

Effective field theory from Relativistic Generalized Uncertainty Principle

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Outline

- 1 Minimum length and Quantum Gravity
- 2 Relativistic Generalized Uncertainty Principle¹
- 3 Quantum Field Theory with Minimum Length^{2 3}
- 4 Summary and outlook

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1. V. Todorinov, P. Bosso, S. Das, *Ann. Phys.* **405**, 92-100 (2019) [arXiv:1810.11761].
 2. P. Bosso, S. Das, V. Todorinov [arXiv:2005.03771 [gr-qc]]
 3. P. Bosso, S. Das, V. Todorinov [arXiv:2005.03772 [gr-qc]].

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Generalized Uncertainty Principle

- Uncertainty relation for a Heisenberg microscope with Newtonian potential

$$\Delta x \Delta p \geq \left(h + \underbrace{\frac{Gh}{c^3}}_{=l_{Pl}^2} \frac{\Delta p^2}{h} \right)$$

- *"Hilbert space representation of the minimal length uncertainty relation" published in 1995⁴*

$$[x, p] = i\hbar(1 + \beta p^2)$$

- This amounts to the following minimum uncertainty in position

$$\Delta x_{\min} = \hbar\sqrt{\beta}$$

4. A. Kempf, G. Mangano, and R. B. Mann, Phys. Rev. **D52**, 1108 (1995) [arXiv:hep-th/9412167]

Problems

- Frame-Dependence of Minimum length
 - ▶ Length is **not** a Lorentz invariant
 - ▶ Generalized Uncertainty Principle adopt this form

$$[x_i, p_j] = i\hbar\delta_{ij}(1 + \beta_1\vec{p}^2) + i\hbar\beta_2 p_i p_j$$

where $i, j \in \{1, 2, 3\}$.

- Composition Law Problem
 - ▶ Modified dispersion relation

$$E^2 = p^2 c^2 + m^2 c^4 \rightarrow E^2 = p^2 c^2 + m^2 c^4 + \mathcal{O}(p^4)$$

- ▶ It can be shown that for a composite system, moving at the same speed

$$\vec{P}_3 \neq \vec{P}_2 + \vec{P}_1 \quad E_3 \neq E_1 + E_2$$

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Relativistic Generalized Uncertainty Principle (RGUP)

- Modification of the Heisenberg uncertainty relation
 \Rightarrow Position and Momentum not canonically conjugate
- Defining auxiliary canonically conjugate variables

$$p_0^\mu = -i \frac{\partial}{\partial x_{0\mu}}, \quad [x_0^\mu, p_0^\nu] = i\eta^{\mu\nu}$$

- Expression of the position and momentum operators

$$x^\mu = x_0^\mu - \alpha \gamma p_0^\rho p_{0\rho} x_0^\mu + \underbrace{\beta \gamma p_0^\mu p_0^\rho x_{0\rho} + \xi \hbar \gamma p_0^\mu}_{\text{breaks isotropy}},$$

$$p^\mu = p_0^\mu (1 + \varepsilon \gamma p_0^\rho p_{0\rho})$$

- Where $\gamma = \frac{1}{(M_{\text{Pl}c})^2}$ and $\alpha, \beta, \varepsilon$ and ξ are numerical coefficients used to fix the model

Relativistic Generalized Uncertainty Principle (RGUP)

- Position-Momentum commutator

$$[x^\mu, p^\nu] = i\hbar (1 + (\varepsilon - \alpha)\gamma p^\rho p_\rho) \eta^{\mu\nu} + i\hbar(\beta + 2\varepsilon)\gamma p^\mu p^\nu \quad 5$$

- Non-commutative spacetime

$$[x^\mu, x^\nu] = i\hbar\gamma \frac{-2\alpha + \beta}{1 + (\varepsilon - \alpha)\gamma p^\rho p_\rho} (x^\mu p^\nu - x^\nu p^\mu)$$

5. C. Quesne and V. M. Tkachuk, Czech. J. Phys. **56**, 1269 (2006) [quant-ph/0612093].

Poincaré group

- Lorentz generators

$$M^{\mu\nu} = p^\mu x^\nu - p^\nu x^\mu = [1 + (\varepsilon - \alpha)\gamma p_0^\rho p_{0\rho}] (x_0^\mu p_0^\nu - x_0^\nu p_0^\mu)$$

- Poincaré algebra

$$[x^\mu, M^{\nu\rho}] = i\hbar[1 + (\varepsilon - \alpha)\gamma p^\alpha p_\alpha] (x^\nu \delta^{\mu\rho} - x^\rho \delta^{\mu\nu}) + i\hbar 2(\varepsilon - \alpha)\gamma p^\mu M^{\nu\rho}$$

$$[p^\mu, M^{\nu\rho}] = i\hbar[1 + (\varepsilon - \alpha)\gamma p^\alpha p_\alpha] (p^\nu \delta^{\mu\rho} - p^\rho \delta^{\mu\nu})$$

$$[M^{\mu\nu}, M^{\rho\sigma}] = i\hbar(1 + (\varepsilon - \alpha)\gamma p^\alpha p_\alpha) (\eta^{\mu\rho} M^{\nu\sigma} - \eta^{\mu\sigma} M^{\nu\rho} - \eta^{\nu\rho} M^{\mu\sigma} + \eta^{\nu\sigma} M^{\mu\rho})$$

- Special line in parameter space, which preserves the Poincaré algebra

$$\alpha = \varepsilon$$

- Position and momentum operators and their commutation relations

$$[x^\mu, p^\nu] = i\hbar (\eta^{\mu\nu} + 2\alpha\gamma p^\mu p^\nu) \quad \text{RGUP!}$$

$$[x^\mu, x^\nu] = -2i\hbar\alpha\gamma (x^\mu p^\nu - x^\nu p^\mu) \quad \text{NC spacetime}$$

Frame dependence and Composition Law Problems

- Form of the RGUP commutator is the same in every frame

$$U[x^\mu, p^\nu]U^{-1} = U \{ i\hbar (1 + (\varepsilon - \alpha)\gamma p^\rho p_\rho) \eta^{\mu\nu} + i\hbar(\beta + 2\varepsilon)\gamma p^\mu p^\nu \} U^{-1}$$

$$[x'^\mu, p'^\nu] = i\hbar (1 + (\varepsilon - \alpha)\gamma p'^\rho p'_\rho) \eta^{\mu\nu} + i\hbar(\beta + 2\varepsilon)\gamma p'^\mu p'^\nu .$$

- The Casimir invariants and therefore the Dispersion relation of the Modified Poincaré group are quadratic in the physical momentum

$$p^\mu p_\mu \text{ and } W^\mu W_\mu$$

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Fixing the model

- Expression of the position and momentum operators

$$x^\mu = x_0^\mu,$$

$$p^\mu = p_0^\mu (1 + \gamma p_0^\rho p_{0\rho})$$

- We find the following expression for Relativistic GUP $\gamma = \gamma_0 / (M_{PlC})^2$

$$[x^\mu, p^\nu] = i (1 + \gamma p^\rho p_\rho) \eta^{\mu\nu} + i 2\gamma p^\mu p^\nu$$

Particular Poincaré group

- Lorentz generators

$$M^{\mu\nu} = p^\mu x^\nu - p^\nu x^\mu = [1 + \gamma p_0^\rho p_{0\rho}] (x_0^\mu p_0^\nu - x_0^\nu p_0^\mu)$$

- Poincaré algebra

$$[x^\mu, M^{\nu\rho}] = i[1 + \gamma p^\alpha p_\alpha] (x^\nu \delta^{\mu\rho} - x^\rho \delta^{\mu\nu}) + i2\gamma p^\mu M^{\nu\rho}$$

$$[p^\mu, M^{\nu\rho}] = i[1 + \gamma p^\alpha p_\alpha] (p^\nu \delta^{\mu\rho} - p^\rho \delta^{\mu\nu})$$

$$[M^{\mu\nu}, M^{\rho\sigma}] = i(1 + \gamma p^\alpha p_\alpha) (\eta^{\mu\rho} M^{\nu\sigma} - \eta^{\mu\sigma} M^{\nu\rho} - \eta^{\nu\rho} M^{\mu\sigma} + \eta^{\nu\sigma} M^{\mu\rho})$$

- Casimir operator and Dispersion relation

$$E^2 = (pc)^2 + (mc^2)^2 \Rightarrow p^\rho p_\rho = -m^2 \implies p_0^\rho p_{0\rho} (1 + 2\gamma p_0^\sigma p_{0\sigma}) = -m^2$$

- Fourth order differential equation using $p_0^2 = -\partial^\rho \partial_\rho = \square_0$

$$\square_0(1 + 2\gamma \square_0) = -m^2 \quad \text{where} \quad \square_0 = \frac{\partial}{\partial x_0^\mu} \frac{\partial}{\partial x_{0\mu}}$$

Recovering the Lagrangian for scalar and spinor fields

- We use the Ostrogradsky method for recovering the Euler-Lagrange equations for high derivative theories

$$\frac{\partial \mathcal{L}}{\partial \phi} - \partial_{\mu_1} \frac{\partial \mathcal{L}}{\partial (\partial_{\mu_1} \phi)} + \dots + (-1)^m \partial_{\mu_1} \dots \partial_{\mu_m} \frac{\partial \mathcal{L}}{\partial (\partial_{\mu_1} \dots \partial_{\mu_m} \phi)} = 0$$

- The **Equations of Motion** for the Lagrangians we are interested in are

$$\underbrace{\square_0(1 + 2\gamma \square_0)}_{\text{RGUP modified Klein-Gordon equation}} = -m^2$$

$$\underbrace{[i\tau^\mu \partial_\mu (1 + \gamma \partial_\rho \partial^\rho) - m]}_{\text{RGUP modified Dirac equation}} \psi = 0$$

Recovered Lagrangian

- Scalar field Lagrangian

$$\mathcal{L}_{\phi, \mathbb{C}} = \frac{1}{2} (\partial_\mu \phi)^\dagger \partial^\mu \phi - \frac{1}{2} m^2 \phi^\dagger \phi + \underbrace{\gamma \left[(\partial_\nu \partial^\nu \partial^\mu \phi)^\dagger \partial_\mu \phi + \partial_\nu \partial^\nu \partial^\mu \phi (\partial_\mu \phi)^\dagger \right]}_{\text{RGUP corrections}}$$

- Dirac field Lagrangian

$$\mathcal{L}_\psi = \bar{\psi} \left[i \tau^\mu \partial_\mu \underbrace{(1 + \gamma \partial_\rho \partial^\rho)}_{\text{RGUP corrections}} - m \right] \psi$$

- Gauge field Lagrangian

$$\mathcal{L}_A = -\frac{1}{4} F^{\mu\nu} F_{\mu\nu} = -\frac{1}{4} F_0^{\mu\nu} F_{\mu\nu 0} - \underbrace{\frac{\gamma}{2} F_{\mu\nu 0} \partial_\rho \partial^\rho F_0^{\mu\nu}}_{\text{RGUP corrections}}$$

Feynman rules : Propagators

- Scalar field propagator

$$G(x - x') = \int \frac{d^4 p_0}{(2\pi)^4} \frac{-i}{-p_0^2 \underbrace{(1 + \gamma p_0^2)^2}_{\text{RGUP}} + m^2} e^{-ip_0 \cdot (x - x')} .$$

- Dirac propagator

$$G(x - x') = \int \frac{d^4 p_0}{(2\pi)^4} \frac{-i (\tau^\mu p_{0\mu} (1 - \gamma p_{0\rho} p_0^\rho) + m)}{(p_0^\mu p_{0\mu} (1 - \gamma p_{0\rho} p_0^\rho)^2 - m^2)} e^{-ip_0 \cdot (x - x')} .$$

- Gauge field propagator

$$G(x - x') = \int \frac{d^4 q_0}{(2\pi)^4} \frac{-i}{-q_0^2 + 2\gamma q_0^4} e^{-iq_0 \cdot (x - x')} ,$$

Feynman rules : Vertices

- Scalar field : Vertices involving two scalar and up to four gauge bosons

	Powers of γ			
Powers of α	$\alpha^{1/2}$	$\alpha^{1/2}\gamma$	$\alpha^{1/2}\gamma^2$	$\alpha^{1/2}\gamma^3$
	α	$\alpha\gamma$	$\alpha\gamma^2$	$\alpha\gamma^3$
	N/A	$\alpha^{3/2}\gamma$	$\alpha^{3/2}\gamma^2$	$\alpha^{3/2}\gamma^3$
	N/A	$\alpha^2\gamma$	$\alpha^2\gamma^2$	$\alpha^2\gamma^3$
	N/A	$\alpha^{5/2}\gamma$	$\alpha^{5/2}\gamma^2$	$\alpha^{5/2}\gamma^3$
	N/A	$\alpha^3\gamma$	$\alpha^3\gamma^2$	$\alpha^3\gamma^3$

- Dirac field : Vertices involving two fermion and up to three gauge bosons

	Powers of γ	
Powers of α	$\alpha^{1/2}$	$\alpha^{1/2}\gamma$
	N/A	$\alpha\gamma$
	N/A	$\alpha^{3/2}\gamma$

Differential cross sections

- Scalar QED electron-muon scattering

$$\left. \frac{d\sigma}{d\Omega} \right|_{CM} = \frac{1}{4s} \alpha^2 \left(\frac{3 + \cos\theta}{1 - \cos\theta} \right)^2 [1 + 16\gamma E^2(1 - \cos\theta)]$$

- QED electron-muon scattering

$$\left. \frac{d\sigma}{d\Omega} \right|_{CM} = \frac{\alpha^2}{4s} \left[\frac{1}{2} (1 + \cos^2\theta) + \frac{1}{4} \gamma (m_e^2 + m_{\mu\text{on}}^2) (\cos\theta + \cos^2\theta) \right].$$

$$\sigma(e^+e^- \rightarrow \mu^+\mu^-) = \frac{4\pi\alpha^2}{3s} \left[1 + \frac{\gamma}{8} (m_e^2 + m_{\mu\text{on}}^2) \right].$$

- Comparing the corrected amplitude for Xe-Xe scattering measured by ATLAS in CERN⁶ we can put bounds on the parameters

$$\gamma_0 \lesssim 10^{32}$$

6. P. Balek. Nucl. Phys. A, 982:571574, (2019)

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Summary

- We have formulated RGUP leading to **frame independent minimum length** and bypasses the composition law problem
- We formulated and **Effective quantum field theory with minimum length**
- We have calculated the **QG corrected cross sections for electron-muon scattering** and found correlation between the magnitude of the corrections and the mass of the participating particles.
- Applied to existing Quantum experiments give prediction for the scale of Quantum Gravity.

Thank you for your attention !

