



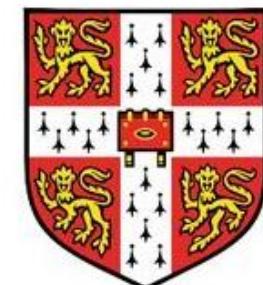
FIAS Frankfurt Institute
for Advanced Studies



Covariant Canonical Gauge Theory of Gravity

David Benisty

In a collaboration with D.Vasak and
J.Struckmeier



UNIVERSITY OF
CAMBRIDGE

Fundamental features of geometry

- The metric $g^{\alpha\beta}$ - the infinitesimal distance between two points:

$$ds^2 = g_{\alpha\beta} dx^\alpha dx^\beta$$

- The connection $\Gamma_{\alpha\beta}^\lambda$ - a curve with covariantly constant tangent vector:

$$\frac{d^2x^\mu}{ds^2} + \Gamma_{\alpha\beta}^\mu \frac{dx^\alpha}{ds} \frac{dx^\beta}{ds} = 0$$

- No Torsion :

$$\Gamma_{\alpha\beta}^\mu = \Gamma_{\beta\alpha}^\mu$$

$$\Gamma^\alpha_{\mu\nu} = \left\{ {}^\alpha_{\mu\nu} \right\} + K^\alpha_{\mu\nu} + L^\alpha_{\mu\nu}$$

Metric Formalism vs. Palatini Formalism

Metric formalism:

- Dependence only on the metric and its derivatives.

- Christoffel Symbol $\{\rho_{\mu\nu}\} = \frac{1}{2}g^{\rho\lambda}(g_{\lambda\mu,\nu} + g_{\lambda\nu,\mu} - g_{\mu\nu,\lambda})$

Palatini Formalism:

- The connection is being an independent degree of freedom in the action.

Is there any correspondence between those formulations?

Metric Formalism

- Field equation

$\frac{1}{2} R^2$:

$$R \left(R^{\mu\nu} - \frac{1}{4} g^{\mu\nu} R \right) - \nabla^\mu \nabla^\nu R + g^{\mu\nu} \nabla^2 R$$

$\frac{1}{2} R_{\mu\nu} R^{\mu\nu}$:

$$R^{\mu\alpha} R_\alpha^\nu - \frac{1}{4} g^{\mu\nu} (R_{\alpha\beta} R^{\alpha\beta}) - \frac{1}{2} \nabla_\gamma (\nabla^\mu R^{\nu\gamma} + \nabla^\nu R^{\mu\gamma}) + \frac{1}{2} \nabla^2 R^{\mu\nu} + \frac{1}{4} \nabla^2 R g^{\mu\nu}$$

$\frac{1}{2} R_{\mu\alpha\nu\beta} R^{\mu\alpha\nu\beta}$:

$$R^{\mu\alpha\beta\gamma} R^\nu_{\alpha\beta\gamma} - \frac{1}{4} g^{\mu\nu} (R_{\alpha\beta\gamma\delta} R^{\alpha\beta\gamma\delta}) + (\nabla_\alpha \nabla_\beta + \nabla_\beta \nabla_\alpha) R^{\mu\alpha\nu\beta}$$

Palatini formalism

- Field equation:

$\frac{1}{2} R^2$:

$$R \left(R^{\mu\nu} - \frac{1}{4} g^{\mu\nu} R \right)$$

$\frac{1}{2} R_{\mu\nu} R^{\mu\nu}$:

$$R^{\mu\alpha} R^\nu_\alpha - \frac{1}{4} g^{\mu\nu} (R_{\alpha\beta} R^{\alpha\beta})$$

$\frac{1}{2} R_{\mu\alpha\nu\beta} R^{\mu\alpha\nu\beta}$:

$$R^{\mu\alpha\beta\gamma} R^\nu_{\alpha\beta\gamma} - \frac{1}{4} g^{\mu\nu} (R_{\alpha\beta\gamma\delta} R^{\alpha\beta\gamma\delta})$$

- Γ variation:

$$K_{\lambda(\alpha)}^{\mu\nu} = (g^{\mu\nu} \nabla_\lambda - \frac{1}{2} \delta_\lambda^\nu \nabla^\mu - \frac{1}{2} \delta_\lambda^\mu \nabla^\nu) R$$

$$K_{\lambda(\beta)}^{\mu\nu} = \nabla_\lambda R^{\mu\nu} - \frac{1}{4} \delta_\lambda^\mu \nabla^\nu R - \frac{1}{4} \delta_\lambda^\nu \nabla^\mu R$$

$$K_{\lambda(\beta)}^{\mu\nu} = \nabla_\sigma R_\lambda^{\mu\sigma\nu} + \nabla_\sigma R_\lambda^{\nu\sigma\mu}$$

A difference in squared gravity terms

- Palatini Formalism:

$$\frac{1}{2} R^2:$$

$$R \left(R^{\mu\nu} - \frac{1}{4} g^{\mu\nu} R \right)$$

$$\frac{1}{2} R_{\mu\nu} R^{\mu\nu}:$$

$$R^{\mu\alpha} R^\nu_\alpha - \frac{1}{4} g^{\mu\nu} (R_{\alpha\beta} R^{\alpha\beta})$$

$$\frac{1}{2} R_{\mu\alpha\nu\beta} R^{\mu\alpha\nu\beta}:$$

$$R^{\mu\alpha\beta\gamma} R^\nu_{\alpha\beta\gamma} - \frac{1}{4} g^{\mu\nu} (R_{\alpha\beta\gamma\delta} R^{\alpha\beta\gamma\delta})$$

- Contribution for the metric formalism:

$$-\nabla^\mu \nabla^\nu R + g^{\mu\nu} \nabla^2 R$$

$$-\frac{1}{2} \nabla_\gamma (\nabla^\mu R^{\nu\gamma} + \nabla^\nu R^{\mu\gamma}) + \frac{1}{2} \nabla^2 R^{\mu\nu} + \frac{1}{4} \nabla^2 R$$

$$+ (\nabla_\alpha \nabla_\beta + \nabla_\beta \nabla_\alpha) R^{\mu\alpha\nu\beta}$$

The metricity constraint

Phys.Rev. D98 (2018) no.4, 044023
[arXiv:1805.09667v2](https://arxiv.org/abs/1805.09667v2)

David Benisty, Eduardo Guendelman

The theorem:

$$\mathcal{L}(g) \text{ 2order} \Leftrightarrow \mathcal{L}(g, \Gamma) + k^{\alpha\beta\gamma} g_{\alpha\beta;\gamma} \text{ 1order}$$

- $\delta k^{\alpha\beta\gamma}: \quad g_{\alpha\beta;\gamma} = 0 \quad \Rightarrow \quad \Gamma_{\mu\nu}^\rho = \{\rho_{\mu\nu}\}$

- $\delta\Gamma_\beta^\alpha: \quad \frac{\delta\mathcal{L}(\kappa)}{\delta\Gamma_{\mu\nu}^\rho} = -k^{\alpha\mu\nu} g_{\rho\alpha} - k^{\alpha\nu\mu} g_{\rho\alpha}$

- $\delta g^{\mu\nu}: \quad G_{(\kappa)}^{\mu\nu} = \frac{\delta\mathcal{L}(\kappa)}{\delta g_{\mu\nu}} = -k_{;\lambda}^{\mu\nu\lambda}$

The contribution for the metric E.o.M

$$g^{\rho\sigma} \frac{\partial \mathcal{L}(\kappa)}{\partial \Gamma_{\mu\nu}^\rho} = -k^{\sigma\mu\nu} - k^{\sigma\nu\mu}$$

$$g^{\rho\nu} \frac{\partial \mathcal{L}(\kappa)}{\partial \Gamma_{\mu\sigma}^\rho} = -k^{\nu\mu\sigma} - k^{\nu\sigma\mu}$$

By changing the indices:

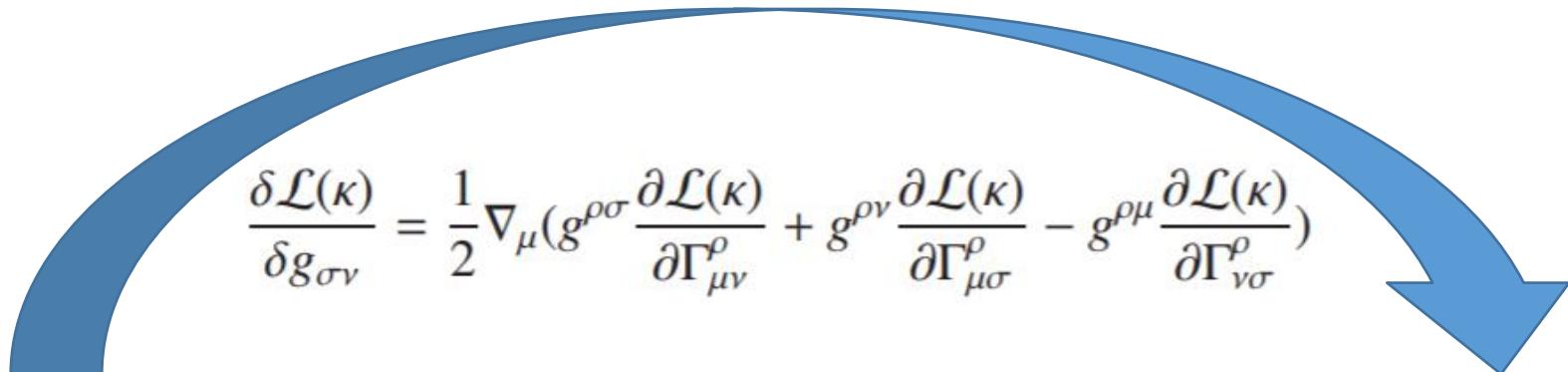
$$g^{\rho\mu} \frac{\partial \mathcal{L}(\kappa)}{\partial \Gamma_{\nu\sigma}^\rho} = -k^{\mu\nu\sigma} - k^{\mu\sigma\nu}$$

- The new contribution:

$$\frac{\delta \mathcal{L}(\kappa)}{\delta g_{\sigma\nu}} = \frac{1}{2} \nabla_\mu \left(g^{\rho\sigma} \frac{\partial \mathcal{L}(\kappa)}{\partial \Gamma_{\mu\nu}^\rho} + g^{\rho\nu} \frac{\partial \mathcal{L}(\kappa)}{\partial \Gamma_{\mu\sigma}^\rho} - g^{\rho\mu} \frac{\partial \mathcal{L}(\kappa)}{\partial \Gamma_{\nu\sigma}^\rho} \right)$$

A correspondence for squared gravity terms

$$G_{(\gamma)}^{\mu\nu} \text{ 2order} = G_{(\gamma)}^{\mu\nu} \text{ 1order} + G_{(\kappa)}^{\mu\nu}$$



$$\frac{\delta \mathcal{L}(\kappa)}{\delta g_{\sigma\nu}} = \frac{1}{2} \nabla_\mu (g^{\rho\sigma} \frac{\partial \mathcal{L}(\kappa)}{\partial \Gamma_{\mu\nu}^\rho} + g^{\rho\nu} \frac{\partial \mathcal{L}(\kappa)}{\partial \Gamma_{\mu\sigma}^\rho} - g^{\rho\mu} \frac{\partial \mathcal{L}(\kappa)}{\partial \Gamma_{\nu\sigma}^\rho})$$

$$K_{\lambda(\alpha)}^{\mu\nu} = (g^{\mu\nu} \nabla_\lambda - \frac{1}{2} \delta_\lambda^\nu \nabla^\mu - \frac{1}{2} \delta_\lambda^\mu \nabla^\nu) R$$

$$K_{\lambda(\beta)}^{\mu\nu} = \nabla_\lambda R^{\mu\nu} - \frac{1}{4} \delta_\lambda^\mu \nabla^\nu R - \frac{1}{4} \delta_\lambda^\nu \nabla^\mu R$$

$$K_{\lambda(\beta)}^{\mu\nu} = \nabla_\sigma R_\lambda^{\mu\sigma\nu} + \nabla_\sigma R_\lambda^{\nu\sigma\mu}$$

$$G_{(\alpha)}^{\mu\nu} \text{ 2order} = G_{(\alpha)}^{\mu\nu} \text{ 1order} + \alpha [-\nabla^\mu \nabla^\nu R + g^{\mu\nu} \square R]$$

$$G_{(\beta)}^{\mu\nu} \text{ 2order} = G_{(\beta)}^{\mu\nu} \text{ 1order} + \beta \left[-\frac{1}{2} \nabla_\gamma (\nabla^\mu R^{\gamma\gamma} + \nabla^\nu R^{\mu\gamma}) + \frac{1}{2} \square R^{\mu\nu} + \frac{1}{4} \square R \right]$$

$$G_{(\gamma)}^{\mu\nu} \text{ 2order} = G_{(\gamma)}^{\mu\nu} \text{ 1order} + \gamma [(\nabla_\alpha \nabla_\beta + \nabla_\beta \nabla_\alpha) R^{\mu\alpha\nu\beta}]$$

A path integral approach

$$\begin{aligned} \mathcal{Z} &= \int \mathcal{D}k \mathcal{D}g \mathcal{D}\Gamma e^{i \int d^4x \sqrt{-g} (\mathcal{L}(g, \Gamma) + k^{\alpha\beta\gamma} g_{\alpha\beta;\gamma})} \\ &= \int \delta(g_{\alpha\beta;\gamma}) \mathcal{D}g \mathcal{D}\Gamma e^{i \int d^4x \sqrt{-g} \mathcal{L}(g, \Gamma)} \\ &\sim \int \delta(\Gamma_{\mu\nu}^\rho - \{\rho_{\mu\nu}\}) \mathcal{D}g \mathcal{D}\Gamma e^{i \int d^4x \sqrt{-g} \mathcal{L}(g, \Gamma)} \\ &= \int \mathcal{D}g e^{i \int d^4x \sqrt{-g} \mathcal{L}(g)} \end{aligned}$$

First order Formalism

Second order Formalism

J. Struckmeier, J. Muench, D. Vasak, J. Kirsch, M. Hanuske, and H. Stoecker

Hamiltonian for dynamic spacetime

$k^{\alpha\beta\gamma}$ as a metric conjugate momenta:

$$\tilde{L} = \tilde{k}^{\alpha\beta\gamma} g_{\alpha\beta,\gamma} - \tilde{H}$$

For a gauge theory that includes general mapping of space time $x \rightarrow X$, we identify the connection $\gamma_{\alpha\beta}^\eta$ as the gauge field of these C.T.

Condition for canonical transformation under a dynamical spacetime.

$$\begin{aligned} S &= \int (\tilde{k}^{\alpha\beta\gamma} g_{\alpha\beta,\gamma} - \tilde{q}_\alpha^{\beta\gamma\delta} \gamma_{\beta\gamma,\delta}^\alpha - \tilde{H} - \tilde{F}_{1,\beta}^\beta) d^4x \\ &= \int (\tilde{K}^{\alpha\beta\gamma} G_{\alpha\beta,\gamma} - \frac{1}{2} \tilde{Q}_\alpha^{\beta\gamma\delta} \Gamma_{\beta\gamma,\delta}^\alpha - \tilde{H}') d^4X \end{aligned}$$

\tilde{F}_1^β is the generating function.

Gauge Hamiltonian

$$F_3^\mu = -k^{\alpha\beta\mu} G_{\xi\lambda} \frac{\partial X^\xi}{\partial x^\alpha} \frac{\partial X^\lambda}{\partial x^\beta} - q_\eta^{\alpha\beta\mu} \left(\Gamma_{\xi\lambda}^\tau \frac{\partial x^\eta}{\partial X^\tau} \frac{\partial X^\xi}{\partial x^\alpha} \frac{\partial X^\lambda}{\partial x^\beta} + \frac{\partial x^\eta}{\partial X^\tau} \frac{\partial X^\tau}{\partial x^\alpha \partial x^\beta} \right)$$

Yields the transformation rules for the metric:

$$G_{\mu\nu} = g_{\alpha\beta} \frac{\partial x^\alpha}{\partial X^\nu} \frac{\partial x^\beta}{\partial X^\mu},$$

and the connection:

$$\Gamma_{\alpha\beta}^\kappa = \gamma_{\eta\tau}^\xi \frac{\partial x^\eta}{\partial X^\tau} \frac{\partial X^\xi}{\partial x^\alpha} \frac{\partial X^\lambda}{\partial x^\beta} + \frac{\partial x^\xi}{\partial X^\alpha \partial X^\beta} \frac{\partial X^\kappa}{\partial x^\xi}$$

rotation In-homogeneous
 part

$$L = k^{\alpha\beta\gamma} g_{\alpha\beta;\gamma} - \frac{1}{2} q_\alpha^{\beta\gamma\delta} R_{\beta\gamma\delta}^\alpha - H_{Dyn}(g, q, k) + L_m$$

The equations of motion

$$L = k^{\alpha\beta\gamma} g_{\alpha\beta;\gamma} - \frac{1}{2} q_\alpha^{\beta\gamma\delta} R^\alpha_{\beta\gamma\delta} - H_{Dyn}(g, q, k) + L_m$$

$$k^{\alpha\beta\gamma} : g_{\alpha\beta;\gamma} = \frac{\partial H_{Dyn}}{\partial k^{\alpha\beta\gamma}}$$

$$\gamma^\mu_{\nu\rho} : (k^{\alpha\mu\nu} + k^{\alpha\nu\mu}) g_{\alpha\rho} = -\frac{1}{2} \nabla_\beta (q_\rho^{\mu\beta\nu} + q_\rho^{\nu\beta\mu})$$

$$q^\sigma_{\mu\nu\rho} : -\frac{1}{2} R^\sigma_{\mu\nu\rho} = \frac{\partial H_{Dyn}}{\partial q^\sigma_{\mu\nu\rho}}$$

$$g^{\mu\nu} : T^{\mu\nu} = g^{\mu\nu} \left(-k^{\alpha\beta\gamma} g_{\alpha\beta;\gamma} + \frac{1}{2} q_\alpha^{\beta\gamma\delta} R^\alpha_{\beta\gamma\delta} \right) + 2k^{\mu\nu\lambda}_{;\lambda} - \frac{2}{\sqrt{-g}} \frac{\partial \tilde{H}_{Dyn}}{\partial g^{\mu\nu}}$$

$$2k^{\mu\nu\lambda}_{;\lambda} = -\nabla_\gamma \nabla_\alpha (q^{\mu\gamma\nu\alpha} + q^{\nu\gamma\mu\alpha})$$

Sample H_{Dyn}

$$L = k^{\alpha\beta\gamma} g_{\alpha\beta;\gamma} - \frac{1}{2} q_\alpha^{\beta\gamma\delta} R_{\beta\gamma\delta}^\alpha - H_{Dyn}(g, q, k)$$

For a Hamiltonian with a quadratic momentum q :

$$\tilde{\mathcal{H}}_{Dyn} = \frac{1}{4g_1} \tilde{q}_\eta^{\alpha\xi\beta} \tilde{q}_\alpha^{\eta\tau\lambda} g_{\xi\tau} g_{\beta\lambda} \frac{1}{\sqrt{-g}} - g_2 \tilde{q}_\eta^{\alpha\eta\beta} g_{\alpha\beta} + g_3 \sqrt{-g}.$$

The equation of motions: (k, Γ, q, g)

k :
$$g_{\alpha\beta;\gamma} = -\frac{\partial \tilde{\mathcal{H}}_{Dyn}}{\partial \tilde{k}^{\alpha\beta\gamma}} = 0$$

q :
$$q_{\eta\alpha\xi\beta} = g_1 (R_{\eta\alpha\xi\beta} - R_{\eta\alpha\xi\beta}|_{max}),$$

$$R_{\eta\alpha\xi\beta}|_{max} = g_2 (g_{\eta\xi} g_{\alpha\beta} - g_{\eta\beta} g_{\alpha\xi})$$

$\Gamma:$

$$\frac{1}{2} \nabla_\mu (\tilde{q}_\beta^{\alpha\eta\mu} + \tilde{q}_\beta^{\eta\alpha\mu}) = -(\tilde{k}^{\lambda\alpha\eta} + \tilde{k}^{\lambda\eta\alpha}) g_{\lambda\beta}$$

The energy momentum tensor:

$$T^{\mu\nu} = \frac{1}{\kappa^2} G^{\mu\nu} + g_1 Q^{\mu\nu} + g_1 S^{\mu\nu} + g^{\mu\nu} \Lambda$$

where:

$$G^{\mu\nu} = R^{\mu\nu} - \frac{1}{2} g^{\mu\nu} R \quad L = g_1 R^{\alpha\beta\gamma\delta} R_{\alpha\beta\gamma\delta} + \frac{1}{2} R + \Lambda$$

$$Q^{\mu\nu} = R^{\mu\alpha\beta\gamma} R^\nu_{\alpha\beta\gamma} - \frac{1}{4} g^{\mu\nu} R^{\alpha\beta\gamma\delta} R_{\alpha\beta\gamma\delta}$$

$$S^{\mu\nu} = (\nabla_\alpha \nabla_\beta + \nabla_\beta \nabla_\alpha) R^{\mu\alpha\nu\beta}$$

$$g_1 = \frac{3}{16\pi G\Lambda}, \quad g_2 = \frac{1}{3}\Lambda.$$

$H_{Dyn}(g, q)$ without breaking metricity

Similar definitions for q tensors as Ricci tensor and Ricci scalar:

$$q_{\alpha\lambda\beta}^{\lambda} = q_{\alpha\beta}, \quad q = q_{\alpha\beta} g^{\alpha\beta}$$

Lead to the general combination of the covariant Hamiltonian:

$$H_{Dyn}(g, q) = g_0 - \frac{g_1}{2} q - \frac{g_{21}}{4} q^2 - \frac{g_{22}}{4} q^{\mu\nu} q_{\mu\nu} - \frac{g_{23}}{4} q_{\alpha\beta\gamma}^{\lambda} q_{\lambda}^{\alpha\beta\gamma}$$

$$L = k^{\alpha\beta\gamma} g_{\alpha\beta;\gamma} - \frac{1}{2} q_{\alpha}^{\beta\gamma\delta} R_{\beta\gamma\delta}^{\alpha} - H_{Dyn}(g, q) + L_m$$

The E.o.M behave as in the second order formalism, and the energy momentum tensor is conserved.

The equation of motions

$$H_{Dyn}(g, q) = g_0 - \frac{g_1}{2}q - \frac{g_{21}}{4}q^2 - \frac{g_{22}}{4}q^{\mu\nu}q_{\mu\nu} - \frac{g_{23}}{4}q_{\alpha\beta\gamma}^{\lambda}q_{\lambda}^{\alpha\beta\gamma}$$

The E.o.M:

$$\mathbf{k}^{\alpha\beta\gamma}: g_{\alpha\beta;\gamma} = \frac{\partial H_{Dyn}}{\partial k^{\alpha\beta\gamma}} = 0 \quad \Rightarrow \quad \gamma_{\nu\rho}^{\mu} = \begin{Bmatrix} \mu \\ \nu\rho \end{Bmatrix}$$

$$\mathbf{q}_{\mu\nu\rho}^{\sigma}: R_{\mu\nu\rho}^{\sigma} = \frac{\partial H_{Dyn}}{\partial q_{\mu\nu\rho}^{\sigma}} = g_{23}q_{\mu\nu\rho}^{\sigma} + \delta_{\nu}^{\rho} \left(g_{22}q_{\mu\rho} + g_{\mu\rho}(g_{21}q + g_1) \right)$$

Trace δ_{ρ}^{ν} :

$$R_{\mu\rho} = (g_{23} + 4g_{22})q_{\mu\rho} + 4g_{\mu\rho}(g_{21}q + g_1)$$

Another trace $g^{\mu\rho}$:

$$R = (g_{23} + 4g_{22} + 16g_{21})q + 16g_1$$

The momentum of the connection

$$q_{\mu\nu\rho}^\sigma = \alpha_1 R_{\mu\nu\rho}^\sigma + \delta_\nu^\rho \left(\alpha_2 R_{\mu\rho} + g_{\mu\rho} (\alpha_3 q + \alpha_4) \right)$$

$$\alpha_1 = \frac{1}{g_{23}}$$

$$\alpha_2 = -\frac{1}{g_{23}} \frac{g_{22}}{g_{23} + 4g_{22}}$$

$$\alpha_3 = -\frac{1}{16g_{21} + 4g_{22} + g_{23}} \frac{g_{21}}{g_{23} + 4g_{22}}$$

$$\alpha_4 = -\frac{g_1}{16g_{21} + 4g_{22} + g_{23}}$$

Legendre Transform

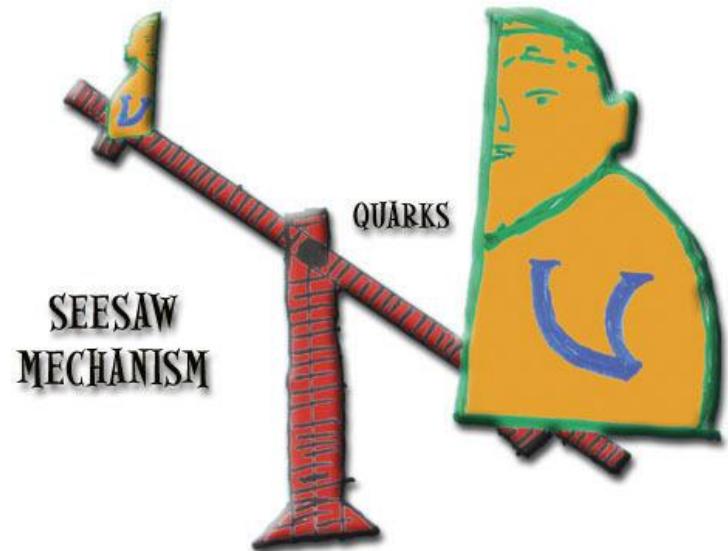
$$-\mathcal{L} = \frac{\alpha_1}{4} R_{\alpha\beta\gamma}^{\lambda} R_{\lambda}^{\alpha\beta\gamma} + \frac{\alpha_2}{4} R_{\mu\nu} R^{\mu\nu} + \frac{\alpha_3}{4} R^2 + \frac{\alpha_4}{4} R + \Lambda$$

The cosmological constant:

$$\Lambda = 4g_1\alpha_4 + g_0 = -\frac{g_1^2}{16g_{21} + 4g_{22} + g_{23}} + g_0$$

Cosmological seesaw mechanism

$$16g_{21} + 4g_{22} + g_{23} \gg 1$$



Gauss Bonnet

$$-\mathcal{L} = R_{\alpha\beta\gamma}^{\lambda} R_{\lambda}^{\alpha\beta\gamma} - 4R_{\mu\nu}R^{\mu\nu} + R^2 - \frac{\alpha_4}{4}R - \Lambda$$

$$g_{21} = 1, \quad g_{22} = -4, \quad g_{23} = 15$$

Conformal gravity

$$\mathcal{L} = -\frac{1}{4}C_{\mu\nu\alpha\beta}C^{\mu\nu\alpha\beta}$$

$$g_{21} = 1, \quad g_{22} = 10, \quad g_{23} = -35$$

$$g_1 = g_0 = 0$$

Coincident General Relativity

$$\mathcal{Q} = g^{\mu\nu} \left(\left\{ {}^\alpha_{\beta\mu} \right\} \left\{ {}^\beta_{\nu\alpha} \right\} - \left\{ {}^\alpha_{\beta\alpha} \right\} \left\{ {}^\beta_{\mu\nu} \right\} \right).$$

$$\mathcal{S} = \int d^n x \left[\frac{1}{2} \sqrt{-g} \mathbb{T} + \lambda_\alpha{}^{\beta\mu\nu} R^\alpha{}_{\beta\mu\nu} + \lambda^\alpha{}_{\mu\nu} \nabla_\alpha g^{\mu\nu} \right].$$

$$\mathcal{S}_G = \int d^n x \left[\frac{1}{2} \sqrt{-g} f(\mathcal{Q}) + \lambda_\alpha{}^{\beta\mu\nu} R^\alpha{}_{\beta\mu\nu} + \lambda_\alpha{}^{\mu\nu} T^\alpha{}_{\mu\nu} \right].$$

Summary



- Modified theories of gravity –
A connection between 2 formulations.
- Hamiltonian gauge theory of gravity –
Momenta for the metric and the connection.
- Quantum Gravity –
No higher derivatives in the action.

