



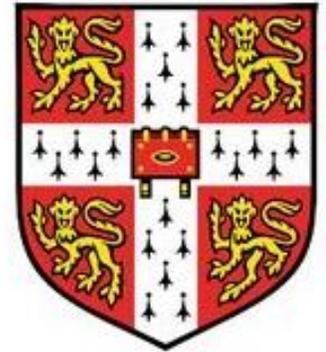
**FIAS** Frankfurt Institute  
for Advanced Studies



# Quantifying the $S_8$ tension: BA & ML

**David Benisty**

Phys.Dark Univ. 31 (2021) 100766

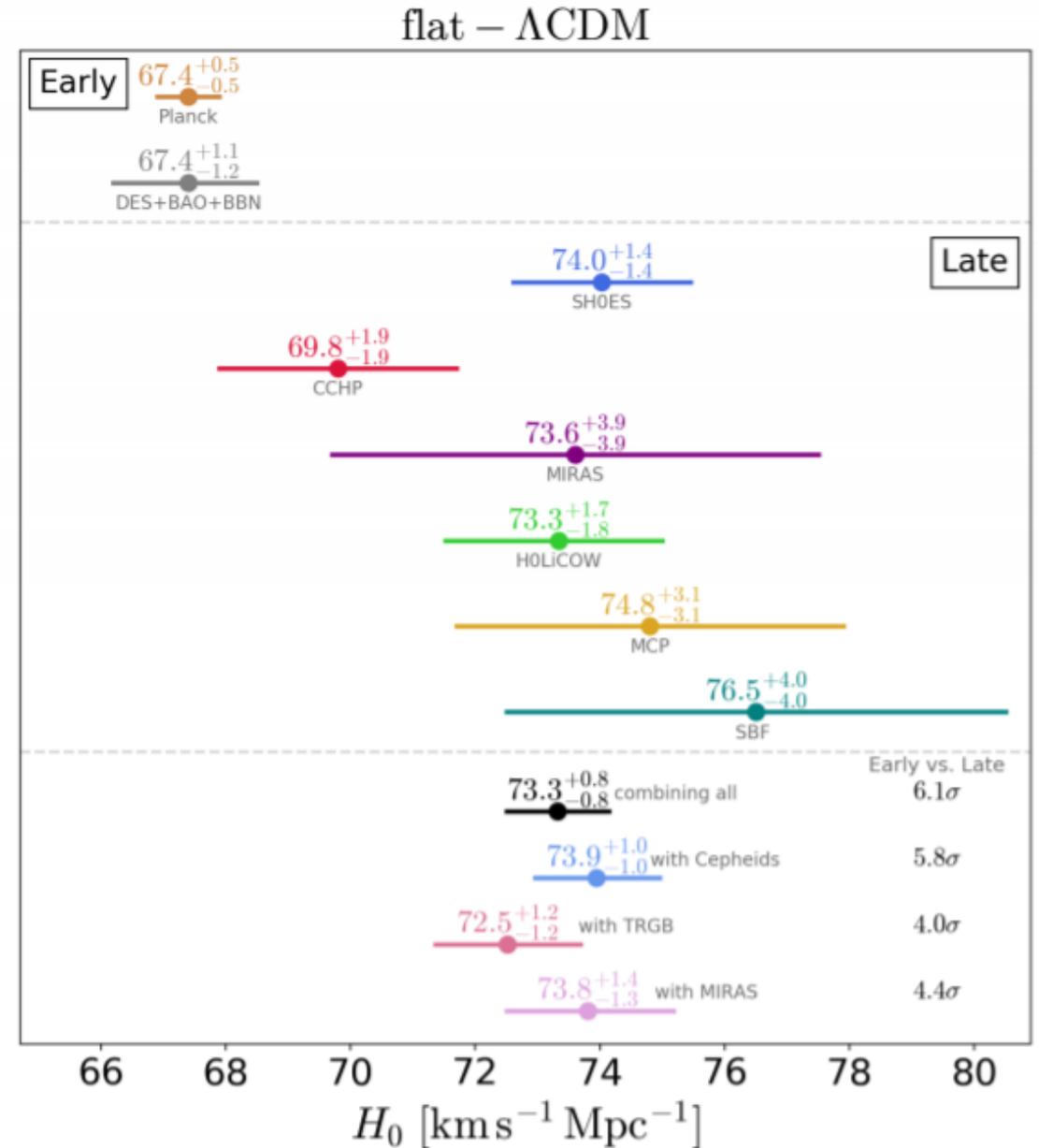


**UNIVERSITY OF  
CAMBRIDGE**

**MG16**  **5-10 JULY 2021**  
**SIXTEENTH MARCEL GROSSMANN MEETING**  
ON RECENT DEVELOPMENTS IN THEORETICAL AND EXPERIMENTAL GENERAL RELATIVITY, ASTROPHYSICS AND RELATIVISTIC FIEL THEORIES

# Tensions in Cosmology

- $4\sigma$  difference between the Hubble constant from the CMB vs. SH0ES.
- More than  $3\sigma$  difference between early and late measurements (CMB & LSS)

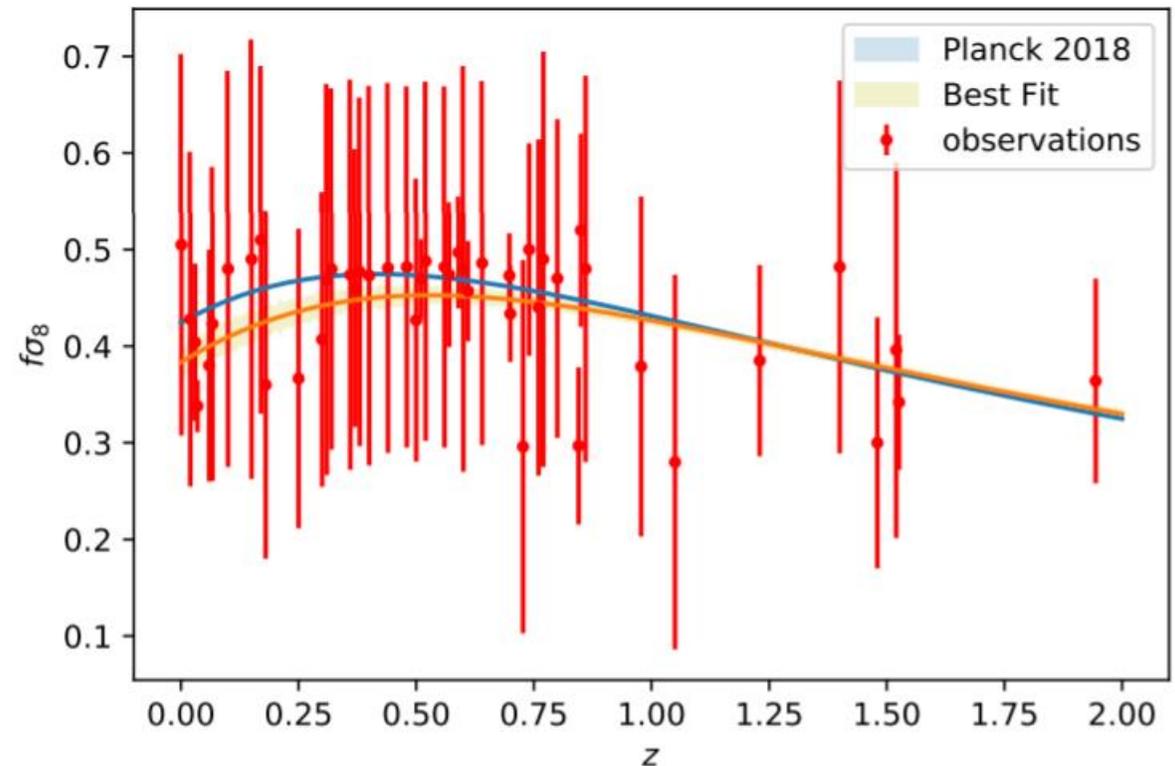


# The $S_8$ tension

- $\sigma_8$  - matter fluctuations averaged in spheres of radius  $8 h^{-1} Mpc$
- The  $S_8$  quantifying the amplitude of late-time matter fluctuations

$$S_8 = \sigma_8 \sqrt{\frac{\Omega_m}{0.3}}$$

- $\sim 50$  data points from the latest experiments:  $0.01 \leq z \leq 1.994$



# Homogenous Cosmology

- The Friedmann Eq. is normalized with:

$$H(z)^2 = H_0^2 (\Omega_m (1+z)^3 + (1 - \Omega_m)(1+z)^{-3(1+w)})$$

- The DE parametrization:  $w = w_0 + (1 - a)w_a$
- Radiation is ignored since the measurements are from the late universe.

# Perturbative Cosmology

- Matter perturbations:  $\delta_m = \delta\rho/\rho$
- The linear matter perturbations (in Fourier space):

$$\delta''_m(a) + \left( \frac{3}{a} + \frac{E'(a)}{E(a)} \right) \delta'_m(a) = \frac{3}{2} \frac{\Omega_m(a)}{a^2} \delta_m(a)$$

- assumption of small scales approximation (no k).

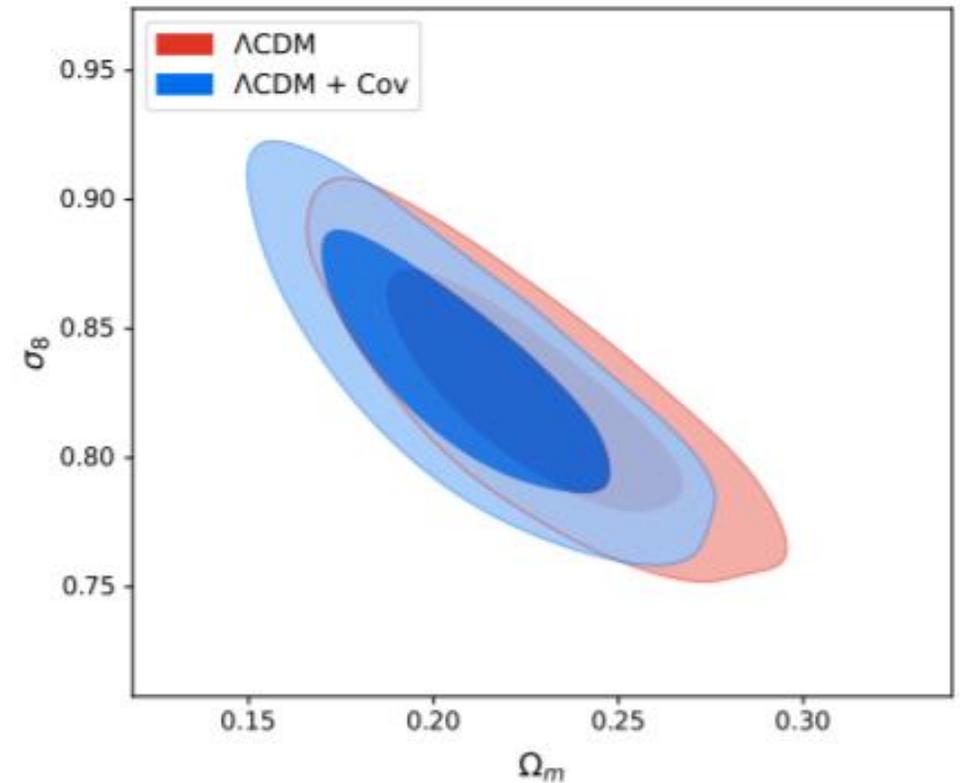
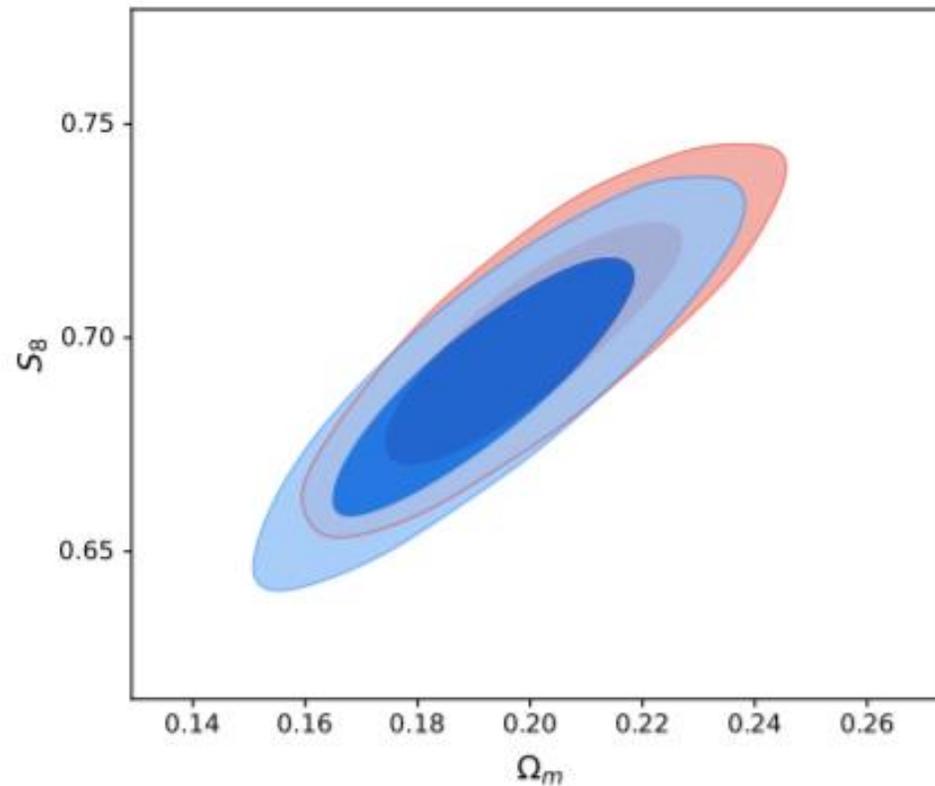
$$f(a) = \frac{d \log \delta_m}{d \log a}, \sigma_8(a) = \sigma_8 \frac{\delta_m(a)}{\delta_m(1)}$$
$$f \sigma_8(a) = a \frac{\delta'_m(a)}{\delta_m(1)} \sigma_8$$

# Uncorrelated and Correlated data

- Likelihood:  $\mathcal{L} = \exp\left(-\frac{\chi^2}{2}\right)$
- $\chi^2 = \sum_{i=1}^N \left( \frac{f\sigma_8(d_i) - f\sigma_8(\theta_i)}{\sigma_y^{(i)}} \right)^2$  for uncorrelated points.
- Can be written as:  $\chi^2 = V_i^T (d - \theta) M_{ij} V_j (d - \theta)$   
 $M_{ij} = (\dots \sigma_{ii} \dots)$  for uncorrelated measurements.
- We test correlations, using  $M_{ij} = 0.5\sigma_i\sigma_j$  with 25% random points.

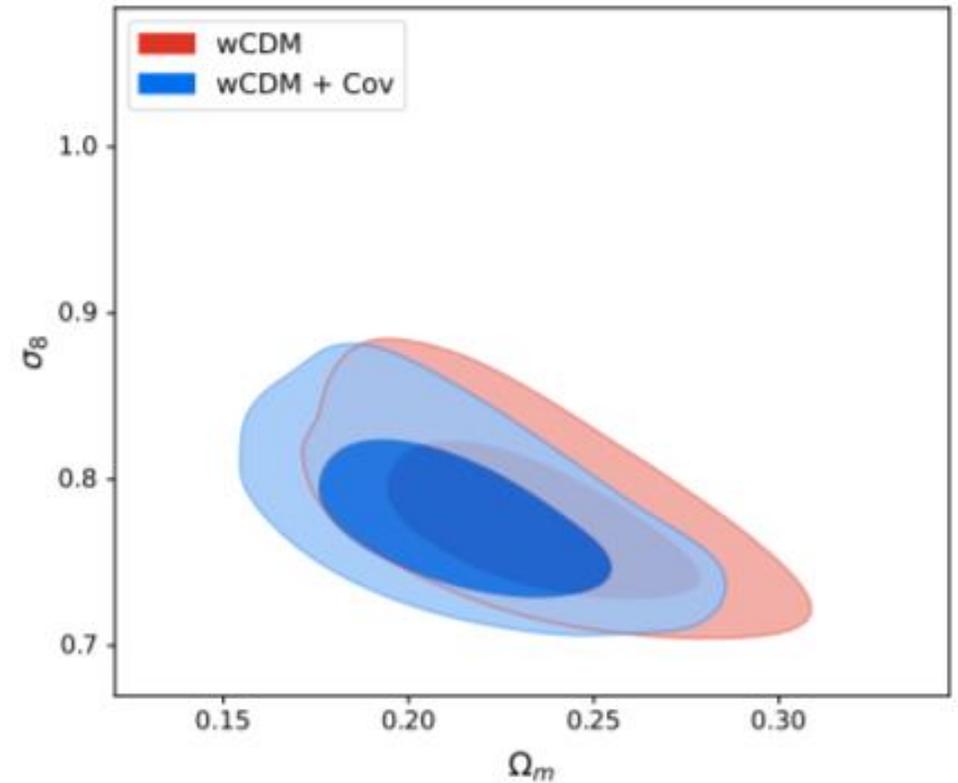
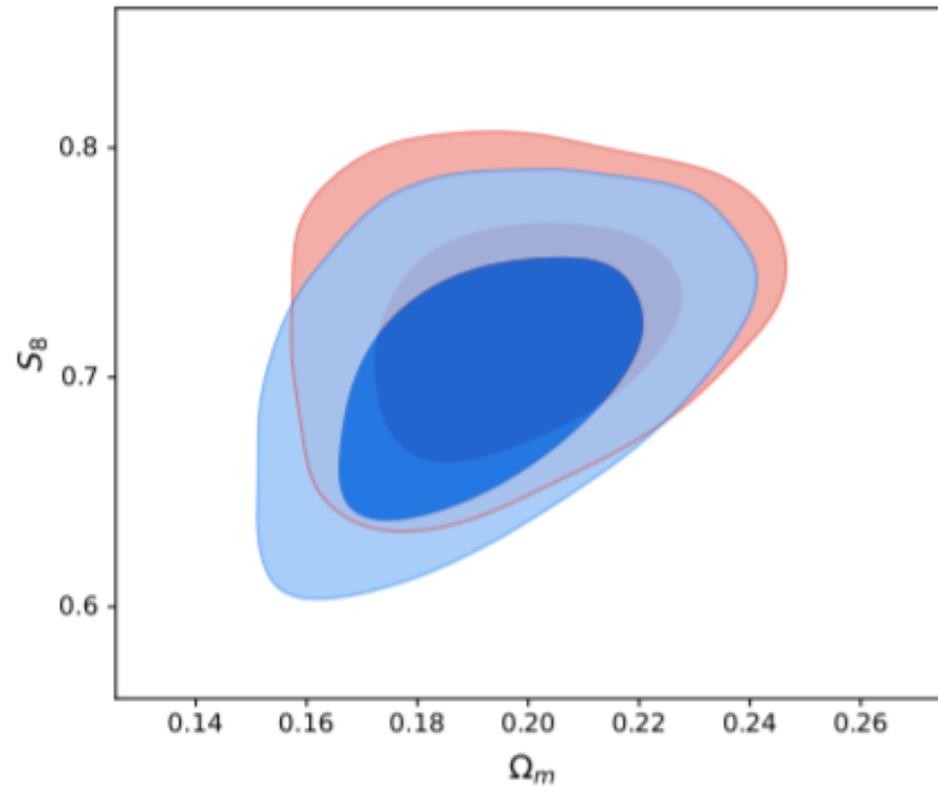
# Bayesian Analysis - $\Lambda$ CDM

Model	$\Omega_m$	$\sigma_8$	$S_8$	w
$\Lambda$ CDM	$0.201^{+0.036}_{-0.033}$	$0.857^{+0.044}_{-0.042}$	$0.700^{+0.038}_{-0.037}$	-1
$\Lambda$ CDM + Ran Cov	$0.192^{+0.037}_{-0.033}$	$0.863^{+0.046}_{-0.043}$	$0.690^{+0.040}_{-0.038}$	-1
Planck 2018	$0.311 \pm 0.019$	$0.811 \pm 0.006$	$0.834 \pm 0.016$	-1



# ***W***CDM

Model	$\Omega_m$	$\sigma_8$	$S_8$	w
$\Lambda$ CDM	$0.201^{+0.036}_{-0.033}$	$0.857^{+0.044}_{-0.042}$	$0.700^{+0.038}_{-0.037}$	-1
$\Lambda$ CDM + Ran Cov	$0.192^{+0.037}_{-0.033}$	$0.863^{+0.046}_{-0.043}$	$0.690^{+0.040}_{-0.038}$	-1
wCDM	$0.198^{+0.038}_{-0.035}$	$0.891^{+0.11}_{-0.095}$	$0.722^{+0.072}_{-0.067}$	$-0.90^{+0.29}_{-0.30}$
wCDM + Ran Cov	$0.192^{+0.038}_{-0.034}$	$0.879^{+0.10}_{-0.088}$	$0.701^{+0.077}_{-0.074}$	$-0.96^{+0.30}_{-0.31}$
Planck 2018	$0.311 \pm 0.019$	$0.811 \pm 0.006$	$0.834 \pm 0.016$	-1



# Gaussian Process Regression

- Unsupervised learning. Assumes Kernel distribution between two points.

- A Kernel dependent  $cov(f(x), f(x')) = K(x, x')$

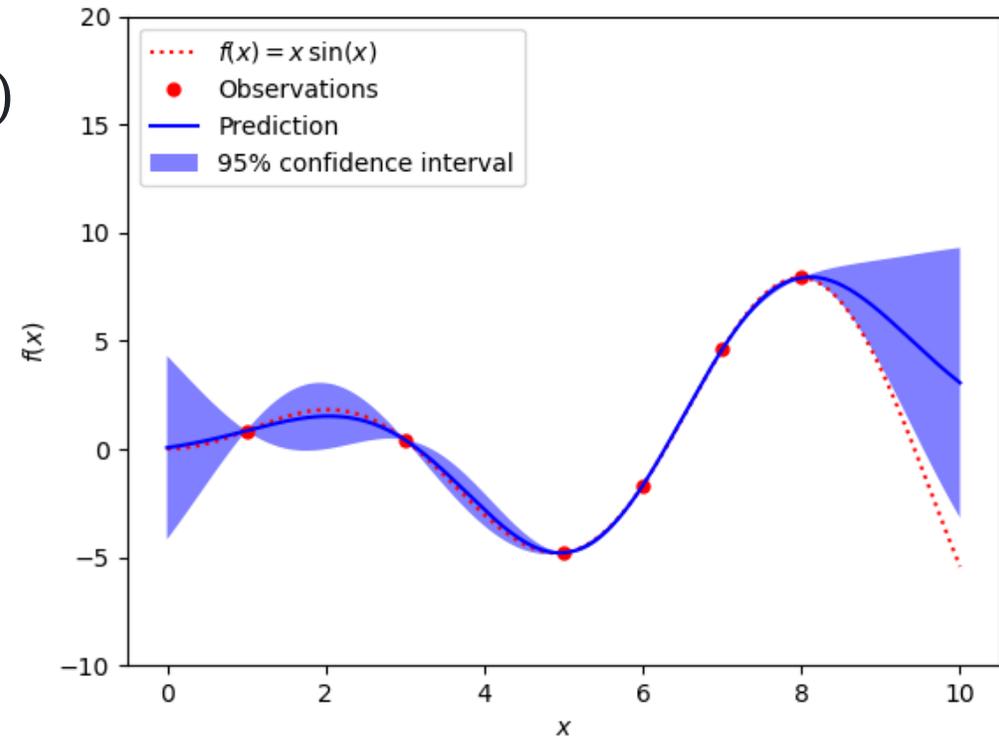
$$k(x, \tilde{x}) = \sigma_f^2 \exp\left(-\frac{(x - \tilde{x})^2}{2l^2}\right), \quad (13)$$

The Matern kernel with  $\nu = 7/2$ :

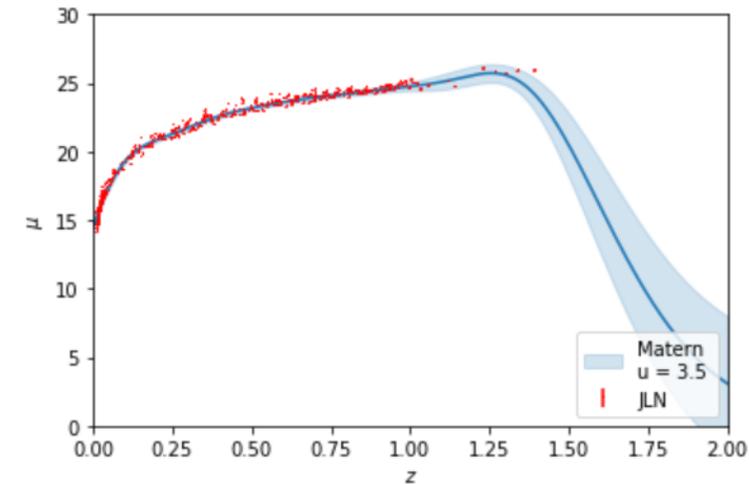
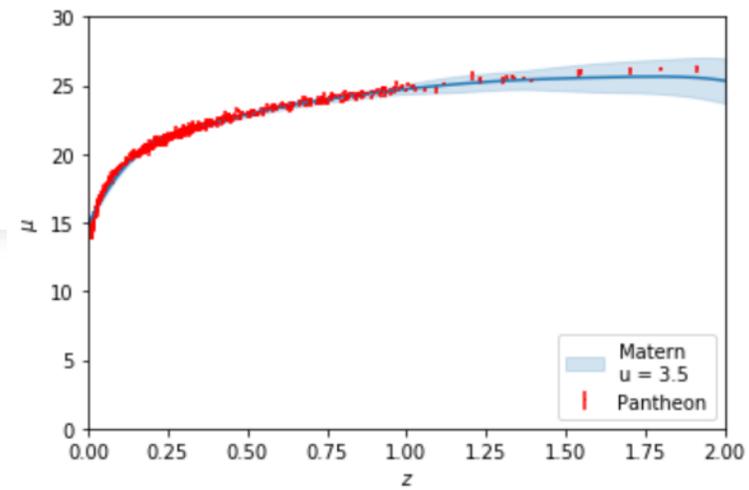
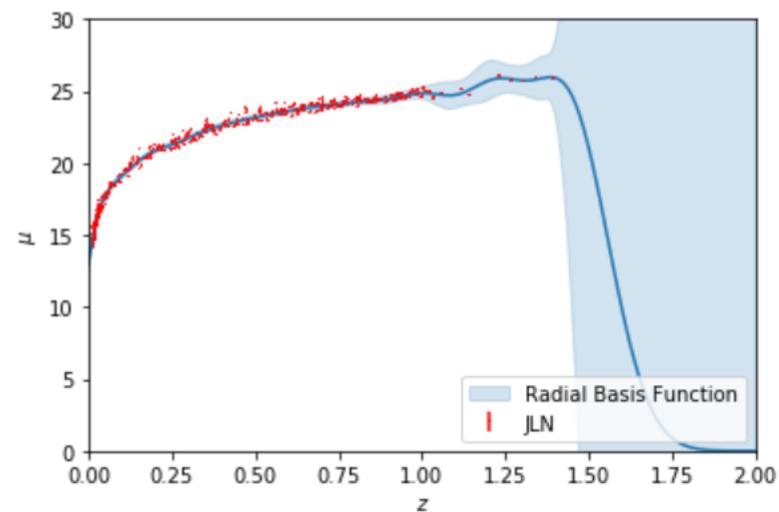
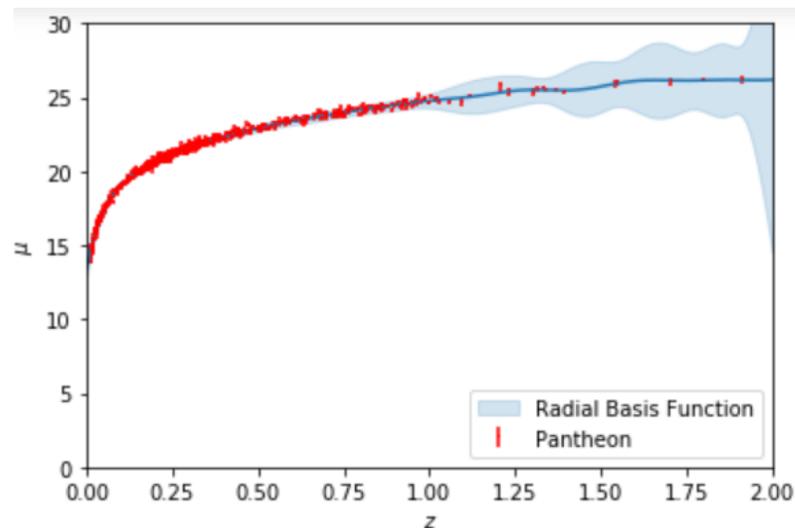
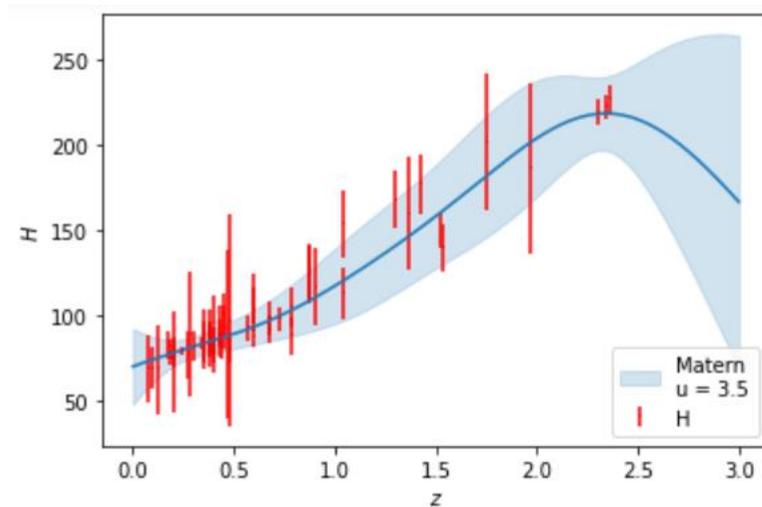
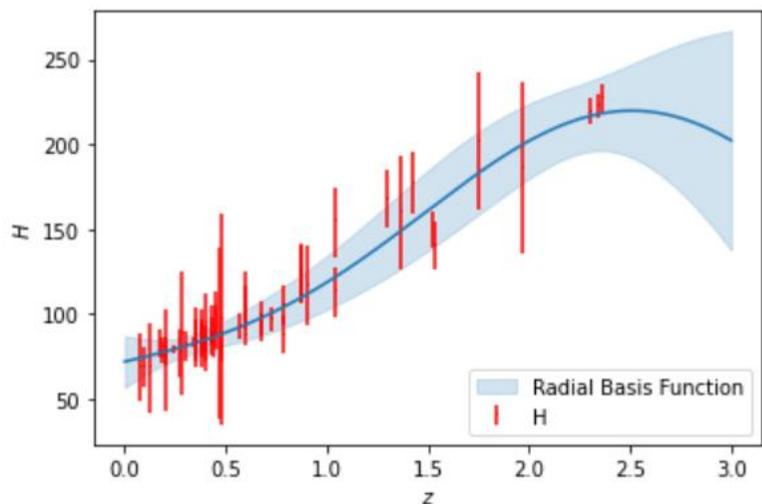
$$k(x, \tilde{x}) = \sigma_f^2 \exp\left(-\sqrt{7}\frac{|x - \tilde{x}|}{l}\right) \left(1 + \sqrt{7}\frac{|x - \tilde{x}|}{l} + 14\frac{(x - \tilde{x})^2}{5l^2} + 7\sqrt{7}\frac{|x - \tilde{x}|^3}{15l^3}\right), \quad (14)$$

and Matern kernel with  $\nu = 9/2$ :

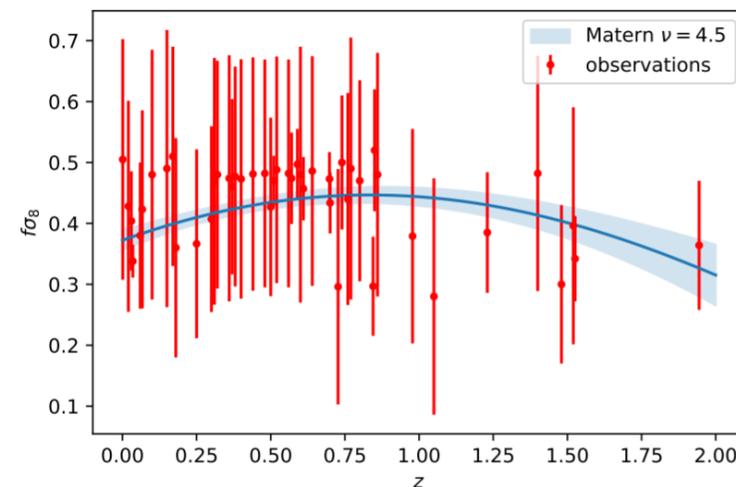
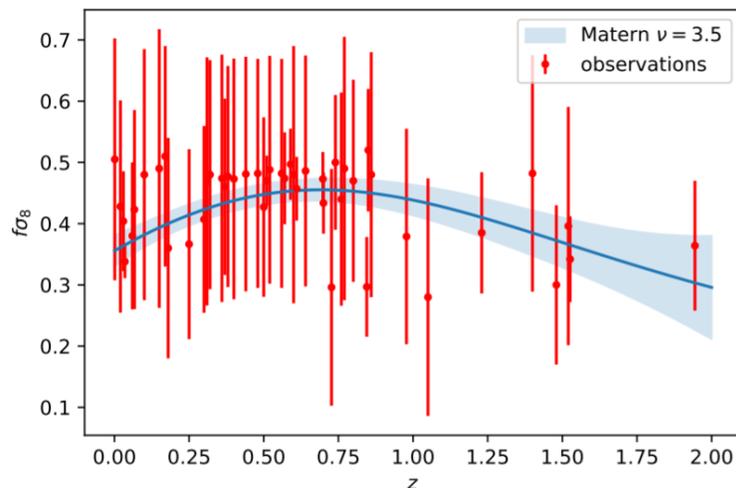
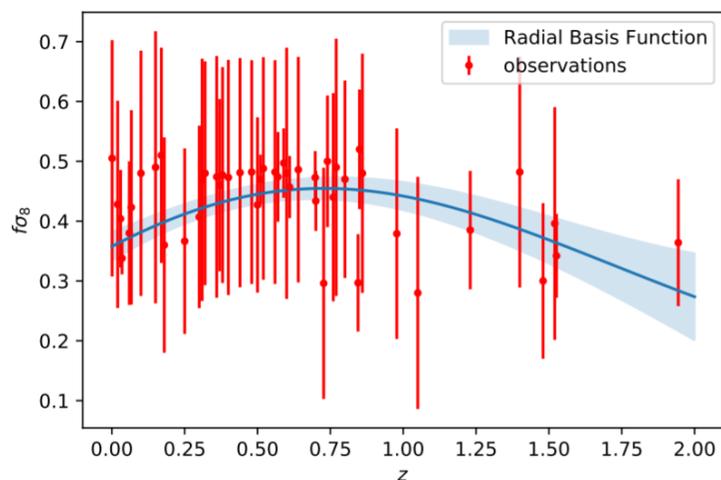
$$k(x, \tilde{x}) = \sigma_f^2 \exp\left(-3\frac{|x - \tilde{x}|}{l}\right) \left(1 + 3\frac{|x - \tilde{x}|}{l} + 27\frac{(x - \tilde{x})^2}{7l^2} + 18\frac{|x - \tilde{x}|^3}{7l^3} + 27\frac{(x - \tilde{x})^4}{35l^4}\right). \quad (15)$$



# Kernel Dependent



# GP for the $f\sigma_8$ data

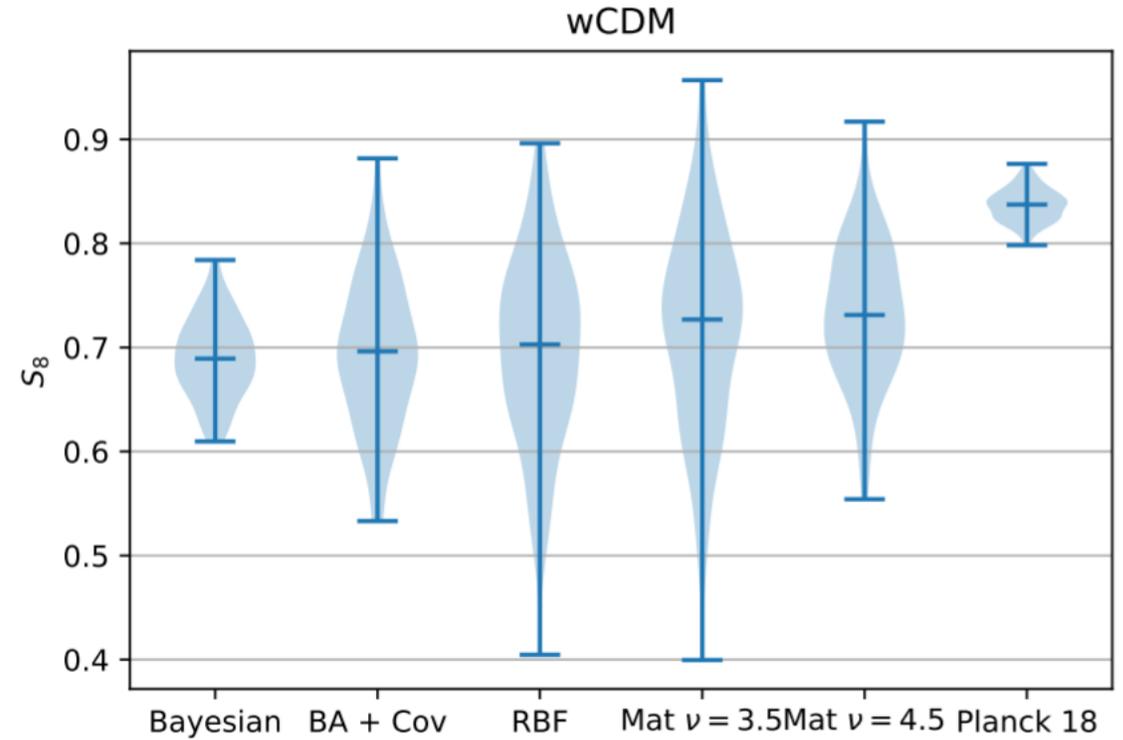
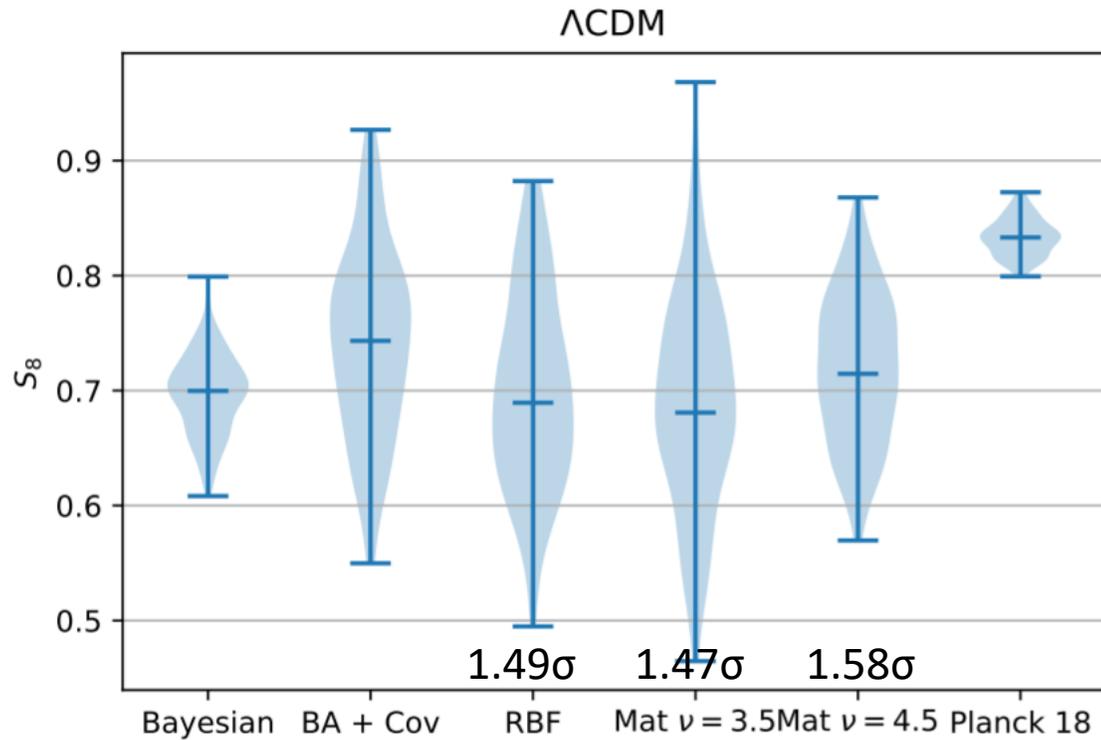


Kernel	RBF	Matern $\nu = 3.5$	Matern $\nu = 4.5$
$f\sigma_8(0)$	$0.358 \pm 0.0241$	$0.356 \pm 0.025$	$0.372 \pm 0.017$

$$f\sigma_8(0.001) = 0.505 \pm 0.0852 \quad \text{MTF} \quad \text{Mon. Not. Roy. Astron. Soc. 471, 3135 (2017)}$$

$$f\sigma_8(0.02) = 0.428 \pm 0.04652 \quad \text{6dFGS} \quad \text{JCAP 1705, 015 (2017)}$$

# Results for the $S_8$ values



$\Lambda$ CDM	RBF	Matern $\nu = 3.5$	Matern $\nu = 4.5$
$\sigma_8$	$0.77^{+0.21}_{-0.16}$	$0.77^{+0.21}_{-0.16}$	$0.79^{+0.19}_{-0.17}$
$\Omega_m$	$0.27^{+0.12}_{-0.11}$	$0.26^{+0.13}_{-0.11}$	$0.27^{+0.12}_{-0.10}$
$S_8$	$0.707^{+0.085}_{-0.085}$	$0.701^{+0.089}_{-0.089}$	$0.731^{+0.063}_{-0.062}$

$\Lambda$ CDM	RBF	Matern $\nu = 3.5$	Matern $\nu = 4.5$
$\sigma_8$	$0.77^{+0.21}_{-0.16}$	$0.77^{+0.21}_{-0.16}$	$0.79^{+0.19}_{-0.17}$
$\Omega_m$	$0.27^{+0.12}_{-0.11}$	$0.26^{+0.13}_{-0.11}$	$0.27^{+0.12}_{-0.10}$
$S_8$	$0.707^{+0.085}_{-0.085}$	$0.701^{+0.089}_{-0.089}$	$0.731^{+0.063}_{-0.062}$

# Final Results

- With B.A there is a clear tension with Planck measurement.
- With ML the tension is reduces to  $\sim 1 \sigma$  for some kernels.
- Still an open window to new physics...