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The Tale of H_0 Crisis and the Gravitational Transition

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We propose a late time gravitational transition at low redshifts $z_t < 0.1$ as a possible solution of both the Hubble and growth tensions. Such a transition would naturally lead to a transition of the intrinsic SNIa luminosity and absolute magnitude M at z_t and could also be accompanied by a transition in the dark energy equation of state parameter w . Thus we would have a late $w - M$ phantom transition ($LwMPT$). Such a model does not belong to the category of dark energy models with late time smooth deformations of the Hubble expansion rate $H(z)$, that as we have shown fail to address the growth tension. Therefore, the $LwMPT$ model has the potential of resolving the growth tension by reducing the growth of density perturbations without affecting the Planck/ Λ CDM background expansion. Finally, we offer observational hints for a gravitational transition that would support the $LwMPT$ hypothesis via the study of the evolution of the baryonic Tully-Fisher relation. Specifically, we use a recently published data compilation, finding hints at $\approx 3\sigma$ level for a transition at critical distances $D_c \simeq 9Mpc$ and $D_c \simeq 17Mpc$.

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