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Interior of Typical Black Holes

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We study an evaporating spherically-symmetric black hole that is statistically most likely to form. That can be considered as a black hole formed adiabatically in a heat bath. We construct it by solving conformal matter fields and the semiclassical Einstein equation in a self-consistent manner. Solving the trace component (using the 4D Weyl anomaly) and a general condition of static radial energy flow without specifying explicit boundary conditions determines uniquely the physical static metric for the interior. This represents a dense object with a near-Planck-scale curvature and a surface (instead of a horizon) just outside the Schwarzschild radius. When the object is taken out of the heat bath, it evaporates emitting a Hawking-like radiation. This should be the typical black hole in quantum theory.

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