

Detecting Planetary-mass Primordial Black Holes with Resonant Electromagnetic Gravitational Wave Detectors

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<https://arxiv.org/abs/2012.12189>

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Gravitational waves

- ~ Interferometry : from nHZ to kHz
- ~ Several interests in high-frequency (MHz to THz)
- ~ Possible astrophysical sources : [arXiv:2011.12414](https://arxiv.org/abs/2011.12414)
 - Transient signals or stochastic background
 - Early Universe : preheating, inflation, phase transitions,...
 - Late Universe : NS mergers, exotic compact objects or

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Primordial black holes

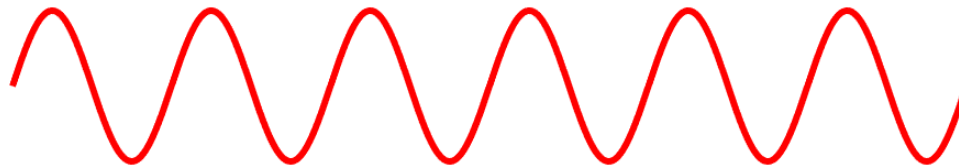
- ~ Dark matter candidate (at least part of)
- ~ Interest in sub-solar detection to point to a primordial origin
- ~ Two mechanisms : primordial binaries or dense PBH halos
- ~ Link between frequency and masses $f_{\text{ISCO}} = \frac{4400 \text{ Hz}}{(m_1 + m_2)/M_{\odot}}$

Electromagnetic waves and GW coupling

- ~ 1962 Gertsenshtein : wave resonance mechanism
- ~ 1970s Grischuck et al. : Using resonant cavities as detectors
- ~ 2019 Ejlli, Cruise et al. : Graviton-photon conversion

GW generation with wave resonance

EMW



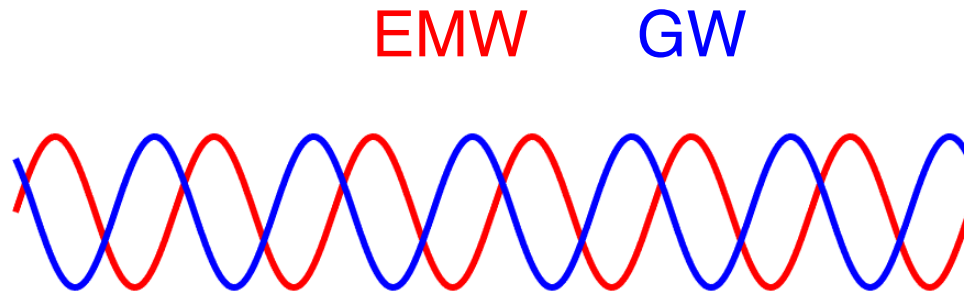
Constant transverse
magnetic field

~ Direct Gersentshtein Effect 1962

~ Amplitude of the generated GW

$$\frac{4GB_0E_0L^2}{c^5\mu_0}$$

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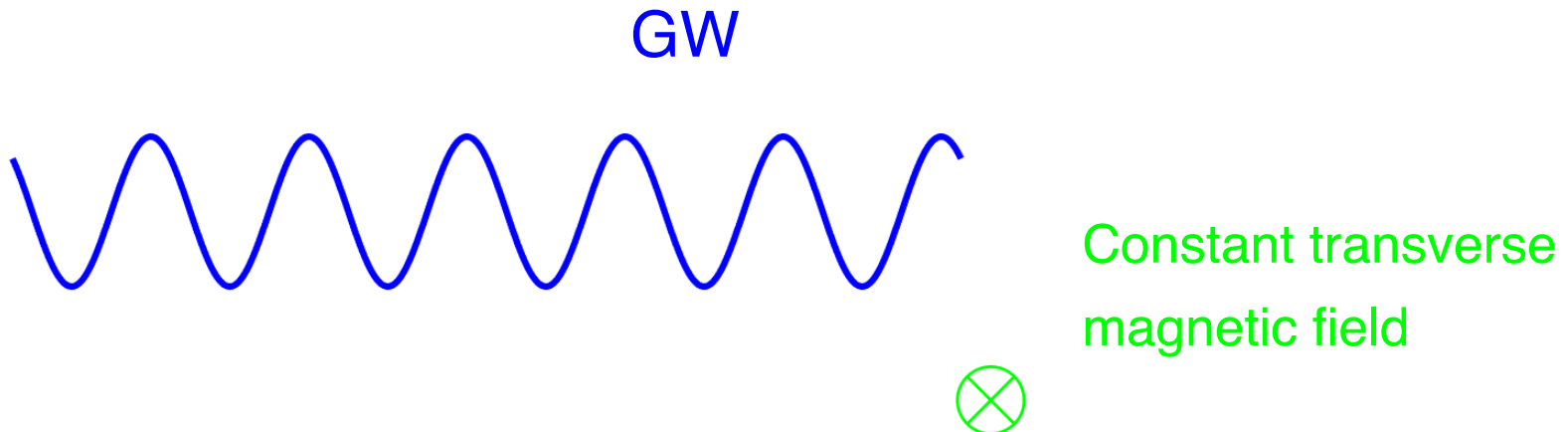
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Direct and inverse Gertsenshtein effects

- ~ Tiny coupling in GW generation $\frac{4G}{c^5 \mu_0} \approx 10^{-46} (T.V.m)^{-1}$
- ~ Generating a strain of $h \approx 10^{-21}$, with $B_0 = 10 \text{ T}$ and $E_0 = 1 \text{ MV/m}$, needs $L \approx 120 \text{ yr}$!!
- ~ Astrophysical application (Zeldovich, 1973)

- ~ Inverse Gertsenshtein Effect **A GW passing through a transverse constant magnetic field produces a faint EM wave**

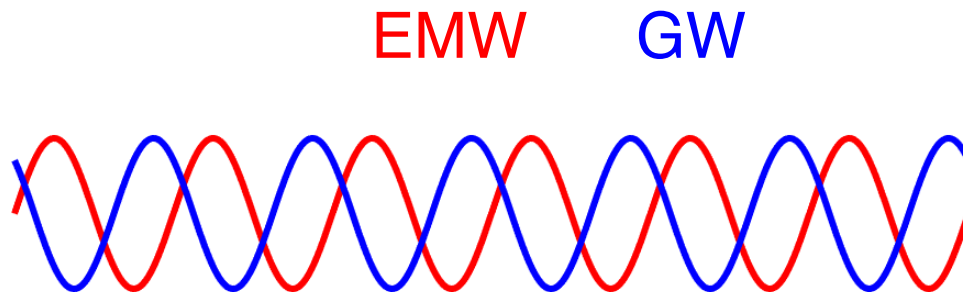


~ Application to the GW detection !

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Constant transverse magnetic field



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Inverse Gertsenshtein Effect

~ Maxwell equations in curved space \Rightarrow EM wave equation

$$g^{\alpha\beta} \nabla_{\alpha} \nabla_{\beta} F_{\mu\nu} + R_{\mu\nu\alpha\beta} F^{\alpha\beta} + R^{\alpha}_{\mu} F_{\nu\alpha} + R^{\alpha}_{\nu} F_{\alpha\mu} = 0$$

~ Background + perturbations

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

$$\partial_{\mu} h^{\mu\alpha} = 0$$

Static external field First order induced field (wave)

$$F_{\mu\nu} = F^{(0)}_{\mu\nu} + F^{(1)}_{\mu\nu}$$

~ Modified Maxwell equations for EM field at first order, with

static field hypothesis $\nabla_{\gamma}^{(\eta)} F^{(0)}_{\alpha\beta} = 0$

$$g^{\alpha\beta} \nabla_{\alpha} \nabla_{\beta} F^{(1)}_{\mu\nu} = -\partial_{\alpha} (\partial_{\mu} h_{\beta\nu} - \partial_{\nu} h_{\beta\mu}) F^{(0)\alpha\beta}$$

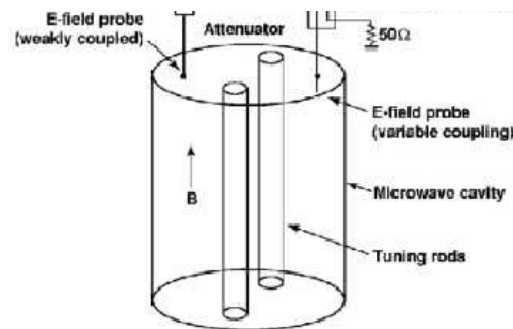
Inverse Gertsenshtein Effect

- ~ Magnetic field transverse to a passing GW
 - ~ theorem by Choquet-Bruhat using plane wave approximation

$$S_{\mu\nu} = 0 \Leftrightarrow J_{\mu}^{\text{eff}} = 0 \Leftrightarrow \Phi_{\alpha} F^{\alpha\mu}{}^{(0)} = q \Phi^{\mu}$$

- ~ same frequency content in the response of the detector
 - => suitable for HFGW signals (like PBH mergers)
 - => different bandwidth than LIGO, LISA, etc.

- ~ problem with ADMX



[arXiv:hep-ex/0701025](https://arxiv.org/abs/hep-ex/0701025)

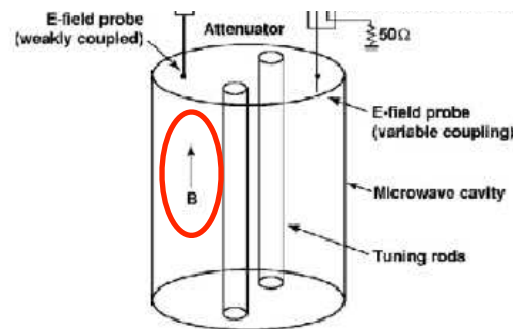
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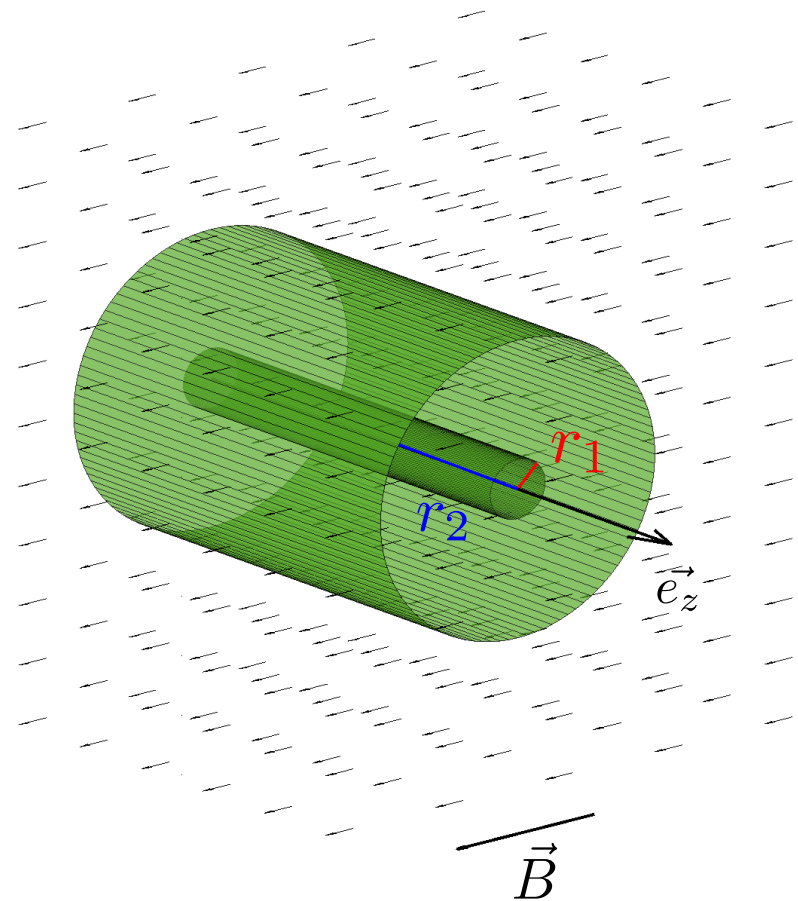
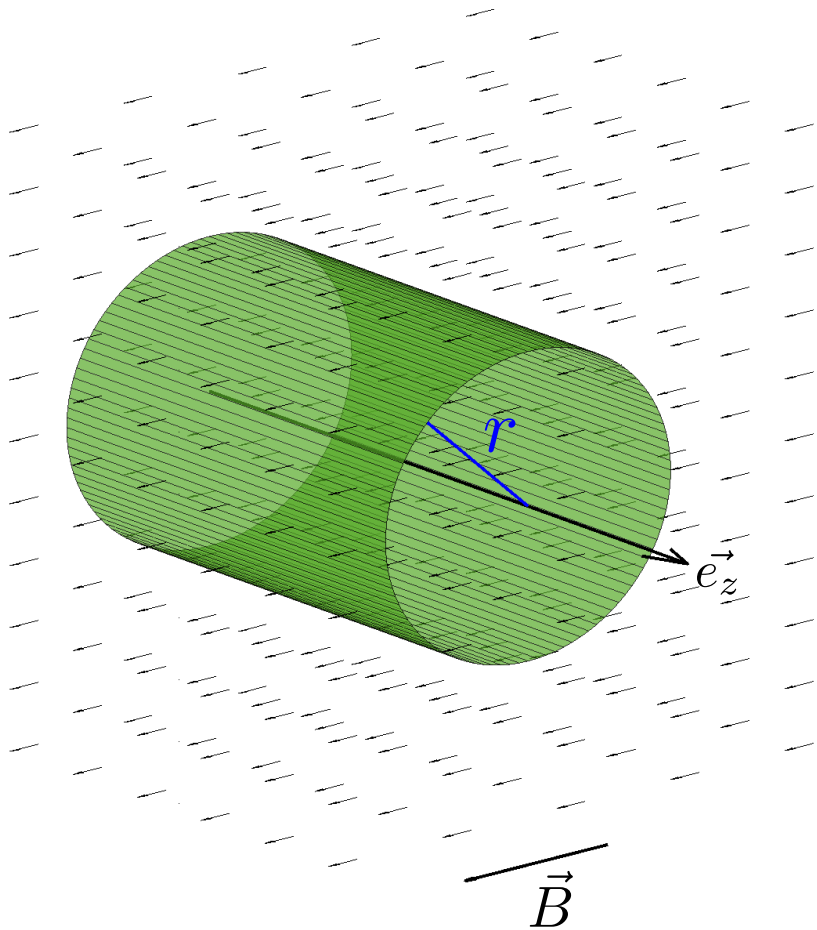


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Case study of two EM GW detectors

~ Two cylindrical cavities : TM (hollow) and TEM (coaxial)

Based on patents PCT/EP2018/086758 & PCT/EP2018/086760



Case study of two EM GW detectors

~ Projection on proper functions of Laplacian

$$\frac{d^2 \hat{b}_{k,m,n}^{r,\phi}}{dt^2} + \Omega_{kn}^2 \hat{b}_{k,m,n}^{r,\phi} = \hat{s}_{k,m,n}^{r,\phi}(t)$$

~ Variation of energy inside the cavity at first order

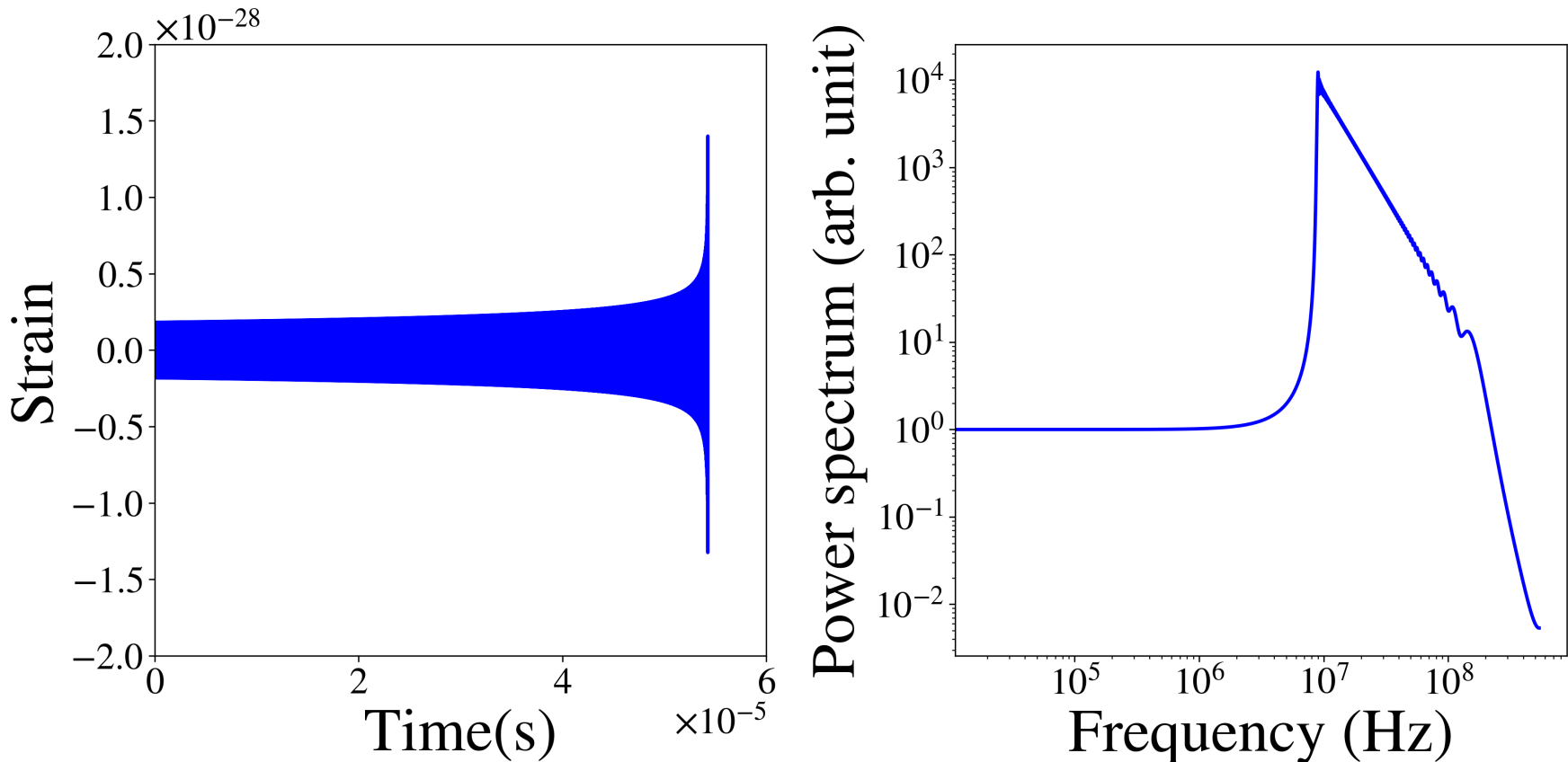
$$\Delta \mathcal{E} = E_{\text{tot}} - E^{(0)} \approx \frac{1}{\mu_0} \int_V \left(\vec{B}^{(0)} \bullet \vec{B}^{(1)} \right) dV$$

$$\Delta \mathcal{E} \approx \frac{2\pi B_0}{\mu_0} \cdot \sum_k \mathcal{I}_k \hat{b}_{k,1,0}(t)$$

$$\hat{s}_{k,1,0}^{r,\phi}(z, t) = \pi B_0 L^2 \mathcal{I}_k \int_{-L/2}^{L/2} \frac{\partial^2 h_+(z, t)}{\partial z^2} dz$$

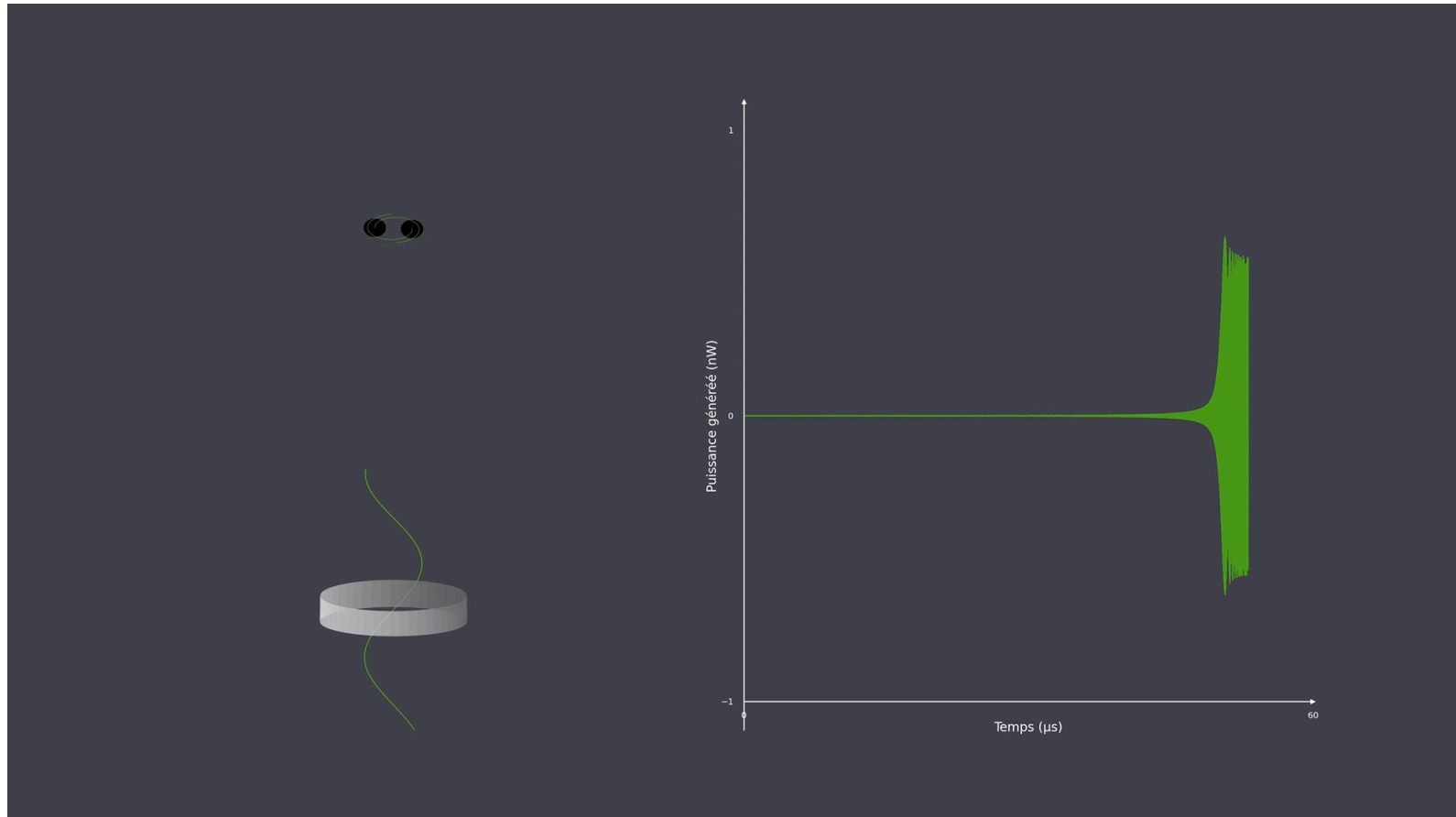
Generation of PBH mergers signals

- ~ Using LALSimulation (PN expansion), start. freq. = $f_{\text{ISCO}}/25$
- ~ Planetary mass : first signal $10^{-5} M_{\odot}$



Resonant response of the detectors $10^{-5} M_{\odot}$

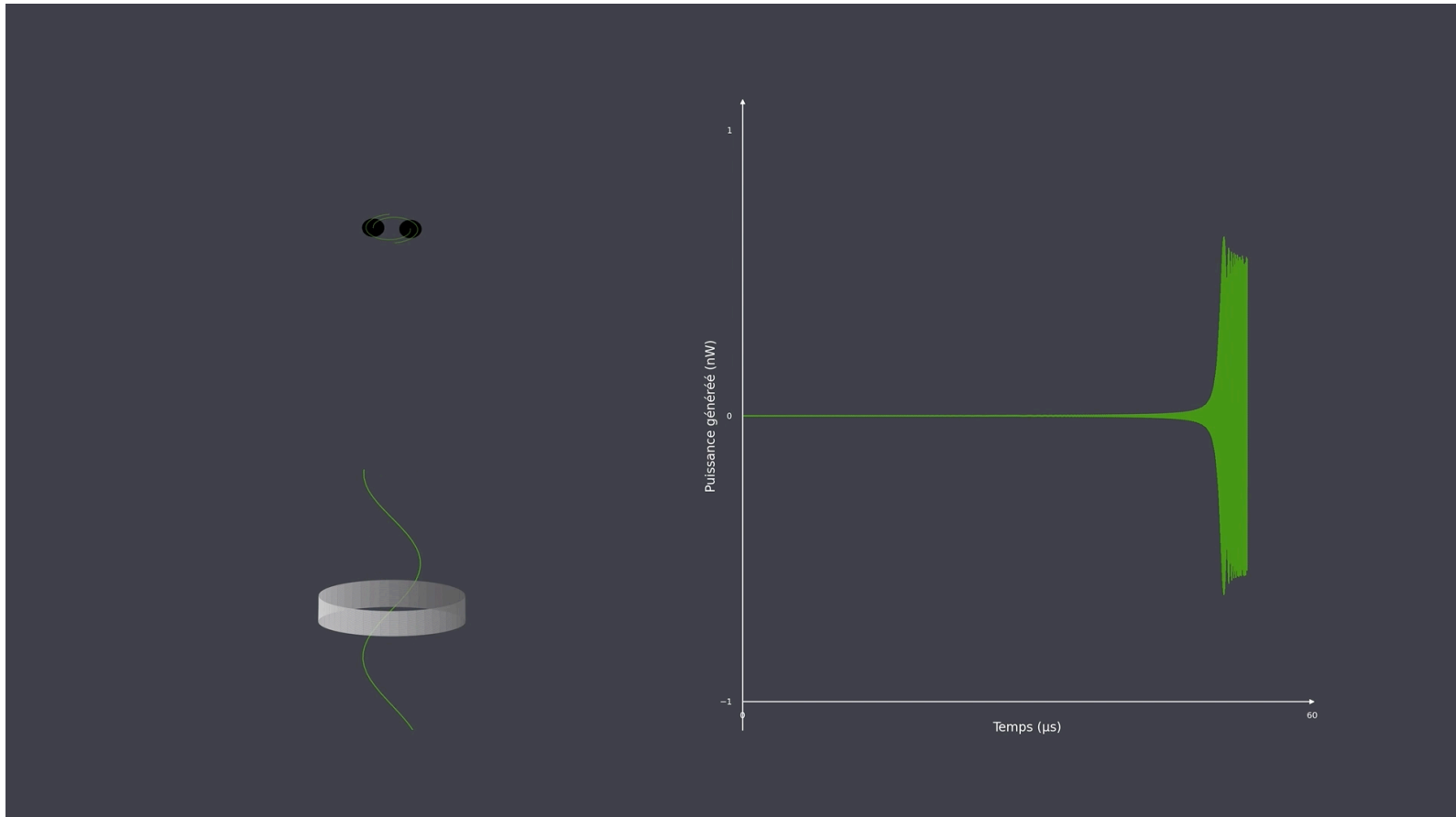
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Other simulation

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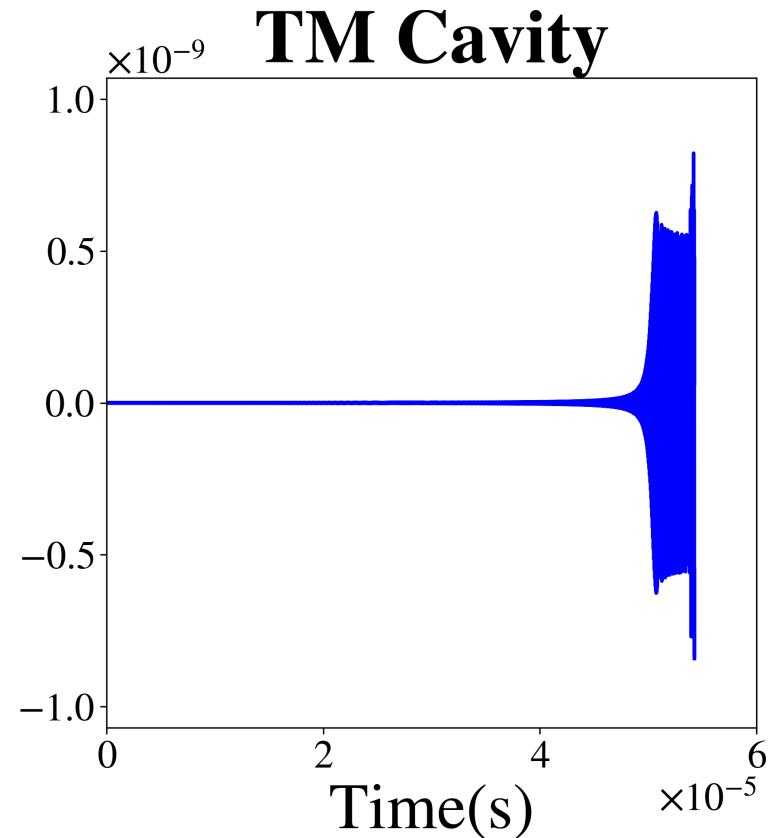
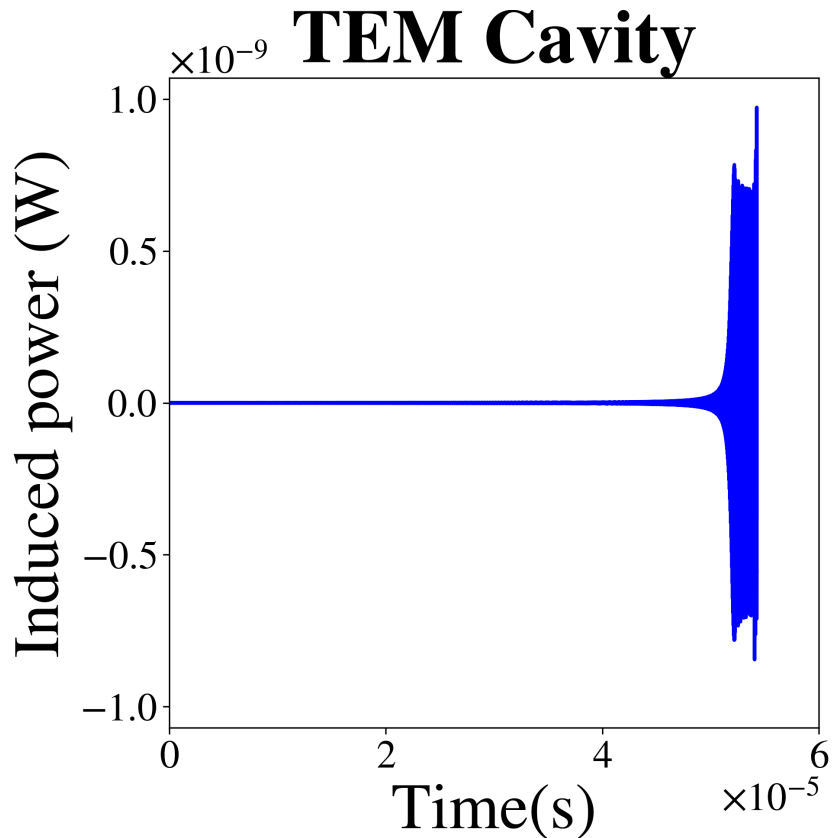
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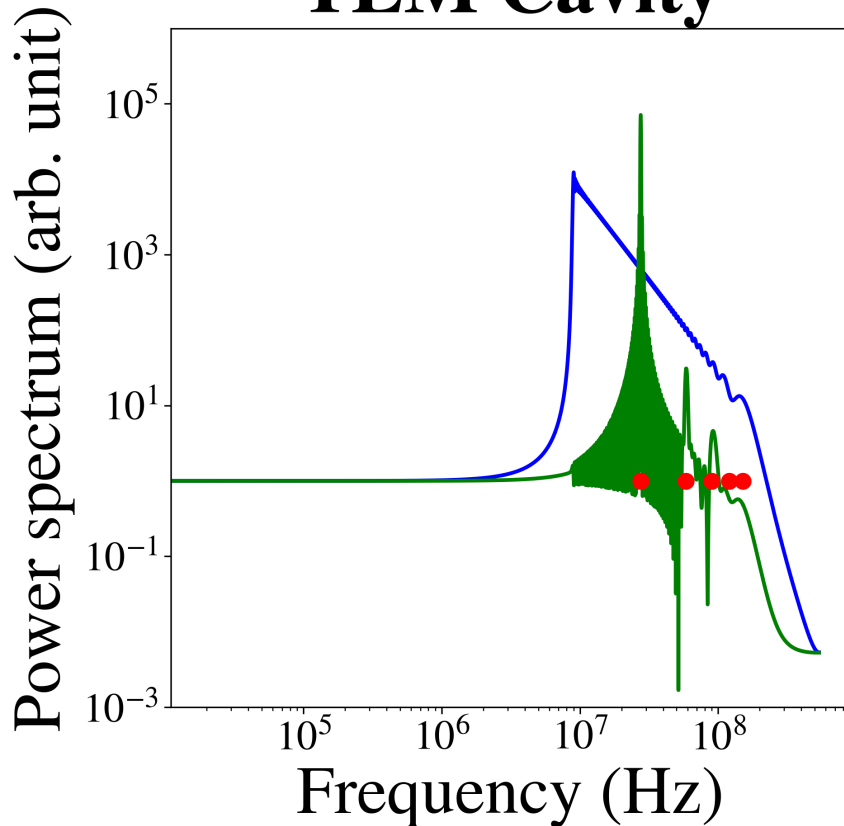


RMS induced power of order 10^{-10} W

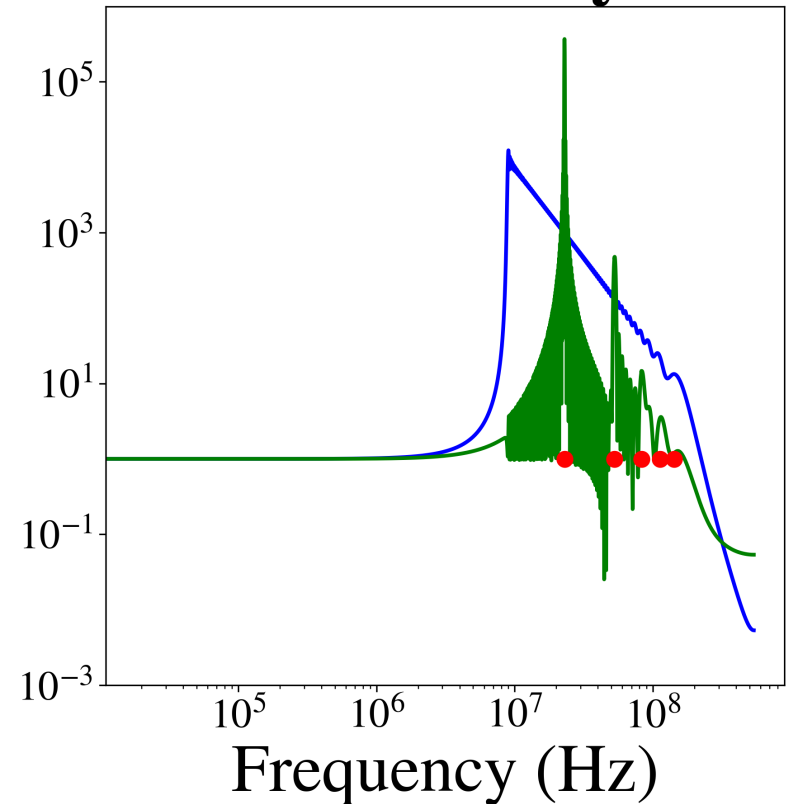
Resonant response of the detectors $10^{-5} M_{\odot}$

$B=5T$ $L=1m$ $r=5m$ $r_1=0.1m$

TEM Cavity

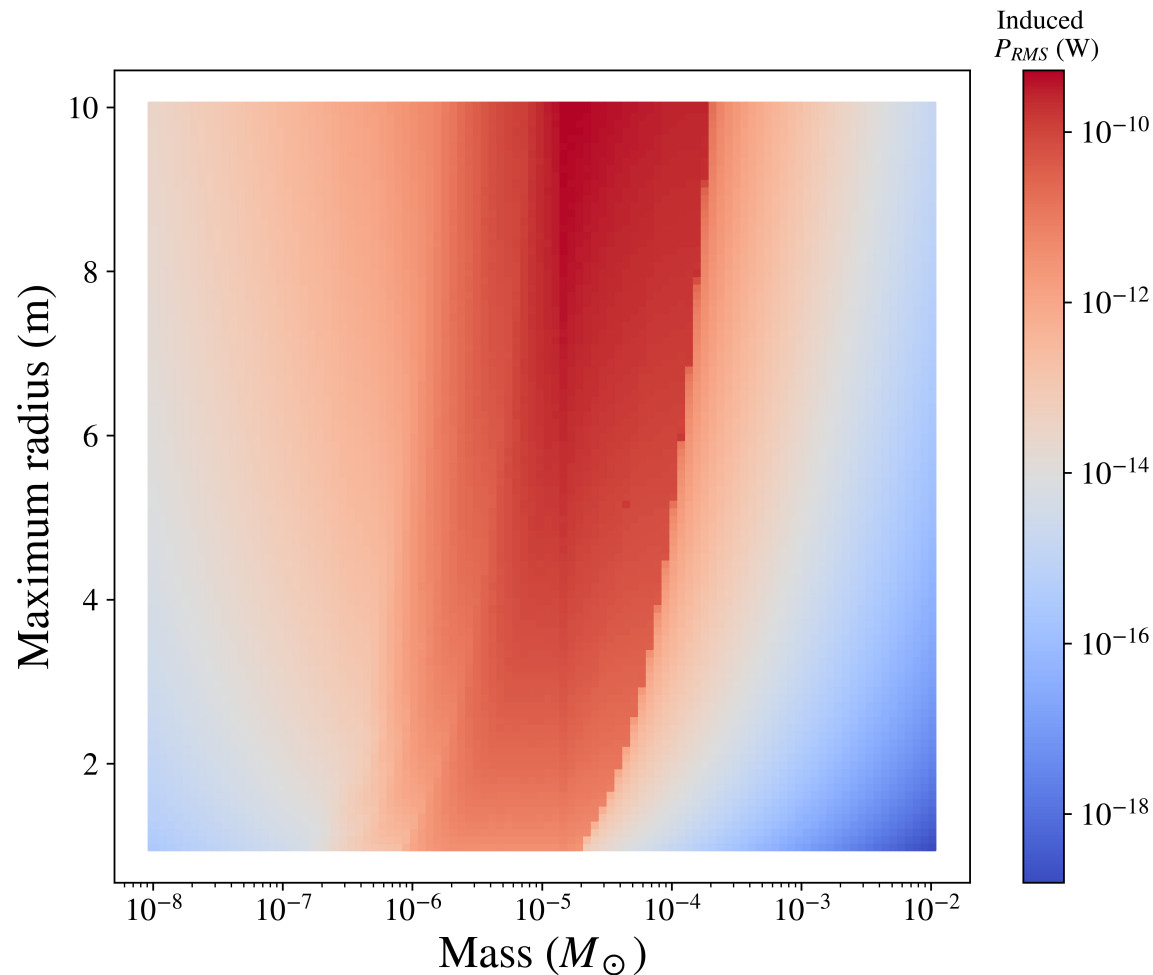


TM Cavity

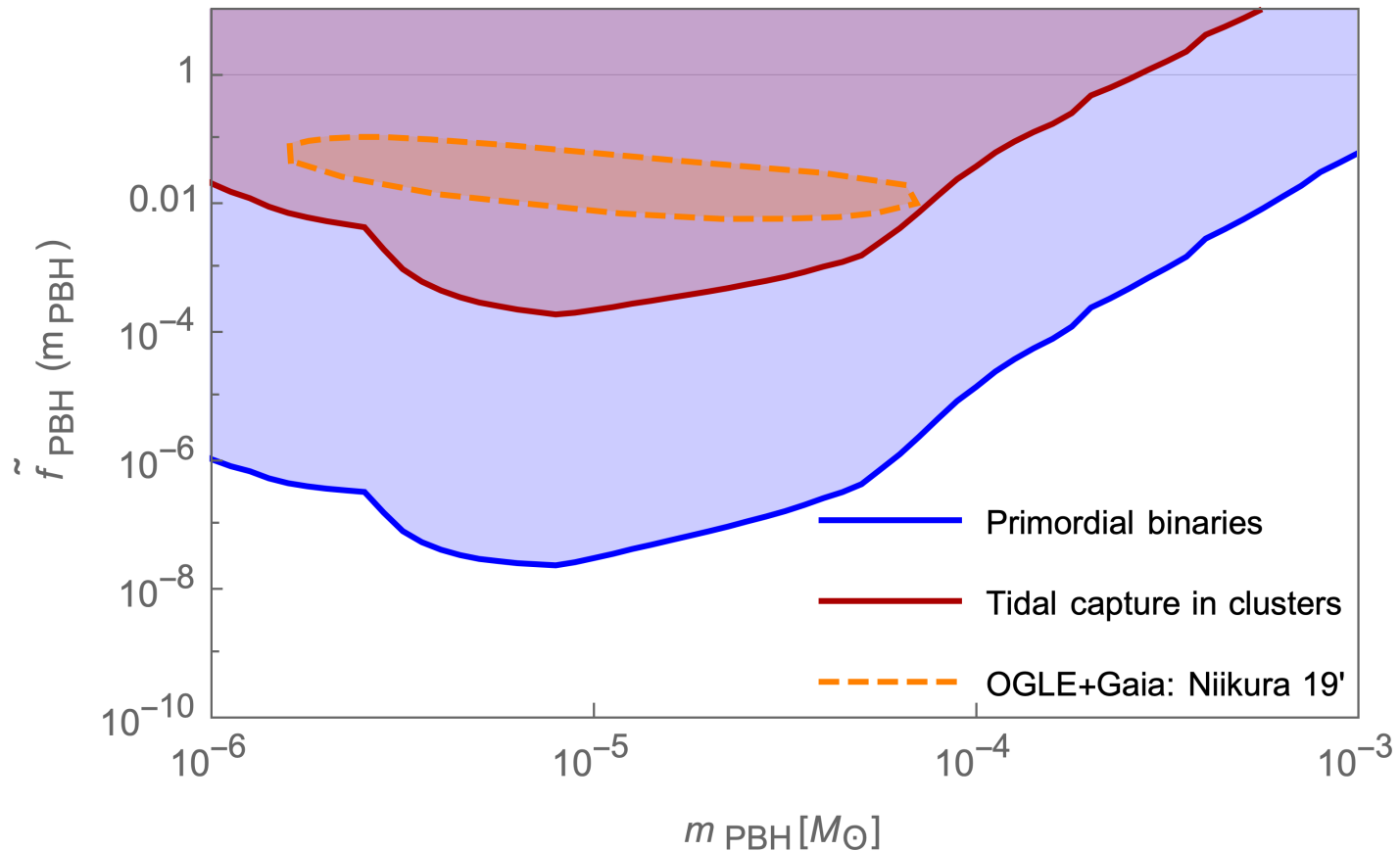


RMS induced power of order 10^{-10} W

Study of the response in function of the mass of the PBH mergers



Study the fraction of DM made of light PBH



Conclusion and perspectives

- ~ Complementary GW detectors
- ~ Study the fraction of DM made of light PBH
- ~ Cosmological GW stochastic background
- ~ Fundamental physics, from early Universe cosmology to exotic compact objects
- ~ Our goal : motivations for experimental development

Thanks for your attention 😊

Bode diagram

