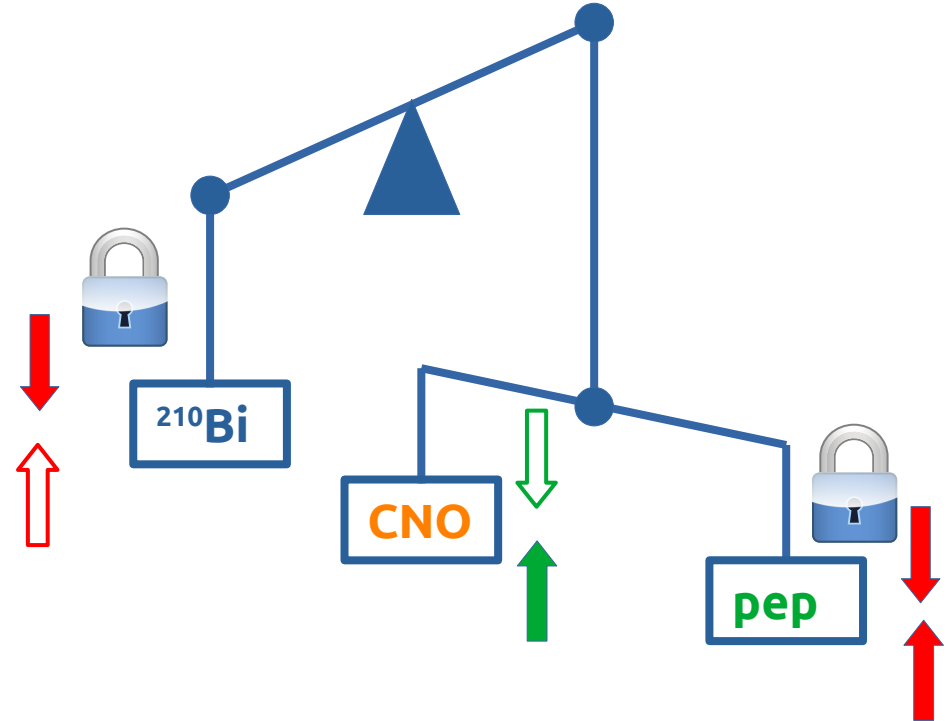
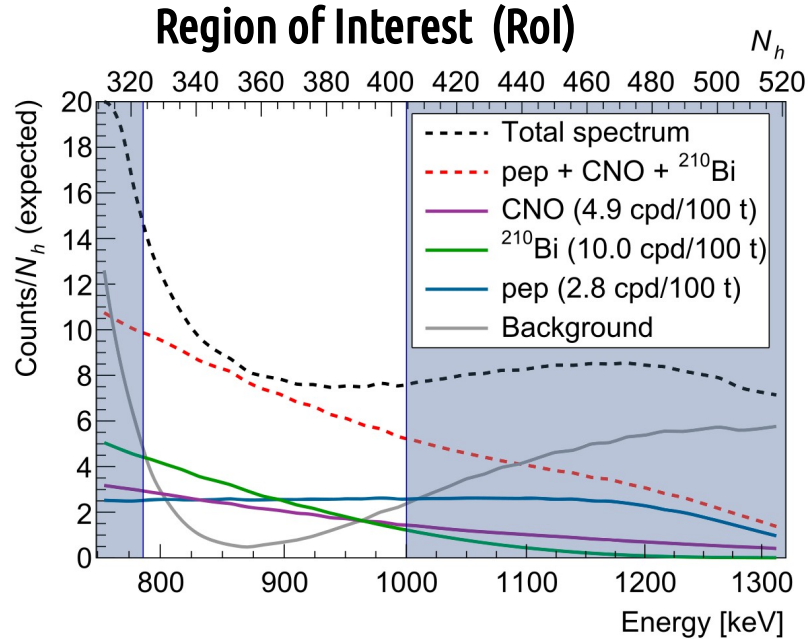
The Borexino energy spectrum



Understanding the spectrum



Sensitivity to CNO neutrinos



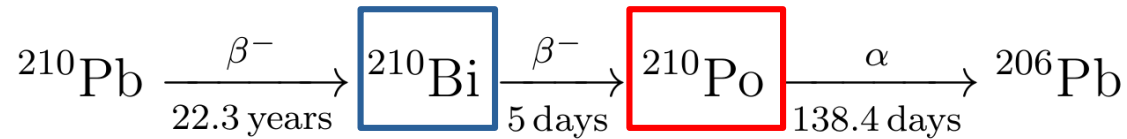
CNO \leftrightarrow pep \leftrightarrow ^{210}Bi correlation

Core of the problem: *counting analysis* in the RoI

Strategy:

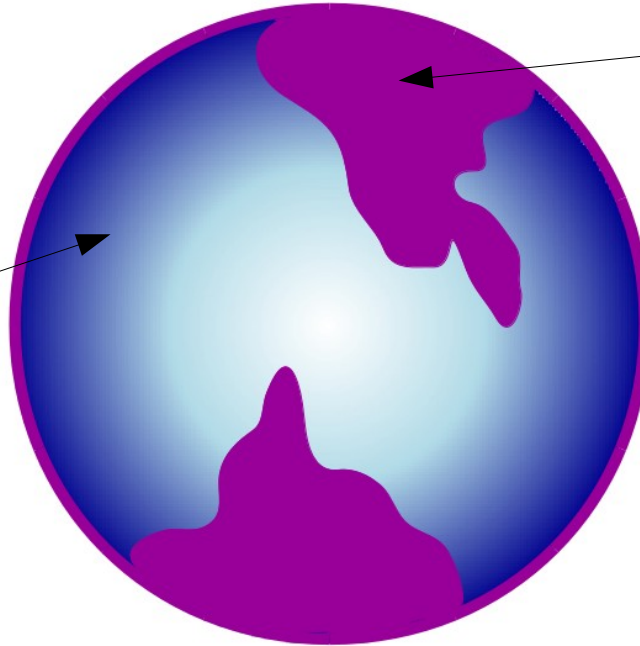
1. independent constraint of pep
2. independent constraint and ^{210}Bi

More about independent ^{210}Bi constraint



Exploiting the A=210 chain
assuming the secular
equilibrium

Diffusion:
Very slow
 $D \sim 10^{-9} \text{ m}^2/\text{s}$
(diffusion coefficient)



Convection:
 ^{210}Po from the outer
regions



Thermal Insulation program

Scenarios:

1. Plateau \rightarrow Symmetric Constraint
2. Minimum \rightarrow Upper Limit

Thermal insulation program

Idea:

Strong and stable temperature **vertical gradient** prevents convective motions

Milestones:

2014: installation of temperature probes

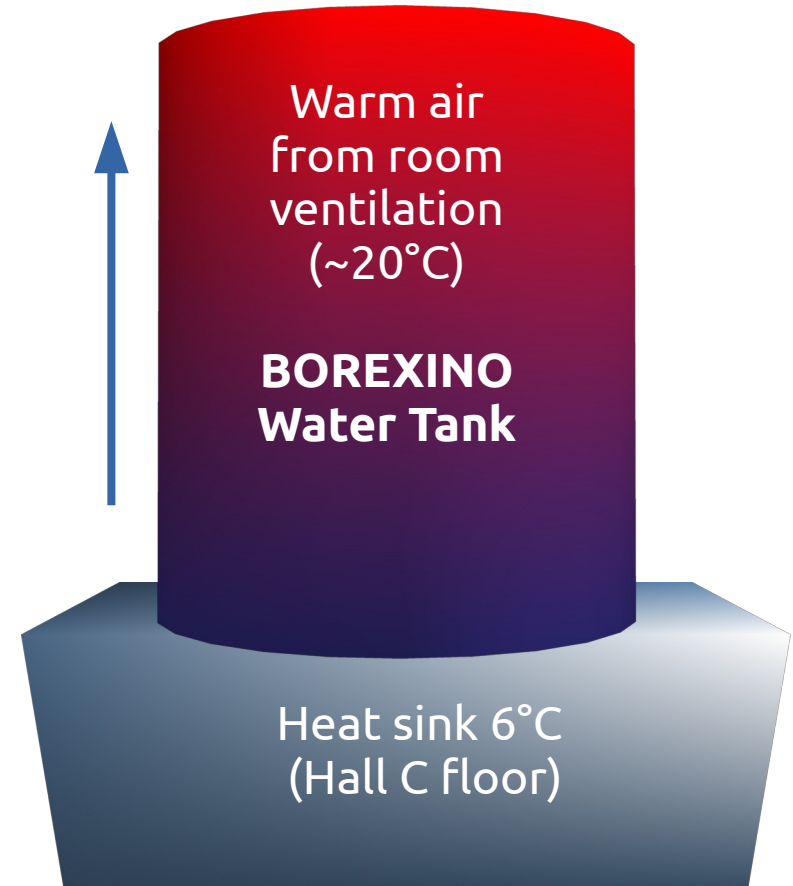
Mid-2015: *beginning* of the insulation program

Late 2015: turning off of the *water recirculation system* in the water tank

2016: first operation of the *active temperature control system* (ATCS)

Early 2019: change of the ATCS *set point*

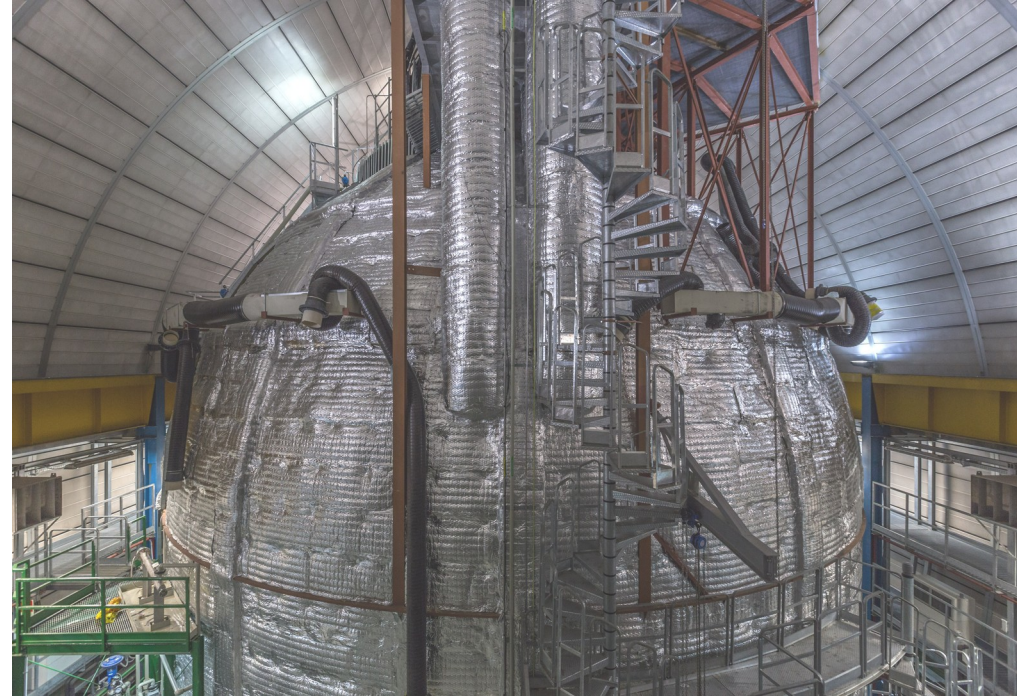
Late 2019: installation and commissioning of the *Hall C ACTS*



Thermal Insulation

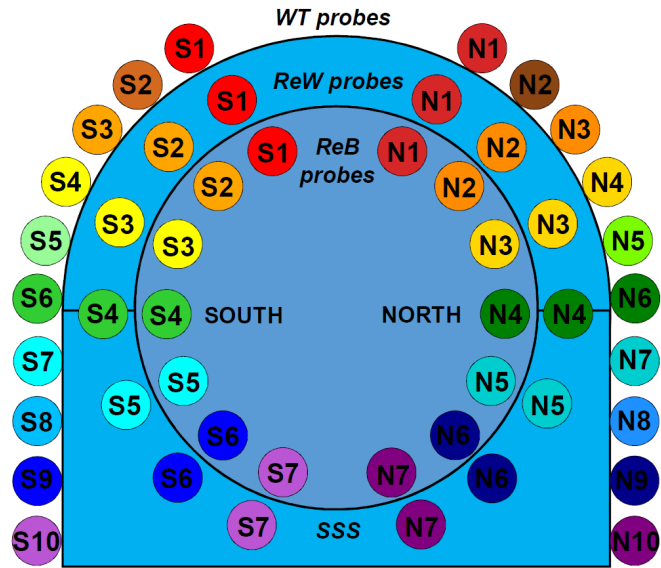


**Before the thermal insulation
(Mid-2015)**

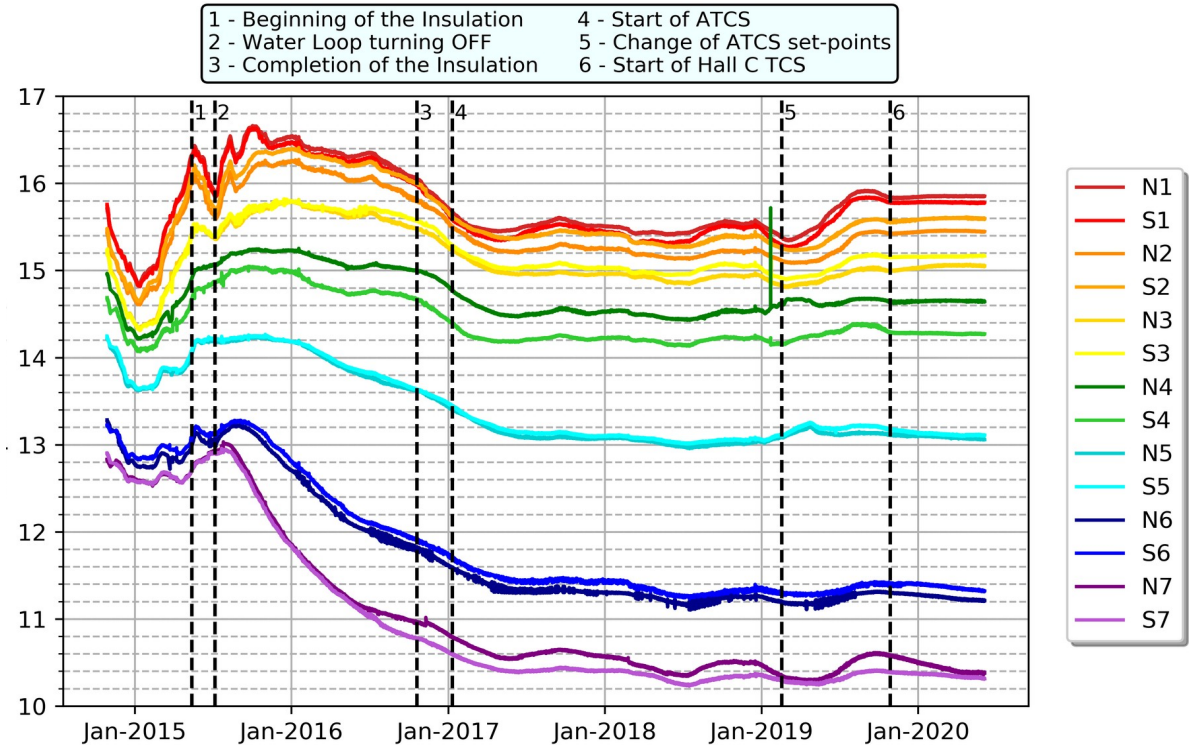


**After the thermal insulation
(Beginning of 2016)**

Effects on the temperatures

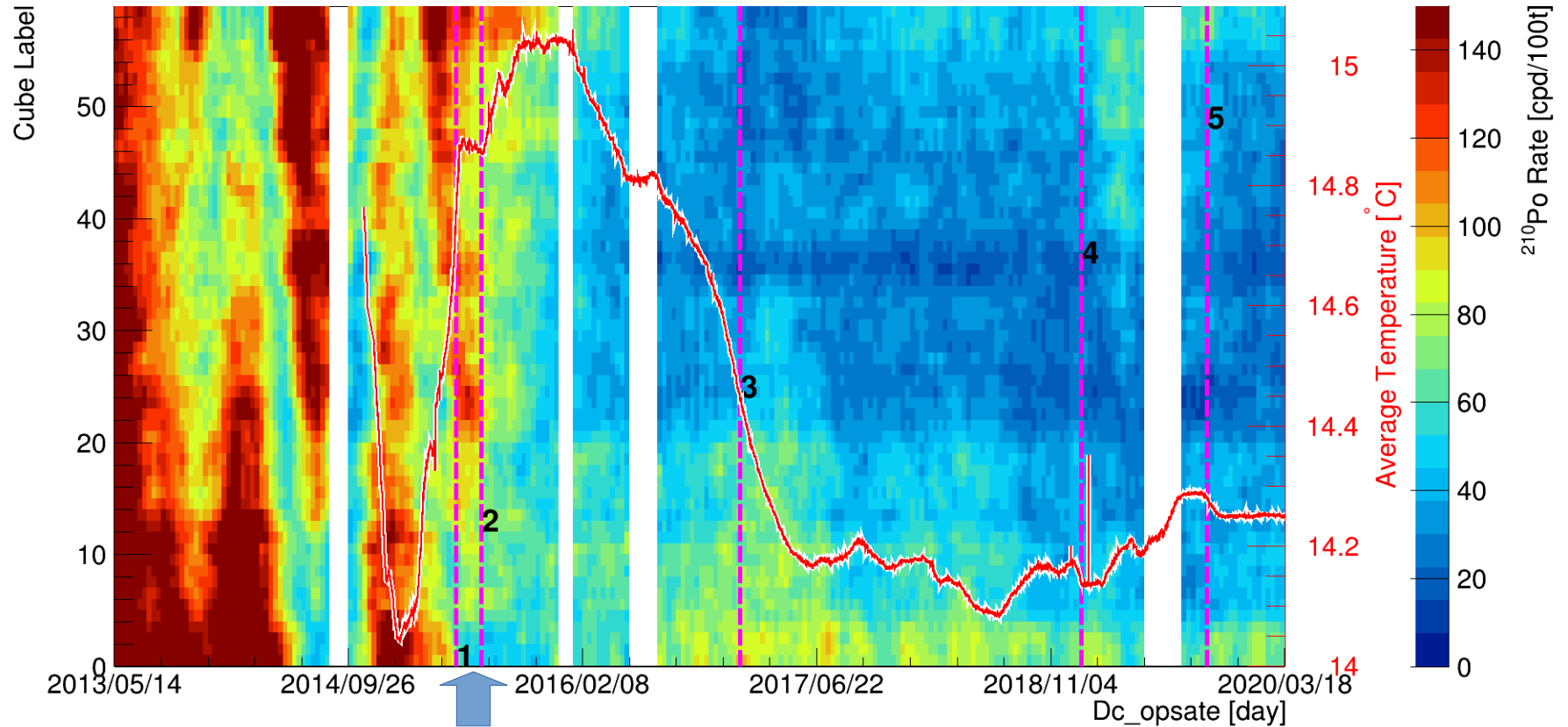


Temperature probes



Probes closer to the inner detectors
with thermal program milestones

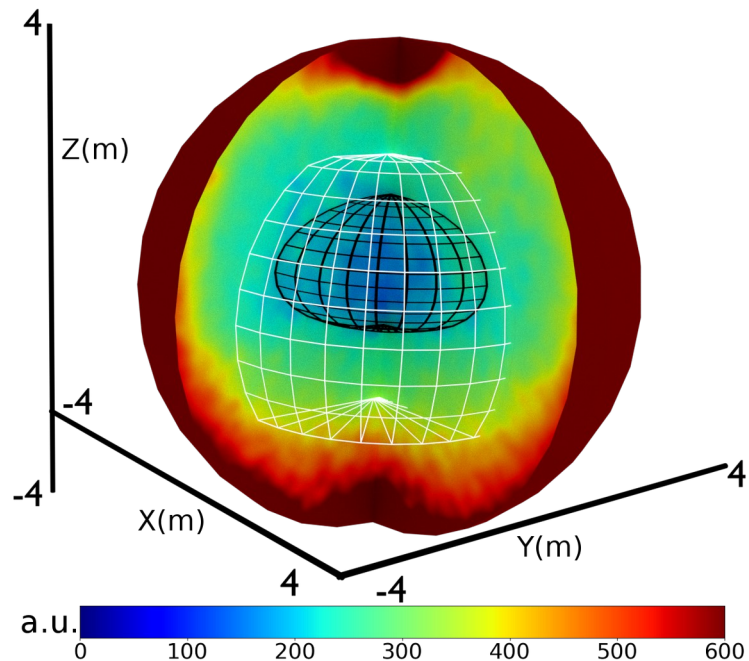
Effects on ^{210}Po



Crucial reduction of the convective motions

The low polonium field

3D View



Minimum at $z \sim 80$ cm

Compatible with numerical fluid dynamics simulations

^{210}Po activity

$$R(\text{minimum}) = R(^{210}\text{Bi}) + R(\text{Vessel})$$

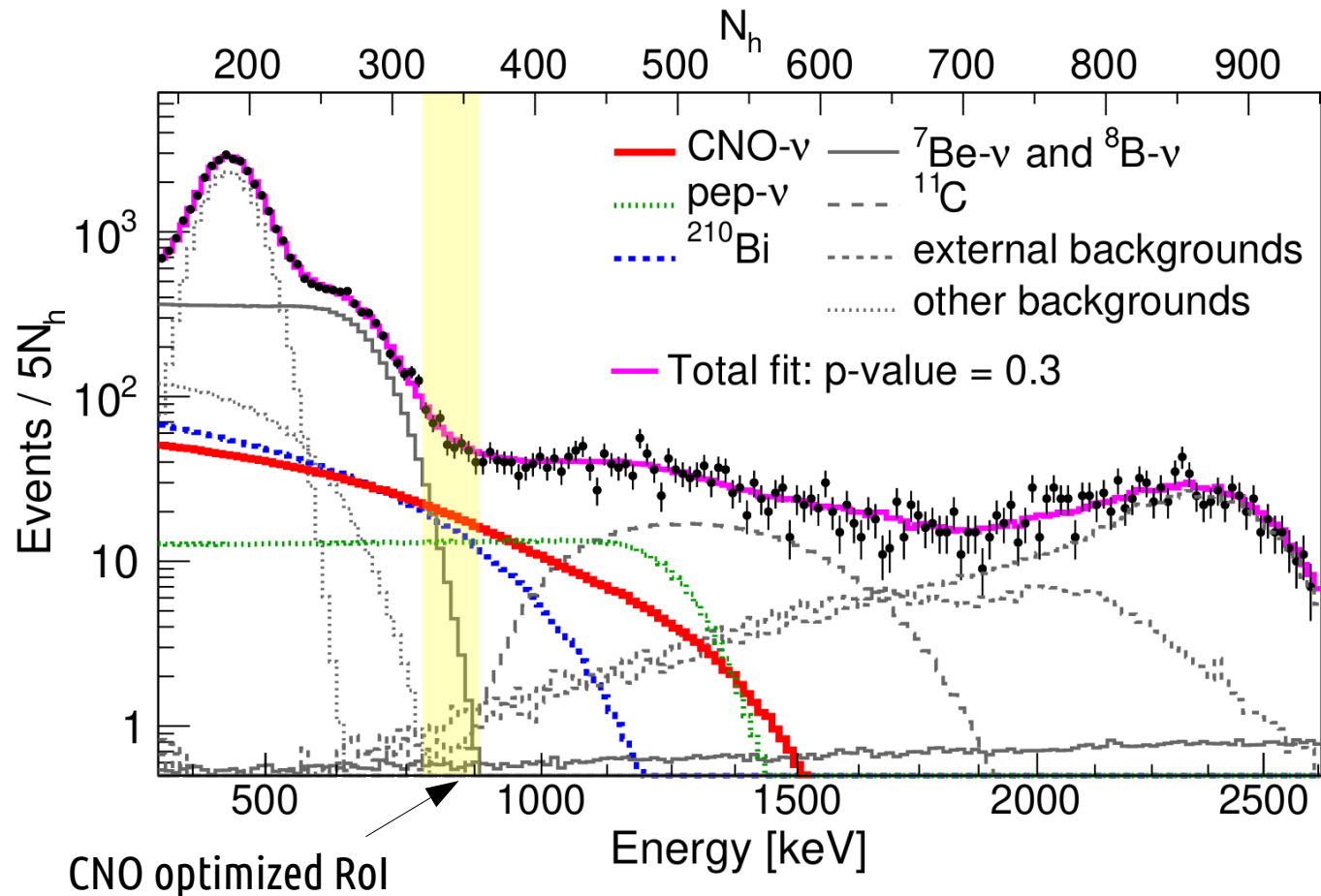
$R(\text{Vessel}) > 0 \rightarrow$ Upper limit \rightarrow lower limit for CNO (anti-correlation)

$$\text{Bi} < (11.5 \pm 1.0) \text{ cpd}/100\text{t} \text{ (stat + sys)}$$

Systematic uncertainty (uniformity): 0.8 cpd/100

$$\text{Final constraint: } ^{210}\text{Bi} < (11.5 \pm 1.3) \text{ cpd}/100\text{t}$$

CNO neutrino analysis



Main ingredients in the spectral analysis:

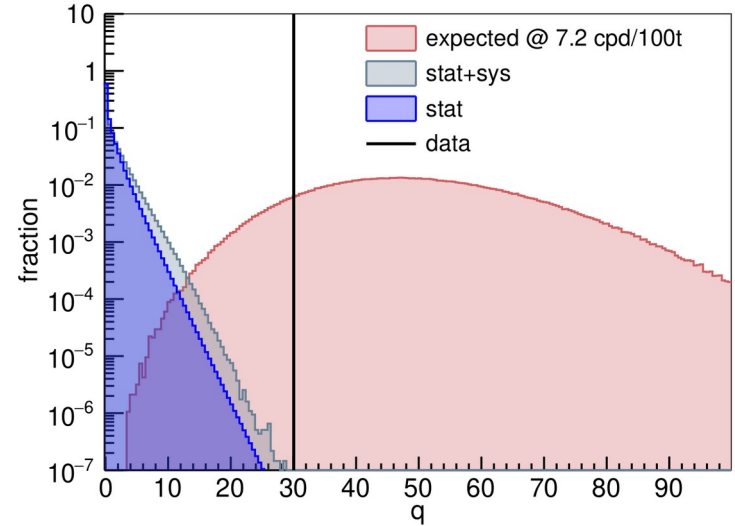
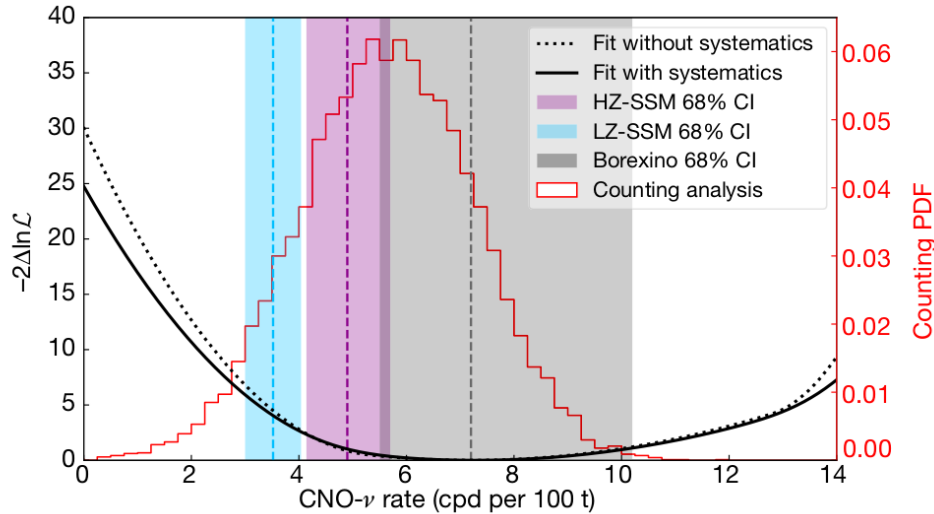
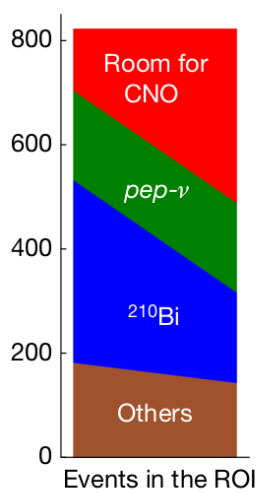
A) pep 1.4%
Symmetric penalty

B) ${}^{210}\text{Bi}$ 11%
Asymmetric penalty

1. Multivariate Montecarlo Fit

2. Counting Analysis (RoI)
Analytical modelling
(consistency check)

Final result



Systematics: Response, resolution, spectral shapes and LY: $\sigma_L = -0.5$, $\sigma_R = +0.6$ (5.1 σ significance)

Hypothesis CNO=0 excluded at 5.0 σ (99%CL) Model compatibility: 0.5 σ (HZ), 1.3 σ (LZ)

Result (68% CL stat + sys) = $7.2_{-1.7}^{+3.0}$ cpd/100t

LZ disfavored at 2.1 σ including other fluxes from pp-chain (Borexino only)

Updating the table with CNO

Neutrinos	References	Rate [cpd/100t]	Flux [cm ⁻² s ⁻¹]
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hep	Nature 2018, PRD 2020	<0.002 (90% CL)	<1.8 × 10 ⁵ (90% CL)
CNO	Nature 2020 (THIS WORK)	$7.2_{-1.7}^{+3.0}$	$7.0_{-2.0}^{+3.0} \times 10^8$

Thank you very much!



G. & V. Cocconi Prize
2021

