

On the Nature of the Ultrarelativistic Prompt Emission (UPE) Phase of GRB 190114C and GRB 180720B

R. Moradi

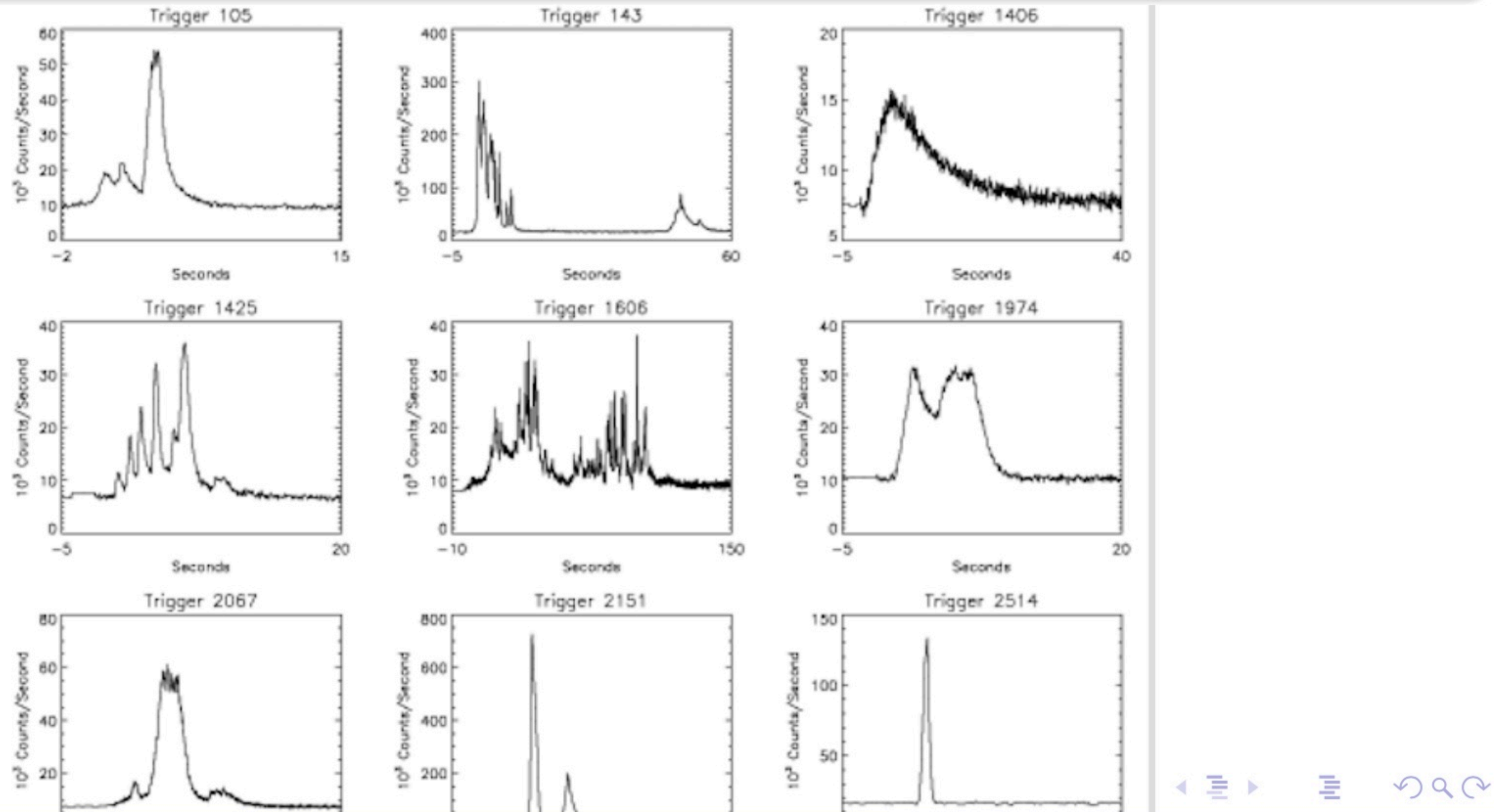
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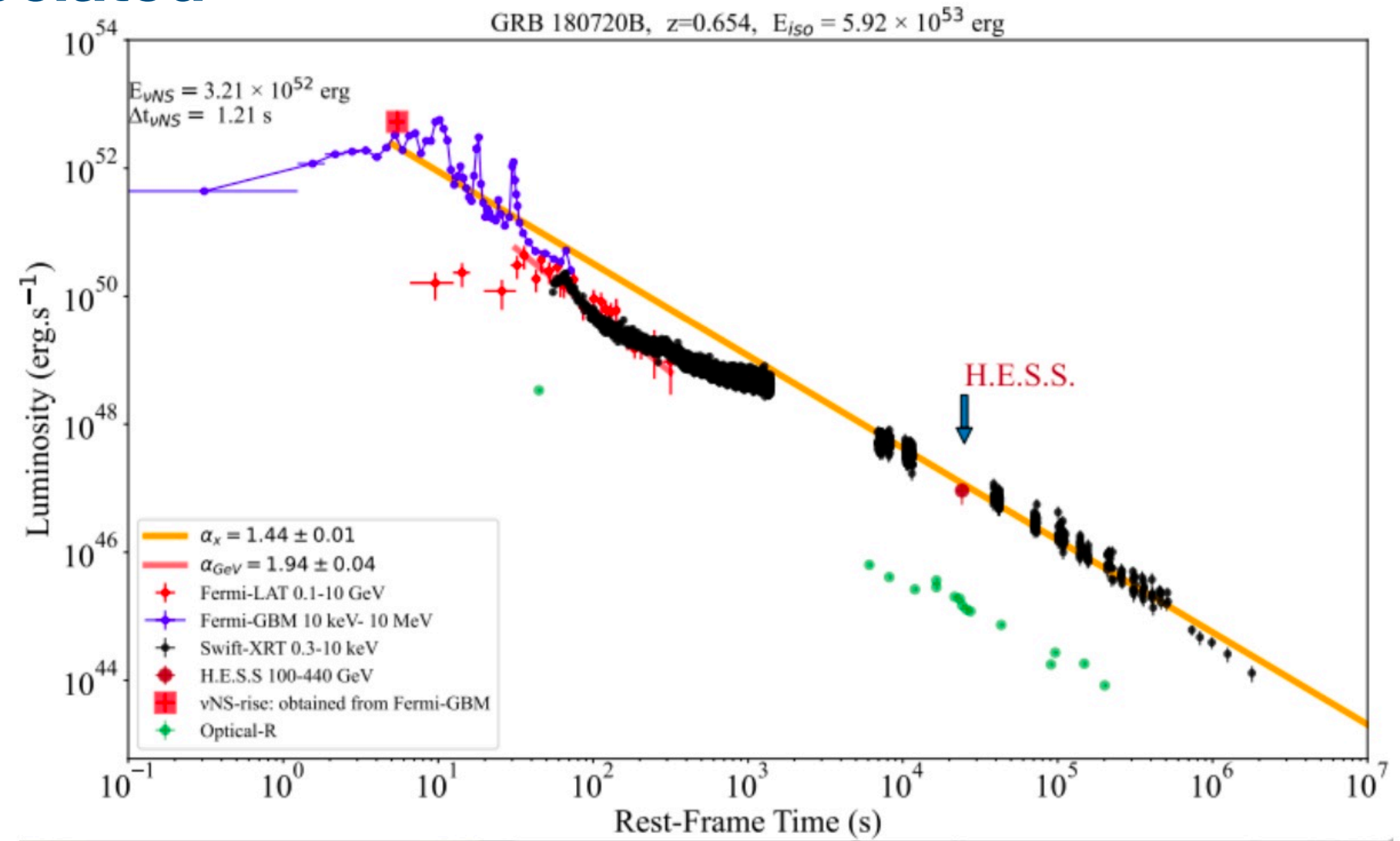
Gamma Ray bursts are

Extremely energetic explosions!

- The most energetic sources in the universe!
- They produce energy equivalent to ~ 1 solar mass in 1-100 second $\rightarrow \sim 10^{54}$ erg!
- if explosion $< 2s \rightarrow$ SHORT GRB.
- if explosion $> 2s \rightarrow$ LONG GRB.
- $\sim 10^{38}$ Tons of TNT explosion! in some seconds!



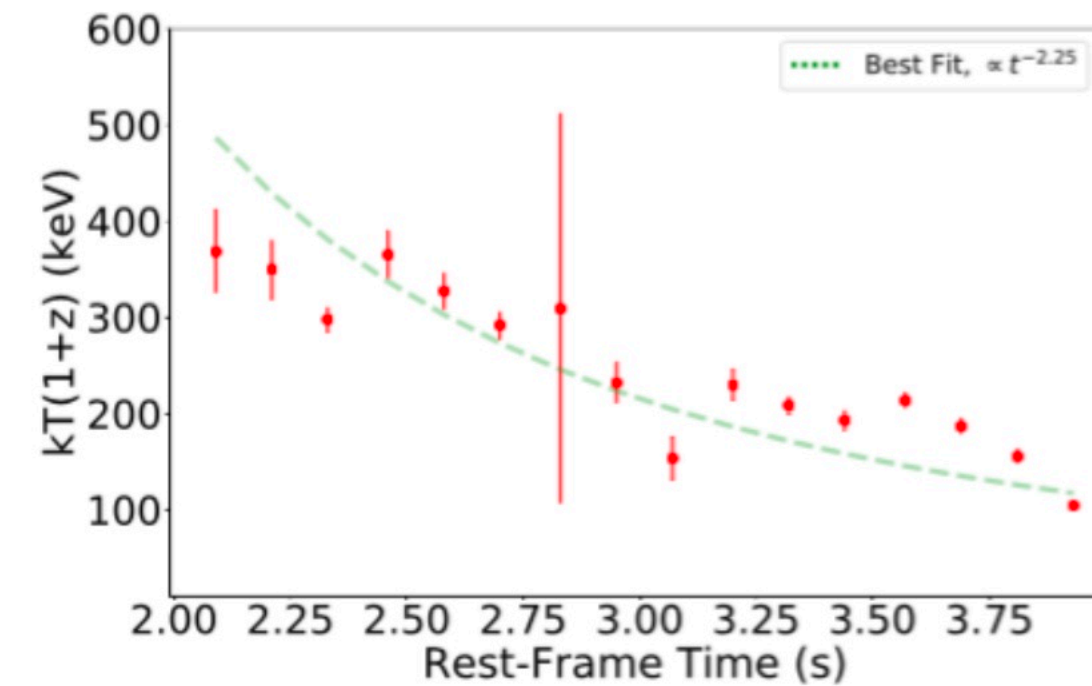
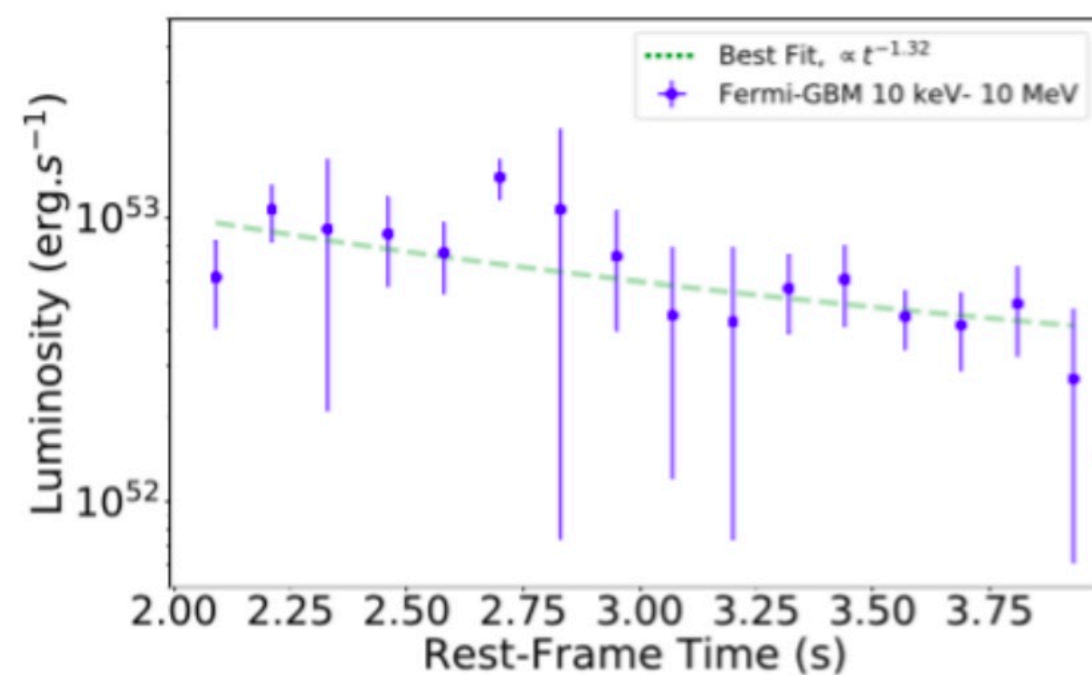
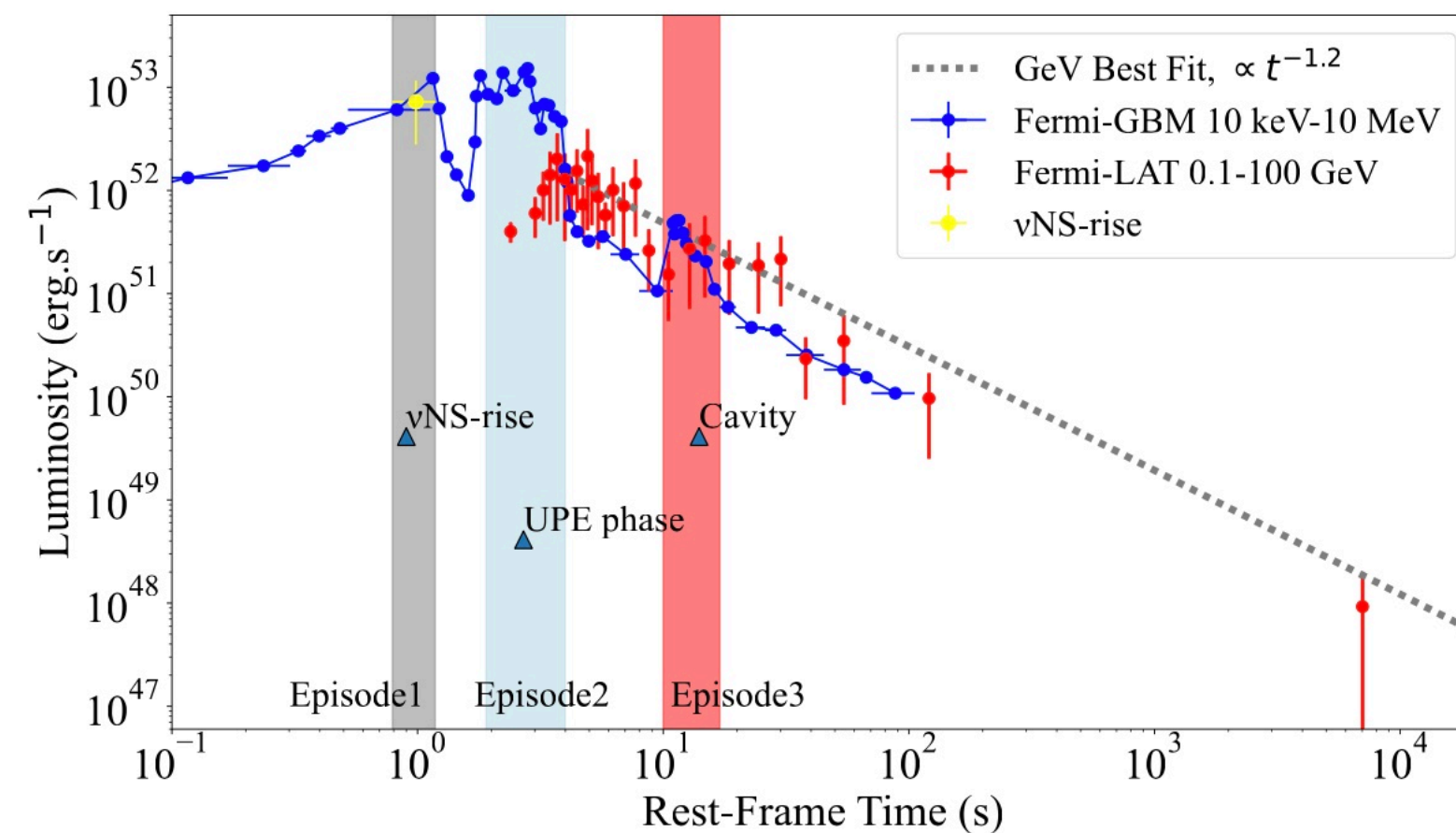
What are the numbers associated with GRBs?



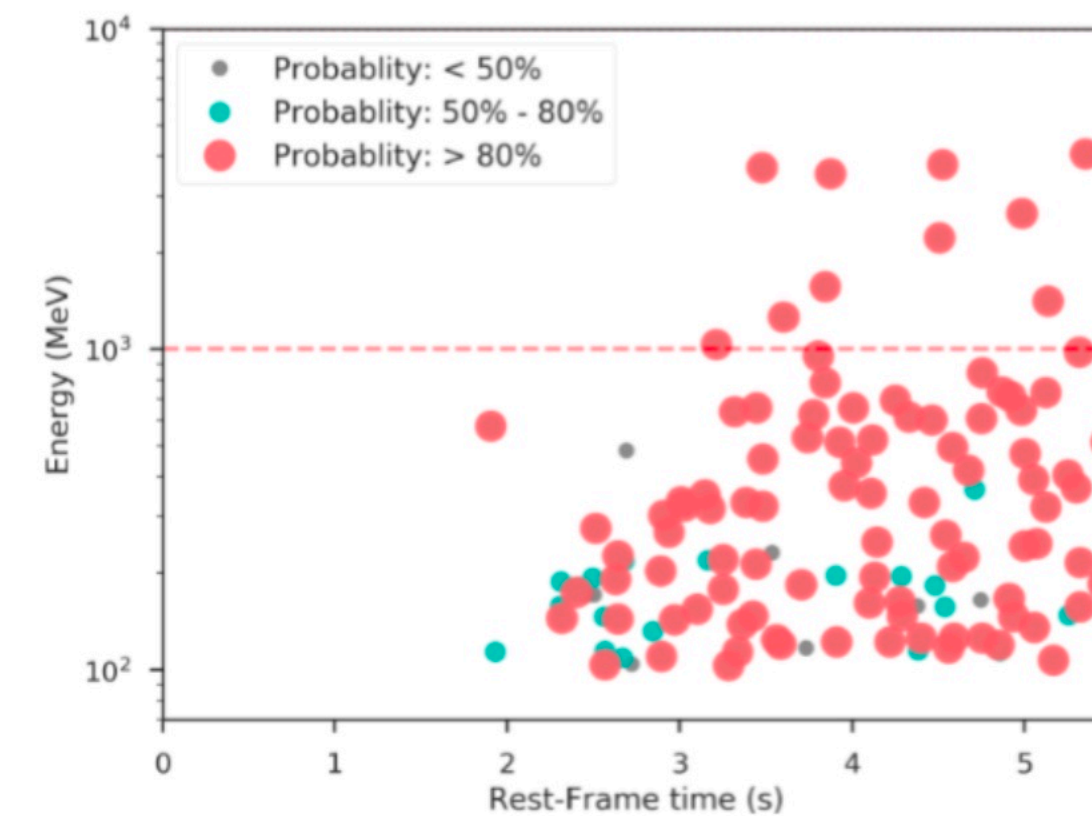
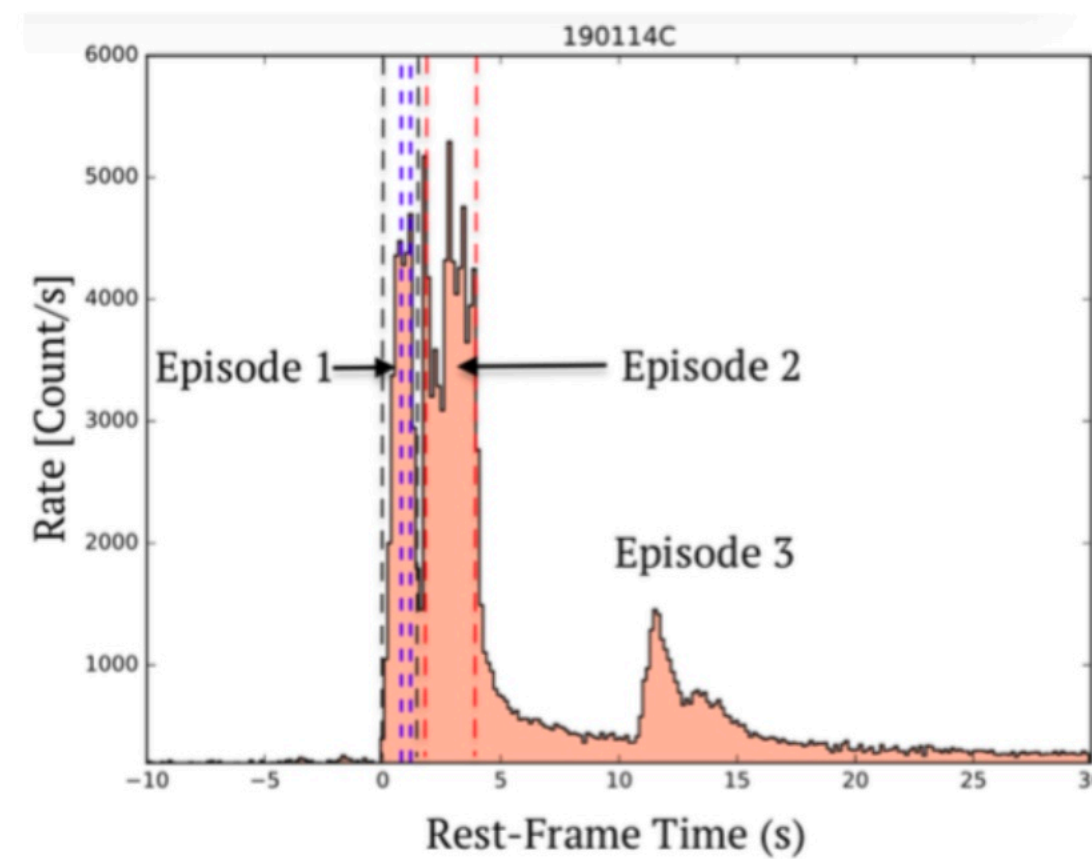
- 180720B
- 18 → 2018
- 07 → July
- 20 → day 20
- B → more than one GRB a day!

GRB 190114C As Binary driven HyperNova of type I (BdHN I)

Moradi, Rueda, Ruffini, Li, et al., PRD 104, 063043 (2021).



$t_1 \sim t_2$	$t_1 \sim t_2$	S	α	E_C	kT	$-\log(\text{posterior}) / (\text{AIC}/\text{BIC})$	F_{GBM}	F_{Total}	F_{Ratio}	F_{Total}
(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)	(11)
2.700-5.500	1.896-3.862	418.62	-0.71 ^{+0.02} _{-0.02}	717.6 ^{+25.4} _{-25.4}	159.0 ^{+3.6} _{-3.6}	-3344/6697/6719	22.49 ^{+2.23} _{-2.23}	111.10 ^{+11.60} _{-11.60}	0.20	1.50e+53
2.700-4.100	1.896-2.879	296.60	-0.51 ^{+0.02} _{-0.02}	696.6 ^{+32.4} _{-32.4}	209.7 ^{+9.1} _{-9.1}	-2675/5360/5381	24.67 ^{+5.35} _{-5.35}	142.50 ^{+21.00} _{-21.00}	0.17	9.64e+52
4.100-5.500	2.879-3.862	318.07	-0.90 ^{+0.02} _{-0.02}	639.3 ^{+31.6} _{-31.6}	130.6 ^{+5.1} _{-5.1}	-2329/5069/5090	25.55 ^{+2.75} _{-2.75}	80.98 ^{+10.07} _{-10.07}	0.32	3.48e+52
2.700-4.400	1.896-2.388	204.30	-0.59 ^{+0.03} _{-0.03}	724 ^{+14.5} _{-14.5}	220.05 ^{+11.1} _{-11.1}	-1182/3774/3796	18.55 ^{+5.92} _{-5.92}	123.90 ^{+39.20} _{-39.20}	0.15	4.19e+52
3.400-4.100	2.388-2.879	235.48	-0.44 ^{+0.04} _{-0.04}	699.8 ^{+17.8} _{-17.8}	196.7 ^{+17.9} _{-17.9}	-2032/4074/4095	31.78 ^{+11.31} _{-11.31}	161.40 ^{+22.30} _{-22.30}	0.20	3.46e+52
4.100-4.800	2.879-3.371	233.97	-0.84 ^{+0.03} _{-0.03}	608.1 ^{+22.1} _{-22.1}	130.4 ^{+5.7} _{-5.7}	-1880/3770/3792	23.94 ^{+1.20} _{-1.20}	85.37 ^{+11.43} _{-11.43}	0.28	2.89e+52
4.800-5.500	3.371-3.862	227.90	-0.96 ^{+0.03} _{-0.03}	679.5 ^{+29.1} _{-29.1}	130.6 ^{+3.1} _{-3.1}	-1809/3628/3649	27.18 ^{+1.01} _{-1.01}	78.20 ^{+11.40} _{-11.40}	0.35	2.65e+52
2.700-3.050	1.896-2.142	148.59	-0.59 ^{+0.03} _{-0.03}	547.7 ^{+14.2} _{-14.2}	240.8 ^{+29.2} _{-29.2}	-1187/2384/2406	19.67 ^{+1.98} _{-1.98}	103.20 ^{+30.60} _{-30.60}	0.19	1.75e+52
3.050-3.400	2.142-2.388	145.04	-0.60 ^{+0.02} _{-0.02}	965.2 ^{+28.8} _{-28.8}	203.5 ^{+14.4} _{-14.4}	-1320/2650/2671	22.87 ^{+2.88} _{-2.88}	152.00 ^{+21.00} _{-21.00}	0.15	2.57e+52
3.400-3.750	2.388-2.633	134.60	-0.63 ^{+0.04} _{-0.04}	885.7 ^{+19.1} _{-19.1}	240.6 ^{+10.9} _{-10.9}	-1224/2458/2480	41.02 ^{+1.91} _{-1.91}	129.10 ^{+33.40} _{-33.40}	0.32	2.18e+52
3.750-4.100	2.633-2.879	187.77	-0.30 ^{+0.05} _{-0.05}	607.8 ^{+40.1} _{-40.1}	151.5 ^{+14.2} _{-14.2}	-1428/2868/2887	23.92 ^{+10.40} _{-10.40}	192.00 ^{+60.30} _{-60.30}	0.12	3.25e+52
4.100-4.450	2.879-3.125	171.81	-0.69 ^{+0.04} _{-0.04}	515.9 ^{+13.4} _{-13.4}	117.3 ^{+5.0} _{-5.0}	-1271/2532/2573	19.13 ^{+4.89} _{-4.89}	92.71 ^{+27.69} _{-27.69}	0.21	1.57e+52
4.450-4.800	3.125-3.371	230.14	-0.94 ^{+0.04} _{-0.04}	702.0 ^{+78.1} _{-78.1}	141.3 ^{+4.8} _{-4.8}	-1254/2518/2539	26.78 ^{+1.47} _{-1.47}	80.78 ^{+17.60} _{-17.60}	0.33	1.37e+52
4.800-5.150	3.371-3.617	166.30	-0.97 ^{+0.04} _{-0.04}	685.1 ^{+68.4} _{-68.4}	140.8 ^{+4.8} _{-4.8}	-1218/2447/2468	31.83 ^{+1.82} _{-1.82}	82.51 ^{+16.62} _{-16.62}	0.39	1.40e+52
5.150-5.500	3.617-3.862	161.51	-0.95 ^{+0.04} _{-0.04}	692.2 ^{+79.1} _{-79.1}	120.0 ^{+4.0} _{-4.0}	-1203/2416/2438	23.19 ^{+1.38} _{-1.38}	73.57 ^{+14.69} _{-14.69}	0.32	1.24e+52
2.700-2.875	1.896-2.019	117.09	-0.58 ^{+0.03} _{-0.03}	470.5 ^{+14.4} _{-14.4}	261.5 ^{+29.0} _{-29.0}	-640/1291/1311	33.68 ^{+2.39} _{-2.39}	112.30 ^{+28.70} _{-28.70}	0.30	9.50e+51
2.875-3.050	2.019-2.142	94.40	-0.68 ^{+0.04} _{-0.04}	627.6 ^{+27.0} _{-27.0}	258.0 ^{+20.1} _{-20.1}	-664/1337/1359	28.45 ^{+2.42} _{-2.42}	98.14 ^{+33.56} _{-33.56}	0.29	8.30e+51
3.050-3.225	2.142-2.265	106.62	-0.59 ^{+0.03} _{-0.03}	937.1 ^{+34.9} _{-34.9}	245.3 ^{+21.0} _{-21.0}	-768/1547/1568	25.71 ^{+1.03} _{-1.03}	169.30 ^{+41.60} _{-41.60}	0.15	1.43e+52
3.225-3.400	2.265-2.388	100.40	-0.70 ^{+0.06} _{-0.06}	1275.9 ^{+215.4} _{-215.4}	208.6 ^{+9.2} _{-9.2}	-669/1349/1369	36.78 ^{+6.93} _{-6.93}	144.90 ^{+37.63} _{-37.63}	0.25	1.29e+52
3.400-3.575	2.388-2.511	98.23	-0.59 ^{+0.04} _{-0.04}	894.0 ^{+106.9} _{-106.9}	255.9 ^{+14.5} _{-14.5}	-702/1414/1436	42.19 ^{+11.43} _{-11.43}	139.30 ^{+34.20} _{-34.20}	0.30	1.16e+52
3.575-3.750	2.511-2.633	93.84	-0.65 ^{+0.05} _{-0.05}	916.0 ^{+114.8} _{-114.8}	229.3 ^{+13.4} _{-13.4}	-730/1471/1492	39.26 ^{+11.59} _{-11.59}	119.50 ^{+30.40} _{-30.40}	0.33	1.01e+52
3.750-3.925	2.633-2.756	126.63	-0.51 ^{+0.02} _{-0.02}	960.9 ^{+30.9} _{-30.9}	204.6 ^{+9.9} _{-9.9}	-808/1627/1648	57.70 ^{+12.81} _{-12.81}	221.10 ^{+48.40} _{-48.40}	0.26	1.87e+52
3.925-4.100	2.756-2.879	141.61	-0.19 ^{+0.02} _{-0.02}	398.7 ^{+12.1} _{-12.1}	216.5 ^{+142.0} _{-142.0}	-734/1479/1499	0.02 ^{+0.23} _{-0.23}	170.20 ^{+12.00} _{-12.00}	0.00	1.44e+52
4.100-4.275	2.879-3.002	122.91	-0.54 ^{+0.06} _{-0.06}	474.1 ^{+45.2} _{-45.2}	162.6 ^{+14.9} _{-14.9}	-758/1526/1547	24.26 ^{+1.79} _{-1.79}	116.10 ^{+32.40} _{-32.40}	0.21	9.82e+51
4.275-4.450	3.002-3.125	122.62	-0.64 ^{+0.08} _{-0.08}	365.0 ^{+48.5} _{-48.5}	107.5 ^{+17.0} _{-17.0}	-675/1360/1380	9.04 ^{+4.69} _{-4.69}	72.20 ^{+19.00} _{-19.00}	0.13	6.11e+51
4.450-4.625	3.125-3.248	111.94	-1.04 ^{+0.08} _{-0.08}	640.0 ^{+108.1} _{-108.1}	161.0 ^{+11.9} _{-11.9}	-640/1290/1310	22.34 ^{+1.95} _{-1.95}	68.54 ^{+11.21} _{-11.21}	0.33	5.80e+51
4.625-4.800	3.248-3.371	123.33	-0.90 ^{+0.05} _{-0.05}	694.2 ^{+34.2} _{-34.2}	146.3 ^{+8.6} _{-8.6}	-734/1477/1499	35.39 ^{+4.46} _{-4.46}	89.91 ^{+18.82} _{-18.82}	0.40	7.60e+51
4.800-4.975	3.371-3.494	129.63	-0.86 ^{+0.06} _{-0.06}	864.0 ^{+119.9} _{-119.9}	135.3 ^{+11.9} _{-11.9}	-744/1498/1519	30.78 ^{+11.55} _{-11.55}	96.50 ^{+22.99} _{-22.99}	0.32	8.17e+51
4.975-5.150	3.494-3.617	107.30	-1.18 ^{+0.09} _{-0.09}	820.2 ^{+115.9} _{-115.9}	149.7 ^{+14.9} _{-14.9}	-683/1378/1398	42.75 ^{+11.12} _{-11.12}	71.51 ^{+17.99} _{-17.99}	0.46	6.05e+51
5.150-5.325	3.617-3.739	108.96	-1.04 ^{+0.06} _{-0.06}	765.2 ^{+119.0} _{-119.0}	130.9 ^{+8.8} _{-8.8}	-697/1404/1426	26.14 ^{+7.02} _{-7.02}	66.70 ^{+20.49} _{-20.49}	0.39	3.64e+51
5.325-5.500	3.739-3.862	121.57	-0.88 ^{+0.06} _{-0.06}	635.3 ^{+98.7} _{-98.7}	108.9 ^{+15.1} _{-15.1}	-736/1483/1504	20.90 ^{+1.81} _{-1.81}	79.48 ^{+28.02} _{-28.02}	0.26	6.72e+51
5.500-5.675	3.862-3.985	111.83	-1.00 ^{+0.08} _{-0.08}	437.7 ^{+82.2} _{-82.2}	73.2 ^{+2.9} _{-2.9}	-587/1184/1205	12.45 ^{+1.10} _{-1.10}	43.09 ^{+17.73} _{-17.73}	0.29	3.64e+51

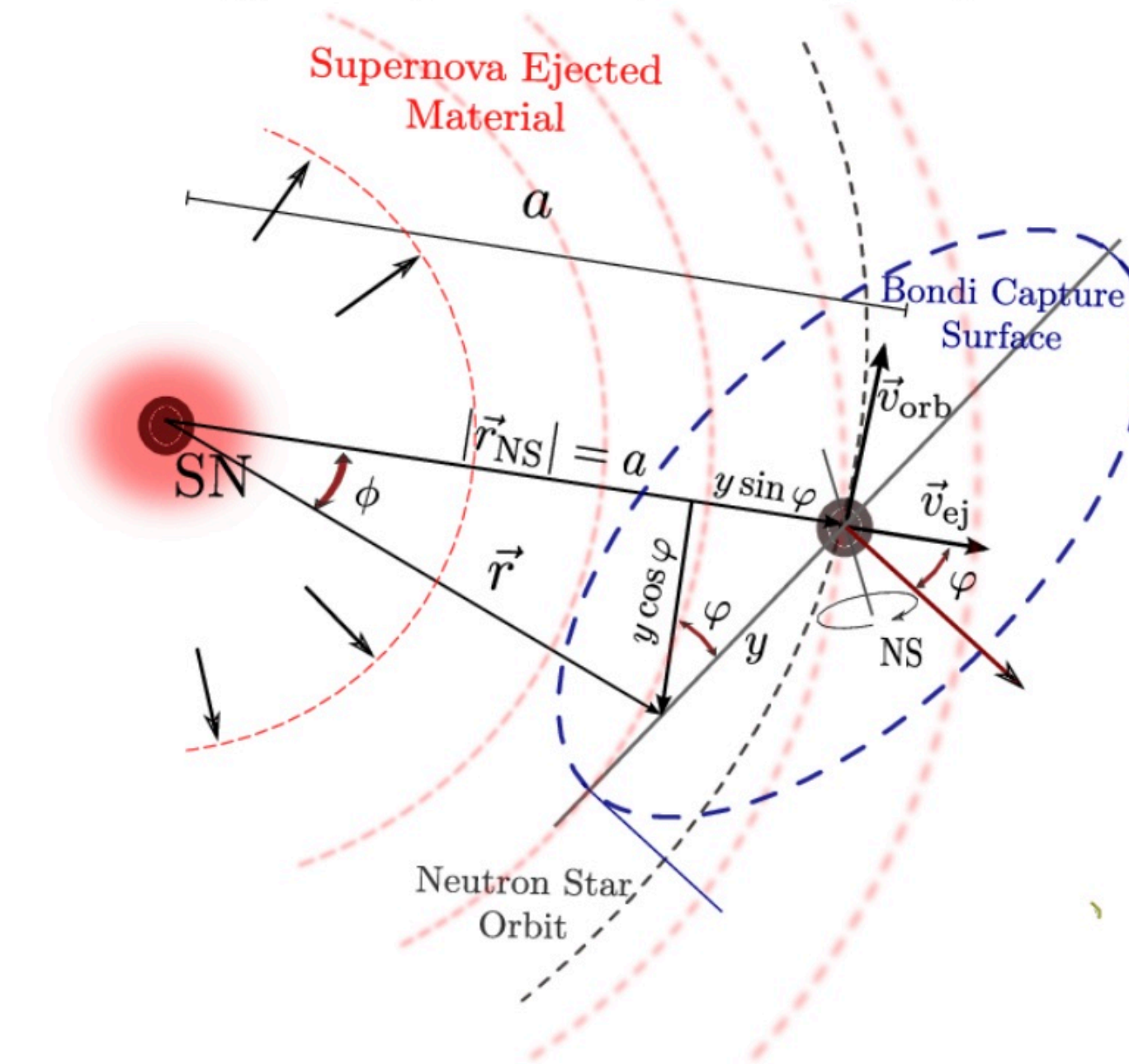


- At 20:57:02.63 UT on 14 January 2019, Fermi-GBM was triggered by GRB 190114C. . Following the 0.37 second at 20:57:03 UT on 14 January 2019, the Neil Gehrels Swift Burst Alert Telescope (BAT) was triggered as well by GRB 190114C.
- The Fermi-LAT had a boresight angle of 68 deg at the trigger time, the GRB remained in the field of view of Fermi-LAT for 150 seconds.
- Nordic Optical Telescope (NOT) announced by GCN23695 the redshift of 0.424.
- At time 15:29:54 GMT on January 15, 2019 we identified by GCN23715 this GRB as a BdHN I, and predicted that an optical SN should appear in the same location of the GRB within 18.8 ± 3.7 days, which indeed was confirmed by Melandri et al.
- This successful prediction and the following detection of TeV radiation by MAGIC have made GRB 190114C as a prototype which all the BdHN phases have been observed.

BdHN I includes three different components:

- 1 a CO core undergoing a SN explosion in presence of a binary NS companion;
- 2 an additional NS originating from the SN explosion indicated as a ν NS (the newborn NS at the center of the SN), accreting the SN ejecta and giving origin to the afterglow;
- 3 the formation of the BH by the hypercritical accretion of the SN ejecta onto the NS reaching its critical mass. The newborn BH originates the GeV emission. Coincidence of these effect, the birth of BH and the onset of the GeV emission has been narrow down by our to 10^{-6} s.

Fryer, et al., ApJ 526, 152 (1999)
Bromberg, et al., ApJ 749, 110(2012)
Götberg, et al., AA 629, A134 (2019)



Rueda & Ruffini, ApJ 758, L7 (2012)
Becerra, et al., ApJ 812, 100 (2015)
Becerra, et al., ApJ 871, 14 (2019)

BdHN model in movies

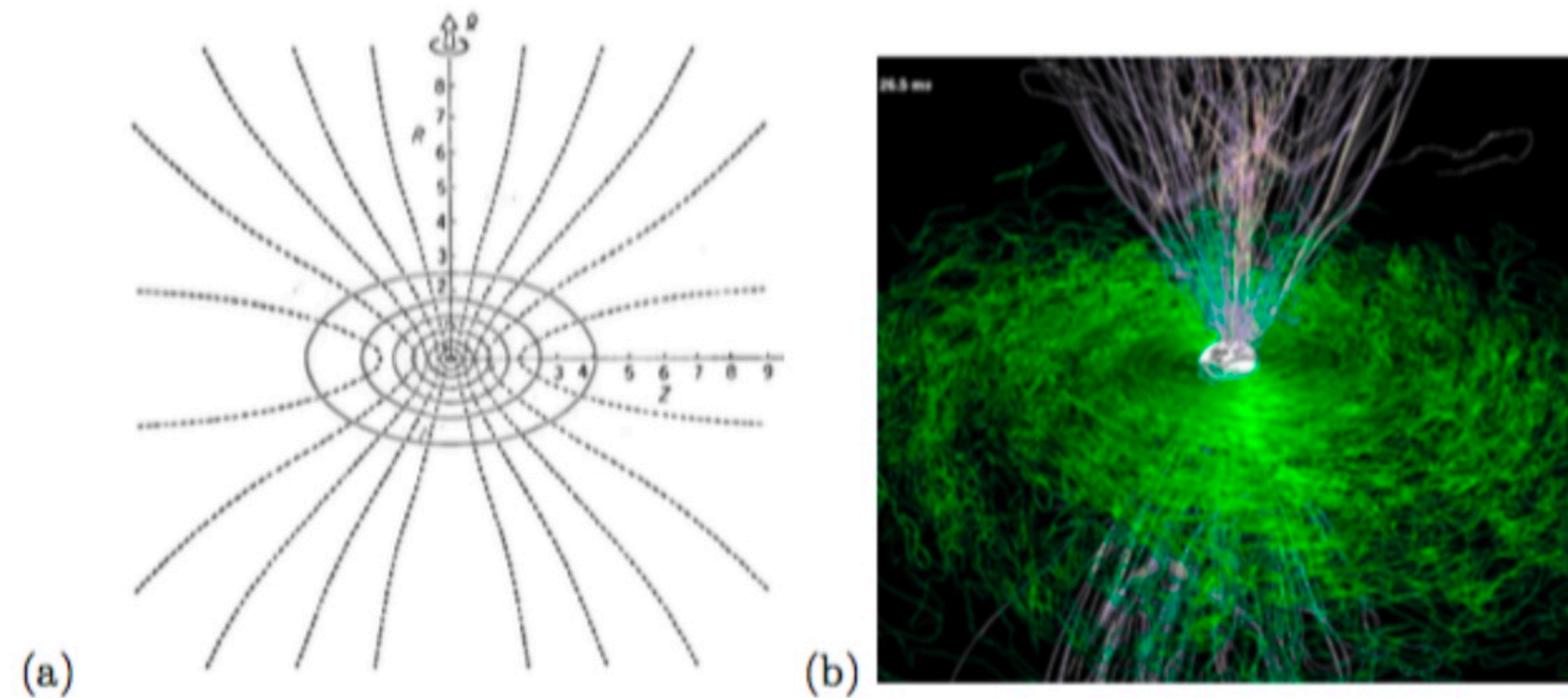
The Correspondence 2016
By J. Tornatore



Olga Kurylenko and Jeremy Irons: The Correspondence

Black Hole

a) Wilson, J. R. 1975, in Annals of the New York Academy of Sciences



b) Rezzolla, L., Giacomazzo, B., Baiotti, L., et al. 2011, ApJL,732, L6

Kerr Metric in a uniform magnetic field

The Kerr space-time metric (geometric units will be considered), which is stationary and axisymmetric, in standard Boyer-Lindquist (BL) coordinates reads

$$ds^2 = - \left(1 - \frac{2Mr}{\Sigma} \right) dt^2 - \frac{4aMr \sin^2 \theta}{\Sigma} dt d\phi + \frac{\Sigma}{\Delta} dr^2 \\ + \Sigma d\theta^2 + \left[r^2 + a^2 + \frac{2Mra^2 \sin^2 \theta}{\Sigma} \right] \sin^2 \theta d\phi^2,$$

where $\Sigma = r^2 + a^2 \cos^2 \theta$ and $\Delta = r^2 - 2Mr + a^2$. The (outer) event horizon is located at $r_+ = M + \sqrt{M^2 - a^2}$.



Accretion to NS -> BH formation+Magnetic field

Killing vectors and Maxwell field

$$\mathcal{L}_\zeta g_{\mu\nu} = \zeta_{\mu;\nu} + \zeta_{\nu;\mu}$$

- Papapetrou: Killing vector in vacuum space-time generates a solution of Maxwell's equations in that space-time.
- The solution of electromagnetic test-field which occurs when a stationary, axisymmetric black hole is placed in an originally uniform magnetic field of strength B_0 aligned along the symmetry axis of rotation of the black hole is

$$\mathbf{F} = \frac{1}{2} B_0 (d\psi + \frac{2J}{M} d\eta)$$

J: angular momentum, M: mass of the black hole, ψ : axial Killing vector, η : timelike Killing vector.

The electromagnetic field of the *inner engine* in the Carter's orthonormal tetrad is:

$$E_{\hat{r}} = \frac{\hat{a}B_0}{\Sigma} \left[r \sin^2 \theta - \frac{\hat{M} (\cos^2 \theta + 1) (r^2 - \hat{a}^2 \cos^2 \theta)}{\Sigma} \right],$$

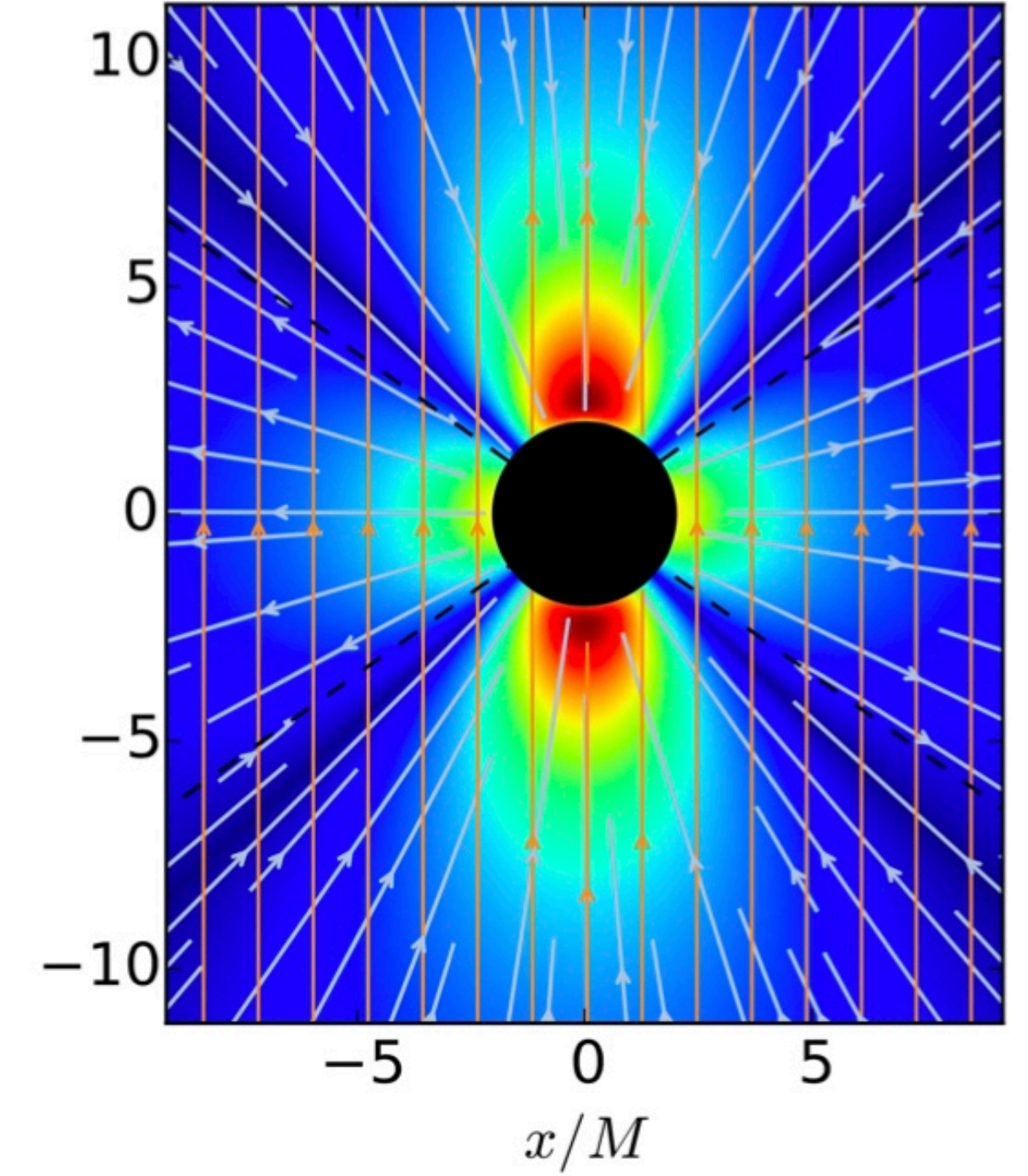
$$E_{\hat{\theta}} = \frac{\hat{a}B_0}{\Sigma} \sin \theta \cos \theta \sqrt{\Delta},$$

$$B_{\hat{r}} = -\frac{B_0 \cos \theta}{\Sigma} \left[-\frac{2\hat{a}^2 \hat{M} r (\cos^2 \theta + 1)}{\Sigma} + \hat{a}^2 + r^2 \right],$$

$$B_{\hat{\theta}} = \frac{B_0 r}{\Sigma} \sin \theta \sqrt{\Delta},$$

where $\Sigma = r^2 + \hat{a}^2 \cos^2 \theta$, $\Delta = r^2 - 2\hat{M}r + \hat{a}^2$, $\hat{M} = GM/c^2$, $\hat{a} = a/c = J/(Mc)$, being M and J the mass and angular momentum of the Kerr BH. The (outer) event horizon is located at $r_+ = (\hat{M} + \sqrt{\hat{M}^2 - \hat{a}^2})$.

The inner engine



The newborn black hole in GRB 191014C proves that it is alive

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ABSTRACT

A multi-decade theoretical effort has been devoted to finding an efficient mechanism to use the rotational and electric extractable energy of a Kerr-Newman black hole (BH), to power the most energetic astrophysical sources such as gamma-ray bursts (GRBs) and active galactic nuclei. We show an efficient general relativistic electro-dynamical process which occurs in the “engine” of a binary driven hypernova. The inner engine is composed of a rotating Kerr BH of mass M and dimensionless spin parameter α , a magnetic field of strength B_0 aligned and parallel to the rotation axis, and a very low-density ionized plasma. We show that the gravitomagnetic interaction between the BH and the magnetic field induces an electric field that accelerates electrons and protons from the environment to ultrarelativistic energies emitting synchrotron radiation. We show that in GRB 190114C of mass $M = 4.4 M_\odot$, $\alpha = 0.4$, and $B_0 \approx 4 \times 10^{10}$ G can lead to a high-energy ($> \text{GeV}$) luminosity of $10^{51} \text{ erg s}^{-1}$. The inner engine is

Study in the undercritical E field regime

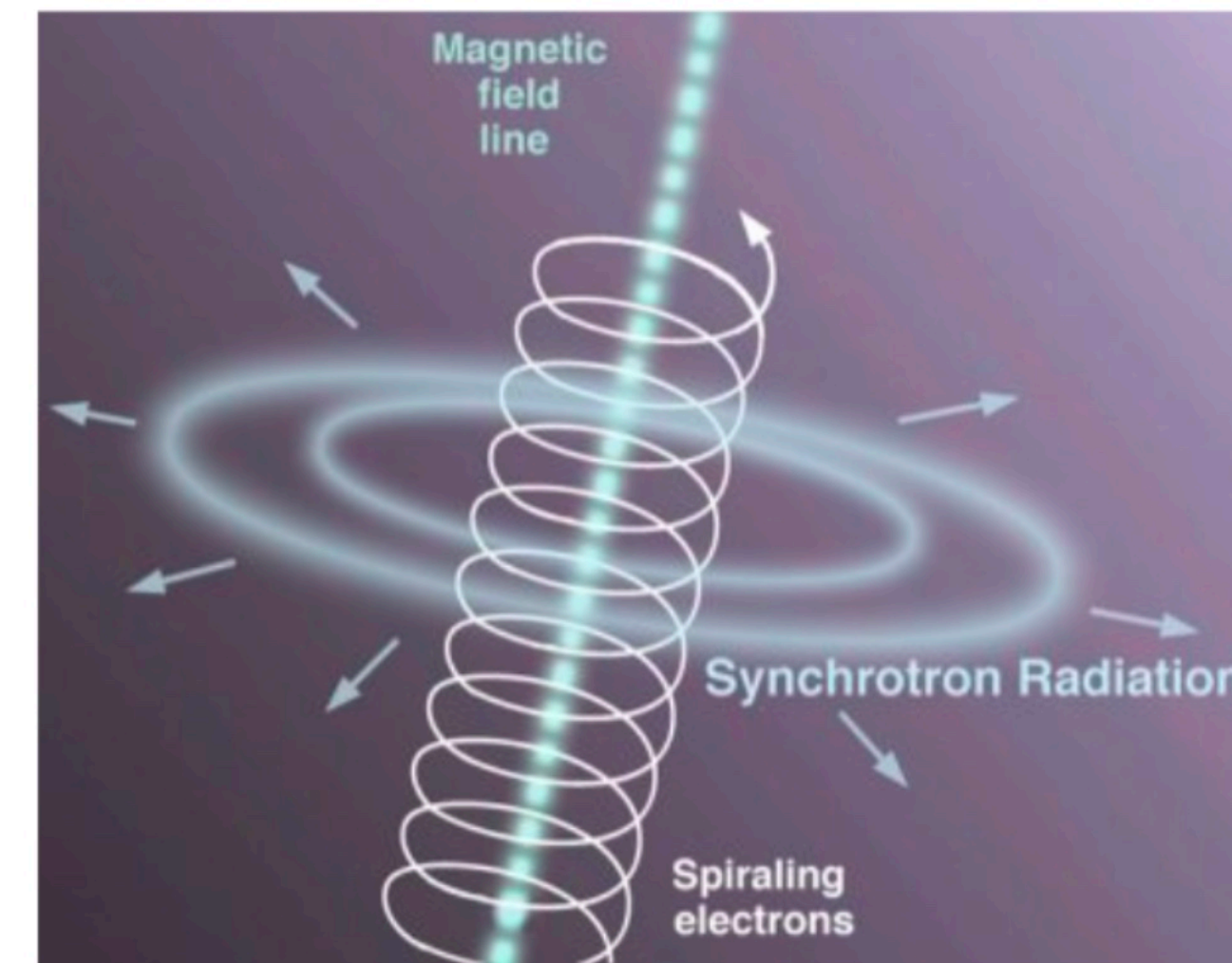
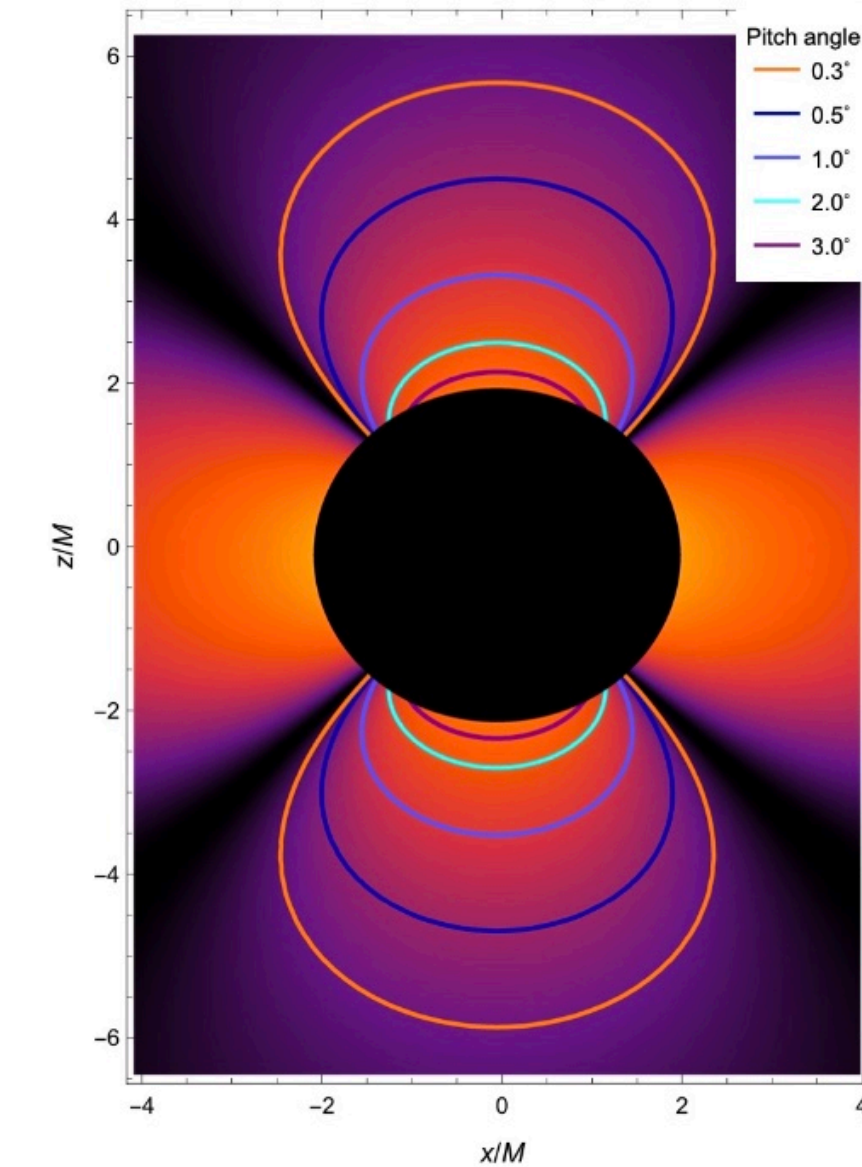
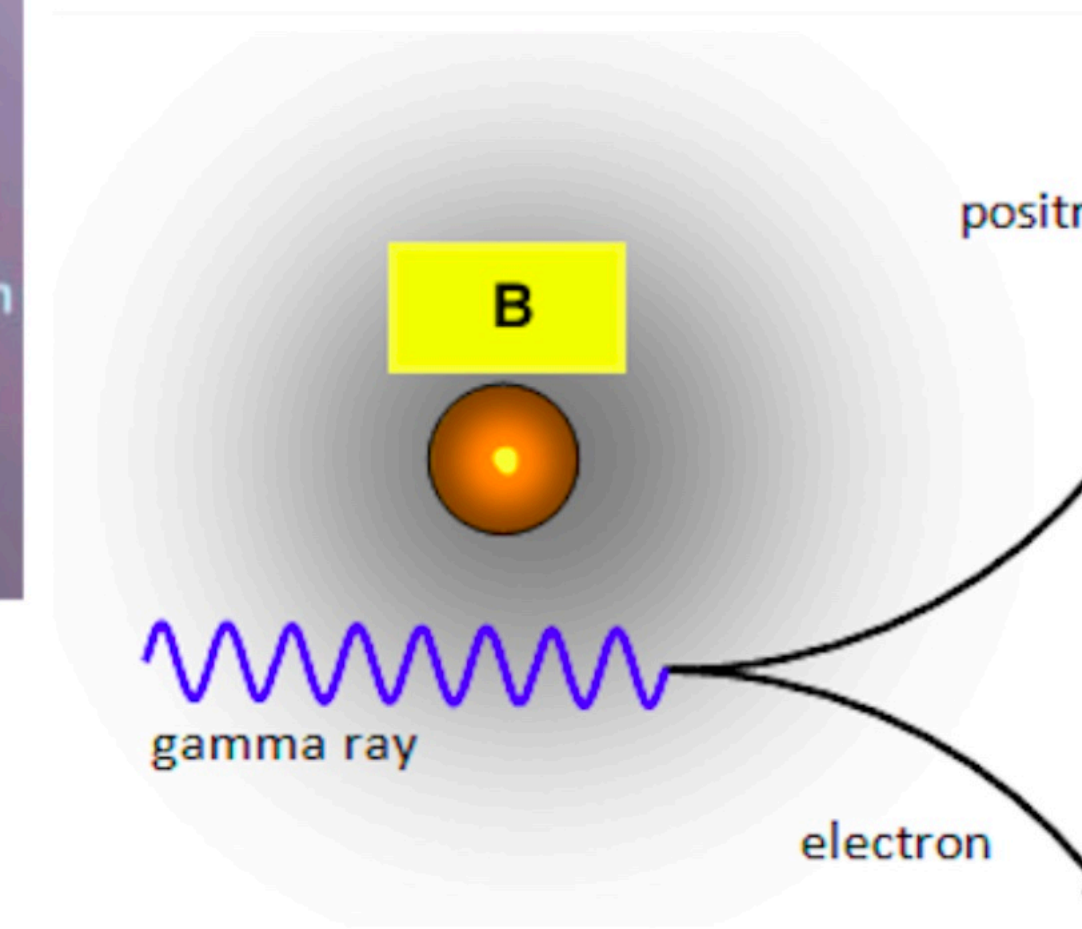


Image: Jon Lomberg/Gemini Observatory.



Mass, Spin and Magnetic Field

$$\begin{aligned}
 E_{\text{GeV}} &= E_{\text{extr}}(\mu, \alpha) \\
 \beta &= 3.67 \times 10^{-4} \alpha^{-1} \\
 t_c(\langle \theta \rangle, \alpha, \beta) &= \tau_{\text{ob},1}(\mu, \alpha, \beta, L_{\text{GeV}}).
 \end{aligned}$$

$$\mu = \left(1 - \sqrt{\frac{1 + \sqrt{1 - \alpha^2}}{2}} \right)^{-1} \frac{E_{\text{GeV}}}{M_{\odot} c^2}.$$

$$\begin{aligned}
 \beta &= \beta(\epsilon_{\gamma}, E_{\text{GeV}}, L_{\text{GeV}}, \alpha) \\
 &= \frac{1}{\alpha} \left(\frac{64}{9} \sqrt{3} \frac{e^2}{\hbar c} \frac{\epsilon_{\gamma}}{B_c^2 r_+(\mu, \alpha)^3} \frac{L_{\text{GeV}}}{e B_c c^2} \right)^{2/7},
 \end{aligned}$$

$$\beta \leq 3.67 \times 10^{-4} \alpha^{-1},$$

with $E_{\text{GeV}} = 1.8 \times 10^{53}$ erg, $L_{\text{GeV}} = 1.46 \times 10^{52}$ erg s⁻¹ and photon energy $\epsilon_{\gamma} = 0.1$ GeV.

For the given energy and luminosity,

- $\beta = 8.9 \times 10^{-4}$, i.e. $B_0 \approx 3.9 \times 10^{10}$ G.
- $\alpha = 0.414$
- $M = 4.447 M_{\odot}$
- $M_{\text{irr}} = 4.346 M_{\odot}$

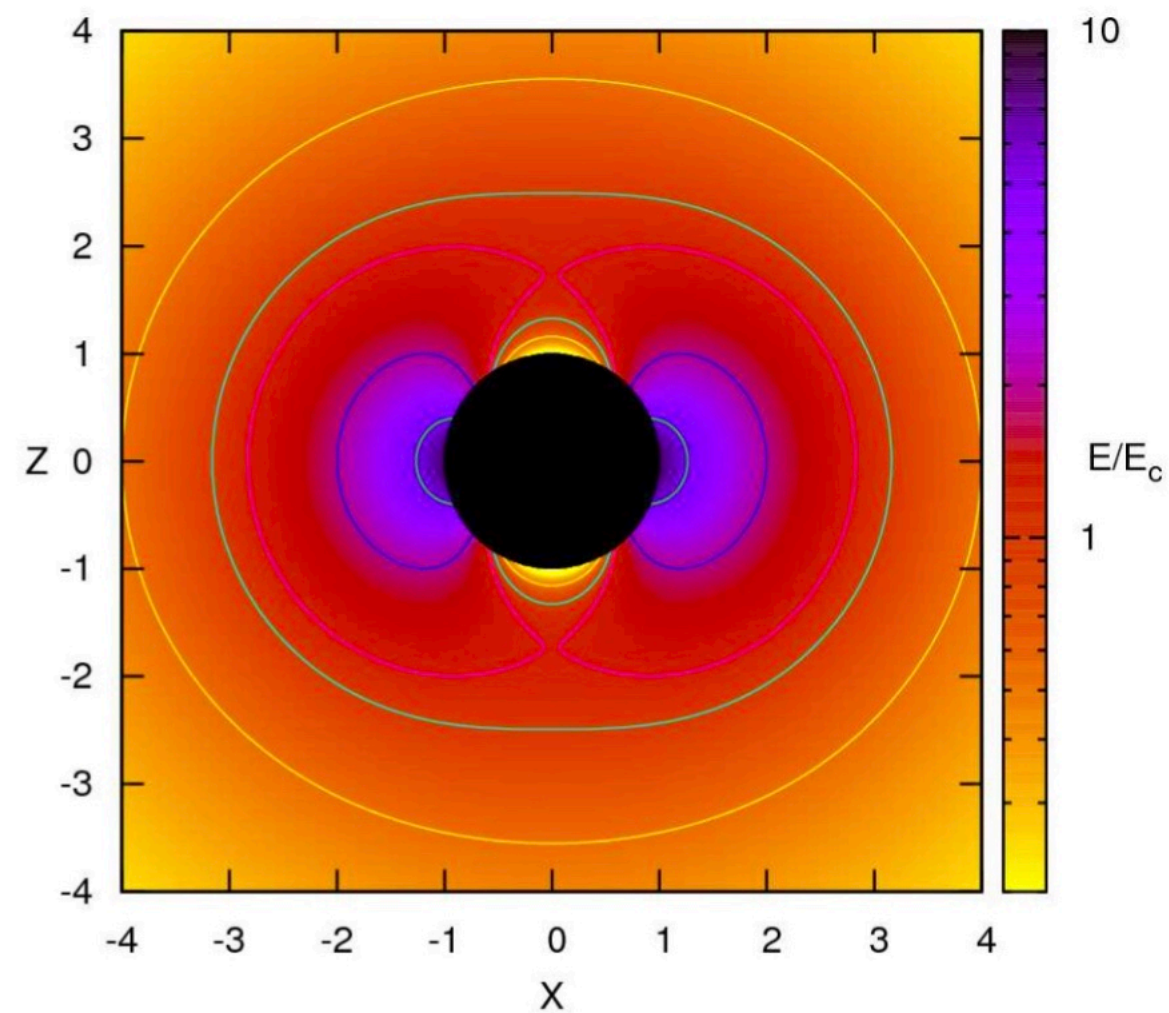
Synchrotron photons in the 0.1 GeV–1 TeV energy band, do not produce pairs if the magnetic field is below $B_0 < 3.9 \times 10^{10}$ G. Therefore, this region is transparent for such photons.

Over Critical E field regime and Ultrarelativistic prompt emission phase

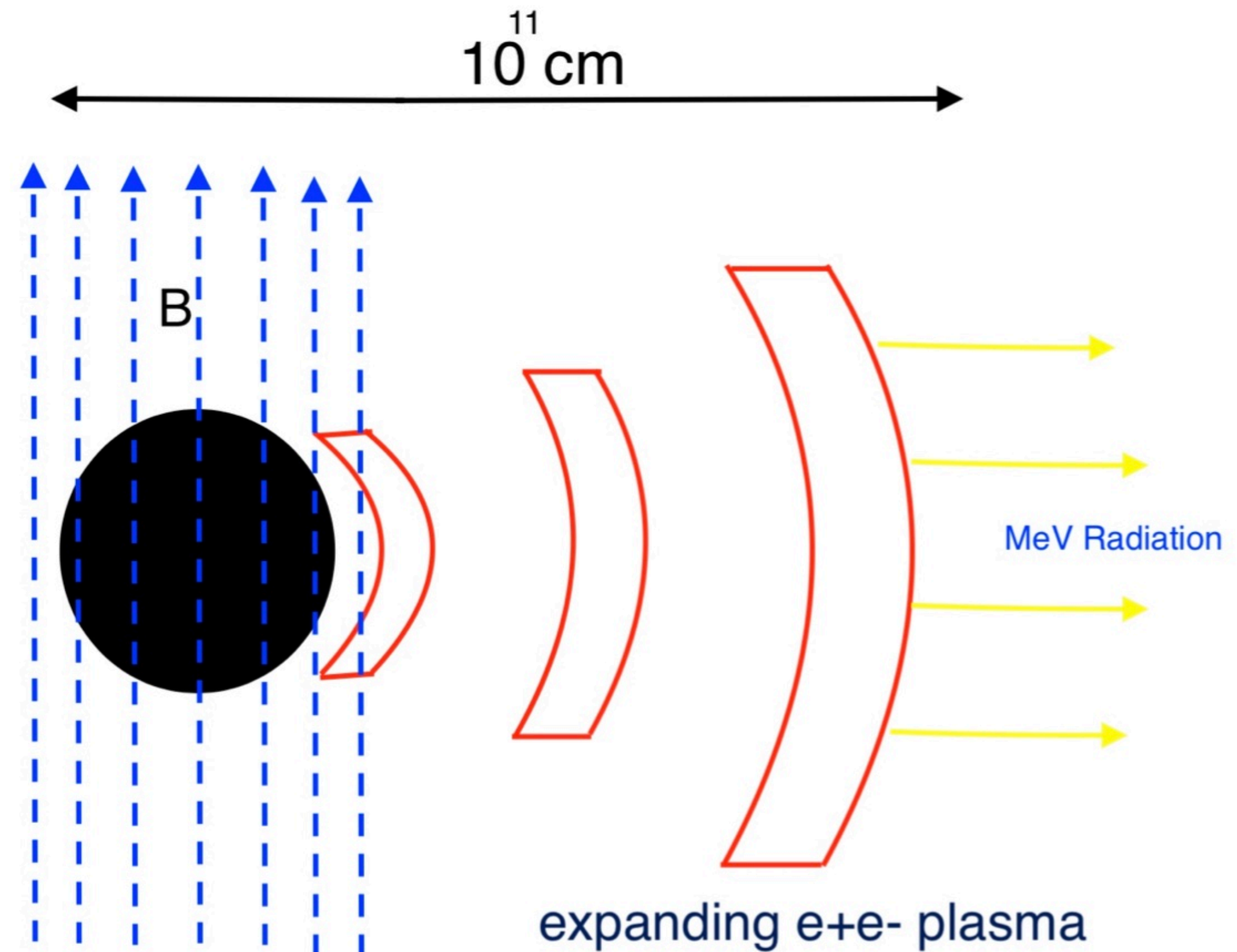
GRB 190114C: UPE mechanism and inner engine

Dyadoregion

C. Cherubini et al PRD 79, 124002 (2009)



$E_c = \frac{m_e^2 c^3}{e \hbar}$ is the critical value of vacuum polarization, where m_e and e are the electron mass and charge, respectively.



Damour and Ruffini, PRL,463 (1975)
 Ruffini, Salmonson, Wilson, and Xue, A&A350, 334 (1999)
 Ruffini, Salmonson, Wilson, and Xue, A&359, 855 (2000)

Nature of the ultrarelativistic prompt emission phase of GRB 190114C

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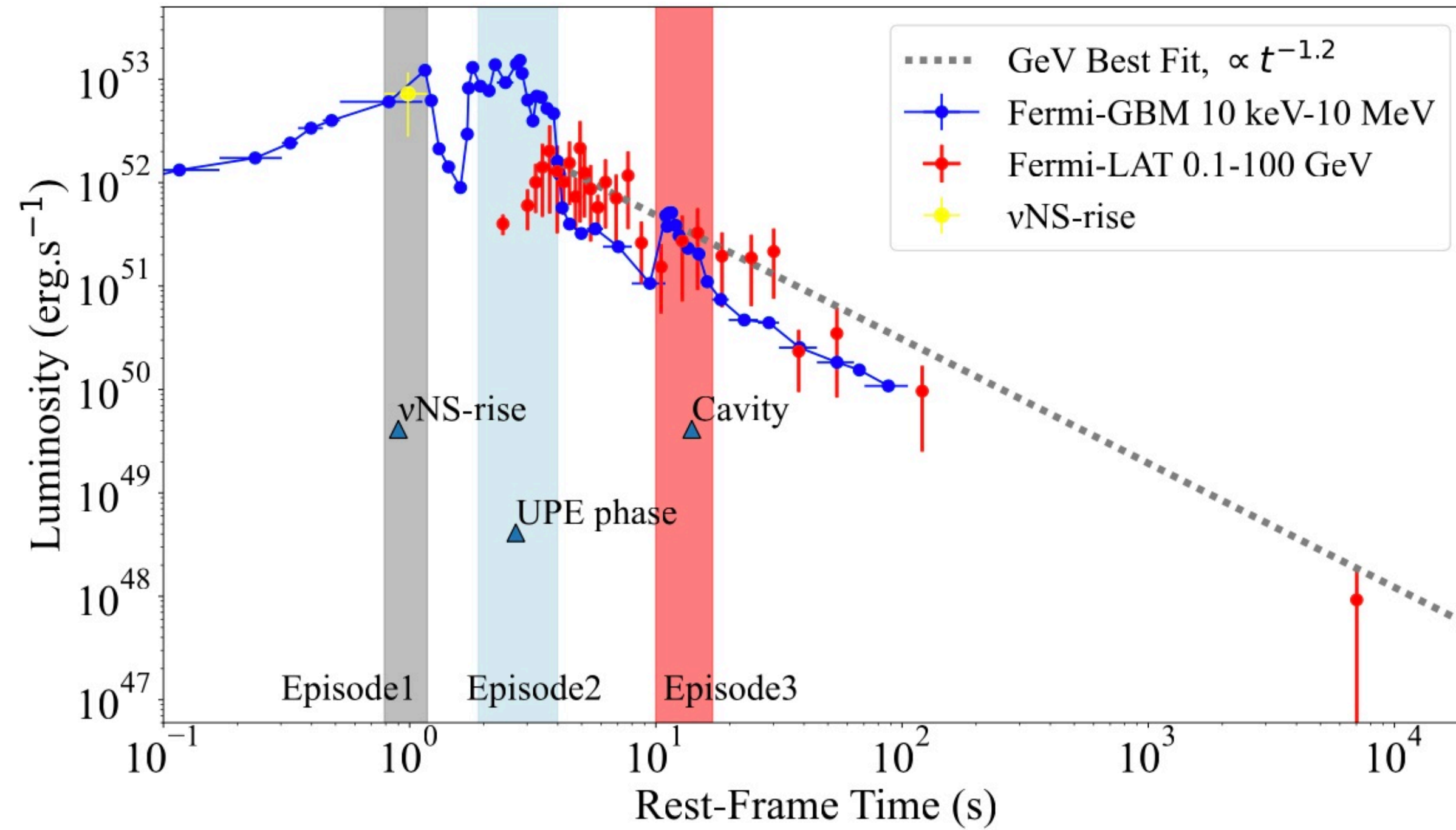


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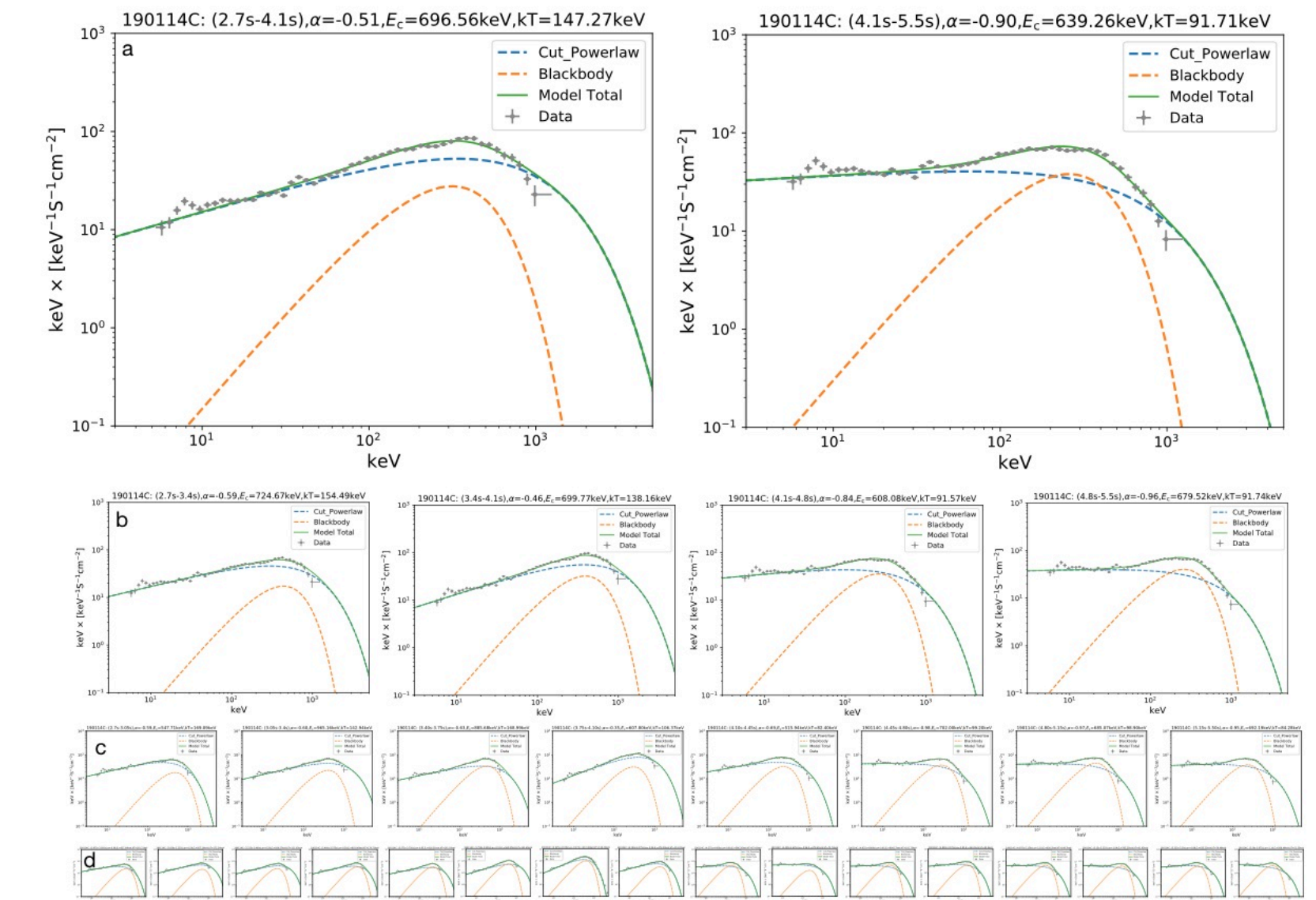
We address the physical origin of the ultrarelativistic prompt emission (UPE) phase of GRB 190114C observed in the interval $t_{\text{rf}} = 1.9\text{--}3.99$ s, by the *Fermi*-GBM in 10 keV–10 MeV energy band. Thanks to the high signal-to-noise ratio of *Fermi*-GBM data, a time-resolved spectral analysis has evidenced a sequence of similar blackbody plus cutoff power-law spectra (BB + CPL), on ever decreasing time

GRB 190114C: Ultra Relativistic Prompt Emission Phase (UPE)

Moradi, Rueda, Ruffini, Li, et al., PRD 104, 063043 (2021).



GRB 190114C: UPE phase



Time-resolved spectral analysis of UPE phase GRB 190114C.

GRB 190114C: UPE mechanism and inner engine

The electromagnetic field of the *inner engine* in the Carter's orthonormal tetrad is:

$$E_{\hat{r}} = \frac{\hat{a}B_0}{\Sigma} \left[r \sin^2 \theta - \frac{\hat{M} (\cos^2 \theta + 1) (r^2 - \hat{a}^2 \cos^2 \theta)}{\Sigma} \right], \quad (1)$$

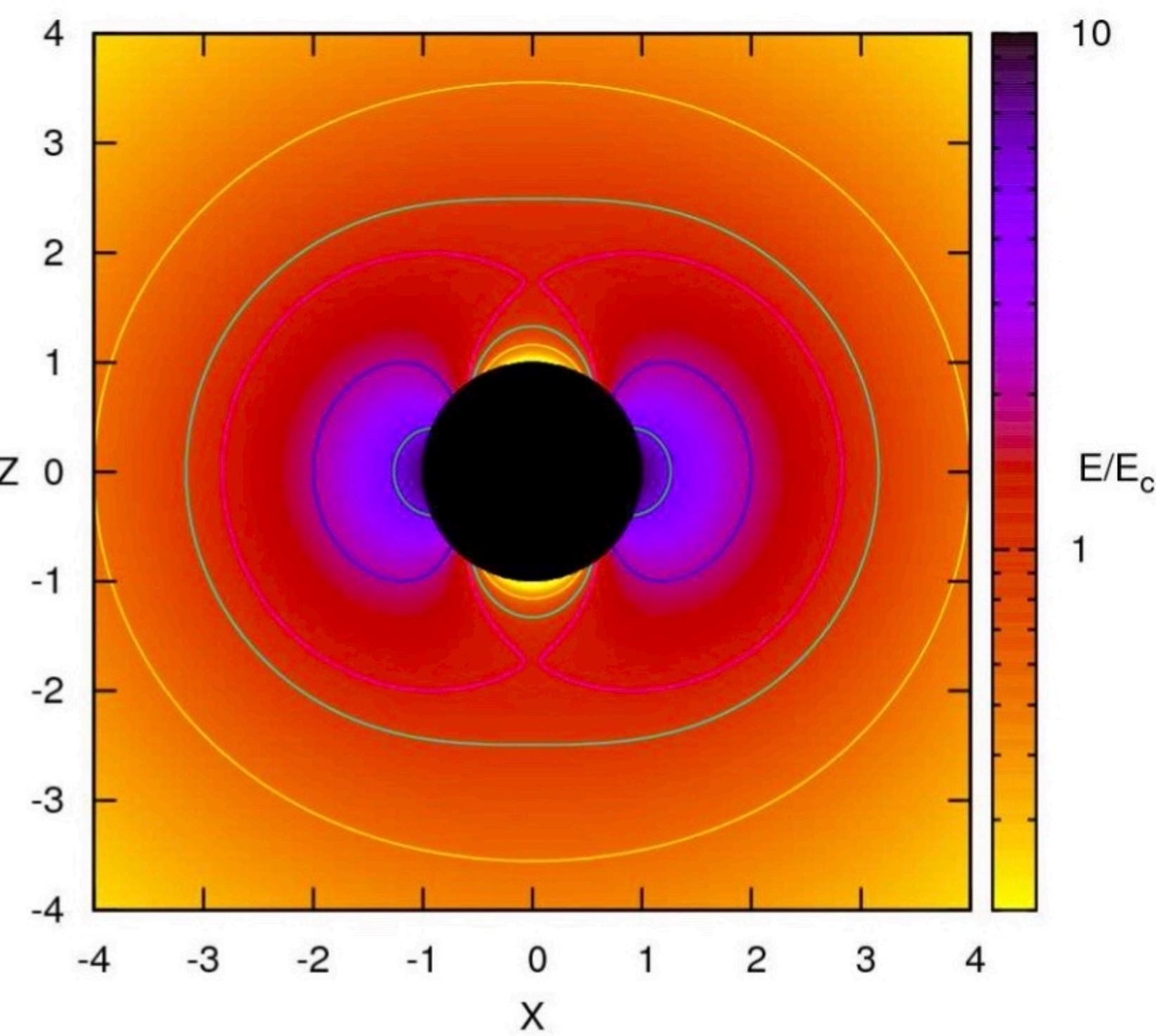
$$E_{\hat{\theta}} = \frac{\hat{a}B_0}{\Sigma} \sin \theta \cos \theta \sqrt{\Delta}, \quad (2)$$

$$B_{\hat{r}} = -\frac{B_0 \cos \theta}{\Sigma} \left[-\frac{2\hat{a}^2 \hat{M} r (\cos^2 \theta + 1)}{\Sigma} + \hat{a}^2 + r^2 \right], \quad (3)$$

$$B_{\hat{\theta}} = \frac{B_0 r}{\Sigma} \sin \theta \sqrt{\Delta}, \quad (4)$$

where $\Sigma = r^2 + \hat{a}^2 \cos^2 \theta$, $\Delta = r^2 - 2\hat{M}r + \hat{a}^2$, $\hat{M} = GM/c^2$, $\hat{a} = a/c = J/(Mc)$, being M and J the mass and angular momentum of the Kerr BH. The (outer) event horizon is located at $r_+ = (\hat{M} + \sqrt{\hat{M}^2 - \hat{a}^2})$.

Analogy with Kerr-Newmann



Dyadoregion

The dyadoregion energy:

$$E_{(r_+, r_d)} = \frac{(2B_0 JG/c^3)^2}{4r_+} \left(1 - \frac{r_+}{r_d}\right) + \frac{(2B_0 JG/c^3)^2}{4\hat{a}} \times \left[\left(1 + \frac{\hat{a}^2}{r_+^2}\right) \arctan\left(\frac{\hat{a}}{r_+}\right) - \left(1 + \frac{\hat{a}^2}{r_d^2}\right) \arctan\left(\frac{\hat{a}}{r_d}\right) \right], \quad (5)$$

where r_d is the radius of the dyadoregion

$$\left(\frac{r_d}{\hat{M}}\right)^2 = \frac{1}{2} \frac{\lambda}{\mu\epsilon} - \alpha^2 + \left(\frac{1}{4} \frac{\lambda^2}{\mu^2\epsilon^2} - 2 \frac{\lambda}{\mu\epsilon} \alpha^2\right)^{1/2} \quad (6)$$

with $\epsilon = E_c M_\odot G^{3/2}/c^4 \approx 1.873 \times 10^{-6}$, and

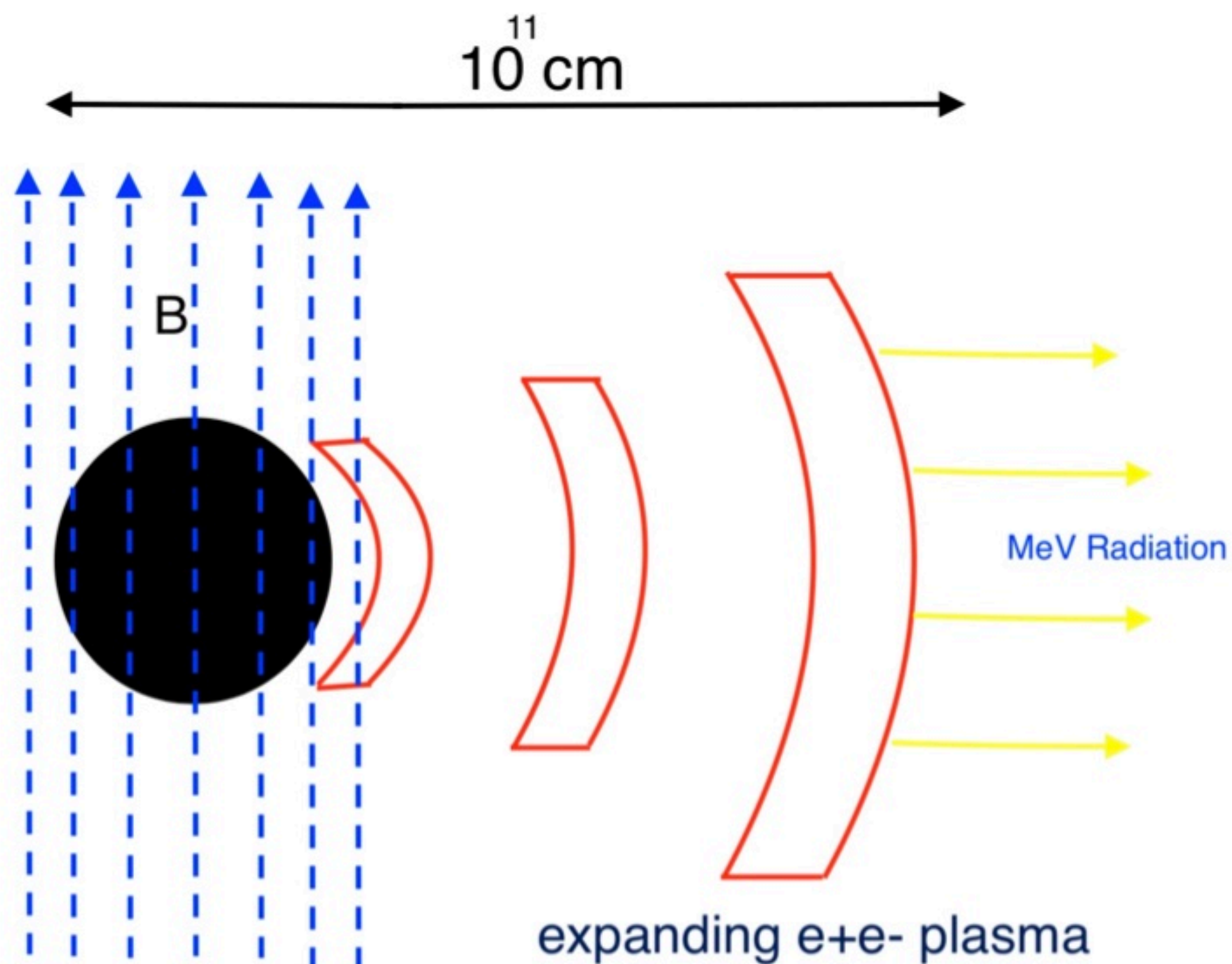
$$\lambda = (2B_0 JG/c^3)/(\sqrt{GM}), \quad (7)$$

is the effective charge-to-mass ratio.

The characteristic width of the dyadoregion, which demonstrates the region in which the electric field around the BH is overcritical is

$$\Delta_d(t) = r_d(t) - r_+(t). \quad (8)$$

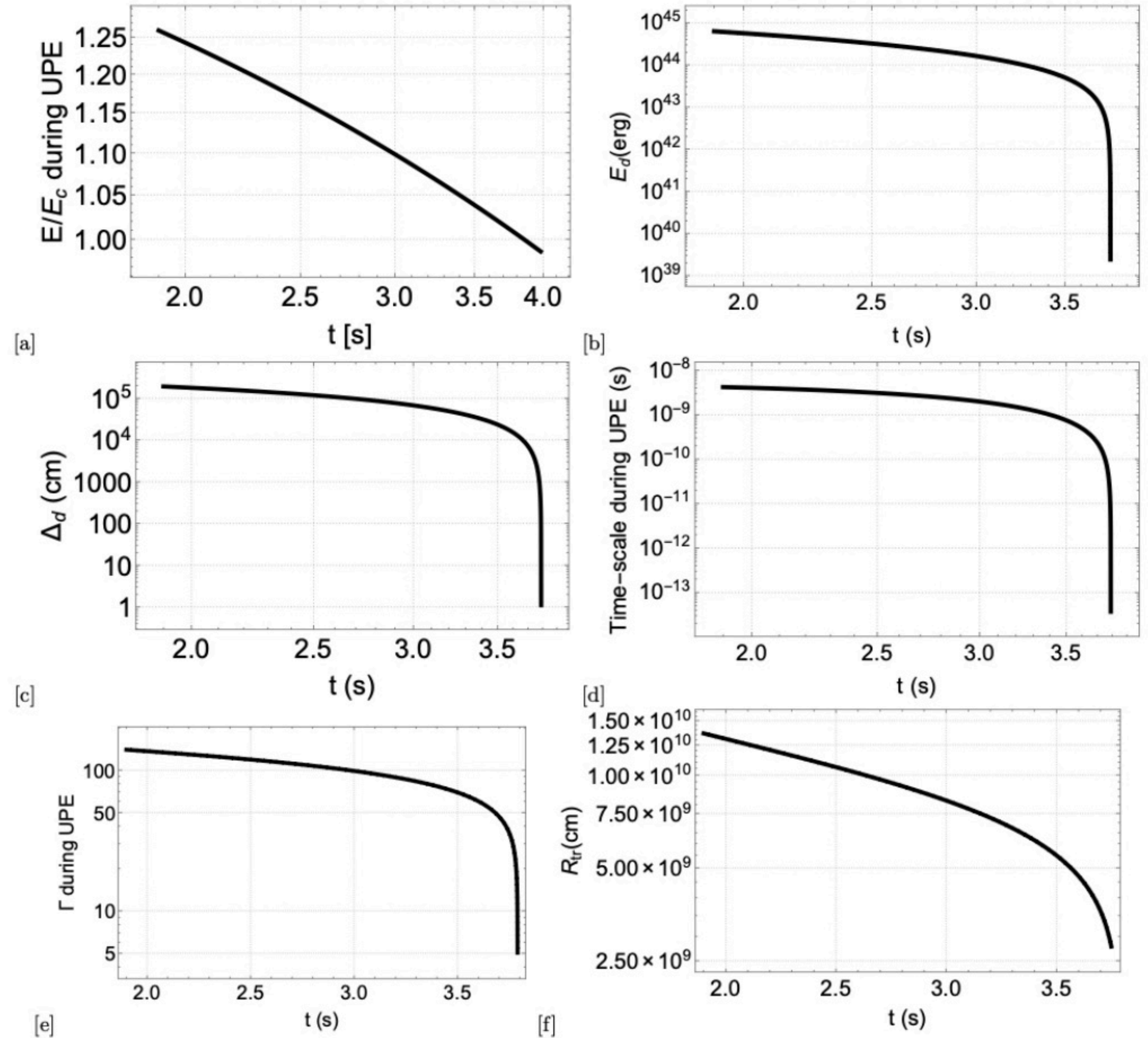
key parameters for calculating the transparency radius of each impulse are:



- its isotropic energy, E_{iso} ,
- its ratio of black body energy to isotropic energy ($E_{\text{P-GRB}}^{\text{obs}}/E_{\text{iso}}$)
- its value of black body temperature in keV, (T_{obs})
- its width $\Delta_{\text{lab}} = \Delta_{\text{d}}$.

Transparency obtained from the lower limit of magnetic field,
 $B_0 = 2.3 \times 10^{14}$ G ($|\mathbf{E}| = E_c$ at the end of UPE phase).

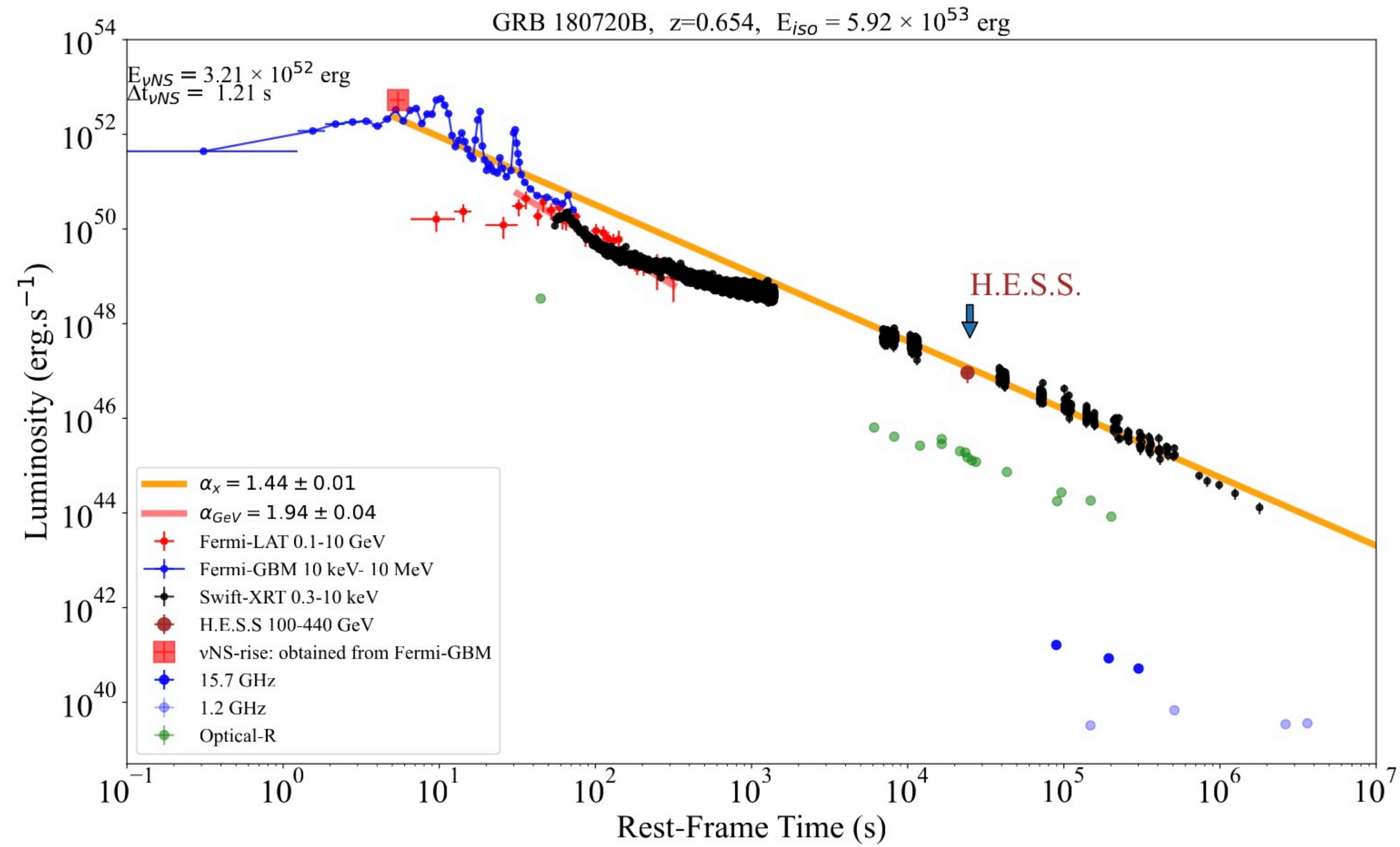
**Not Continuous Emission
But A discrete emission
with time scale of 10^{-9} s**

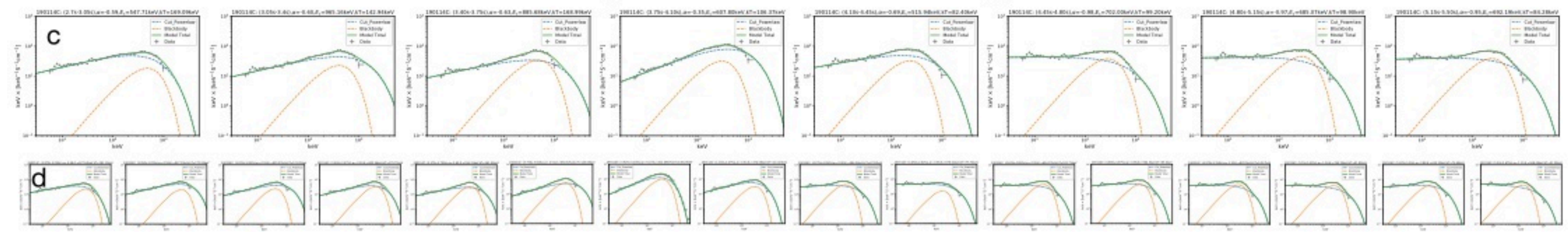
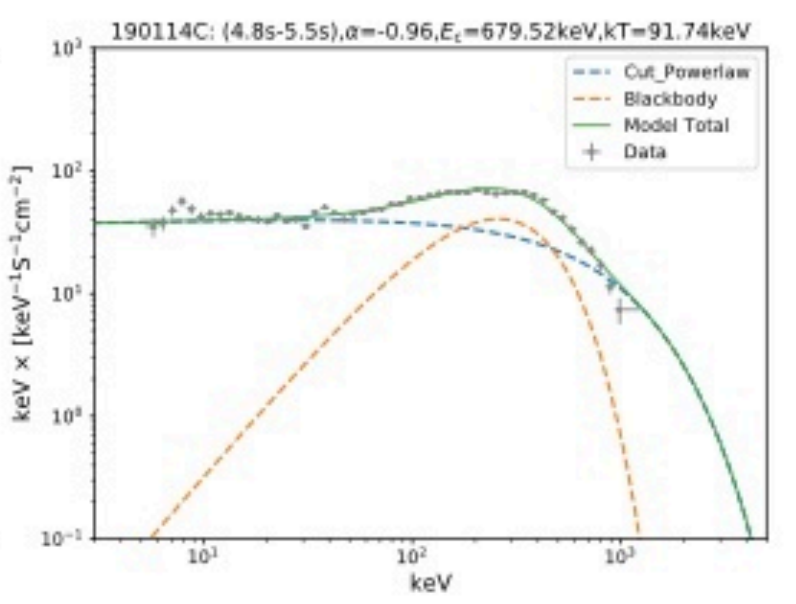
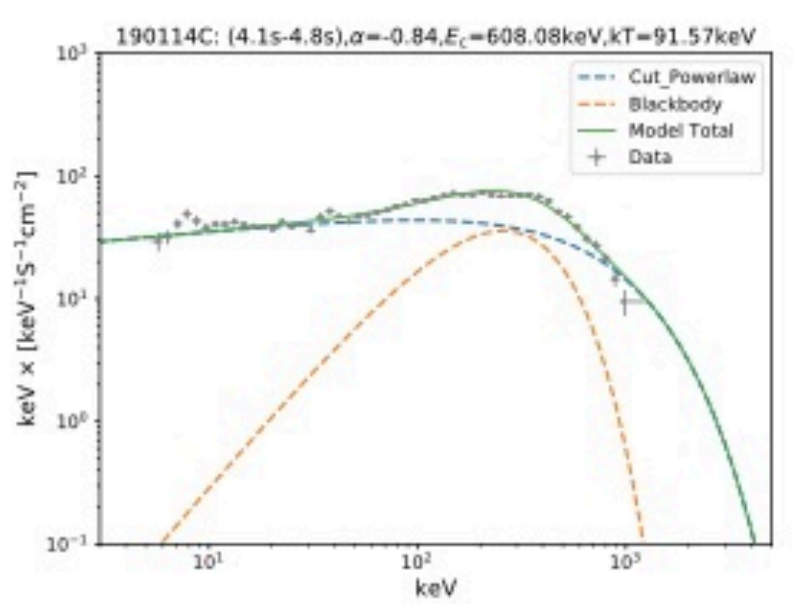
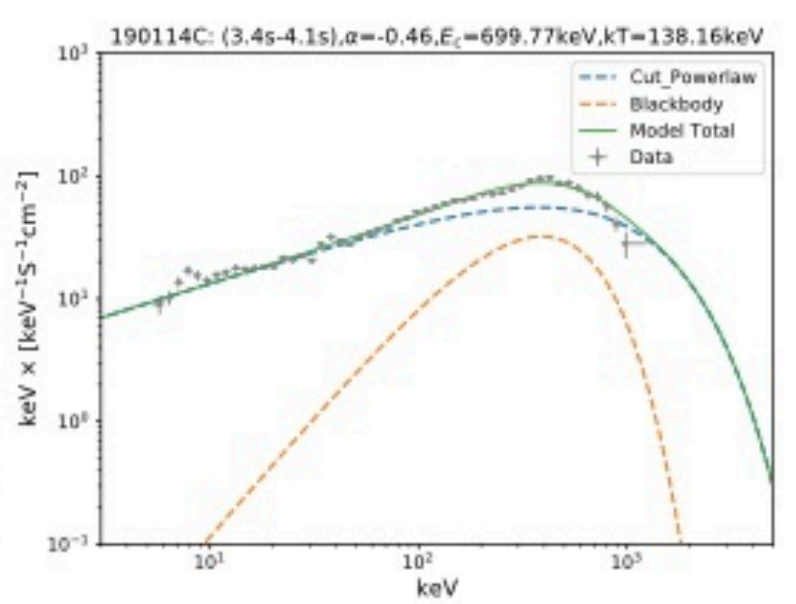
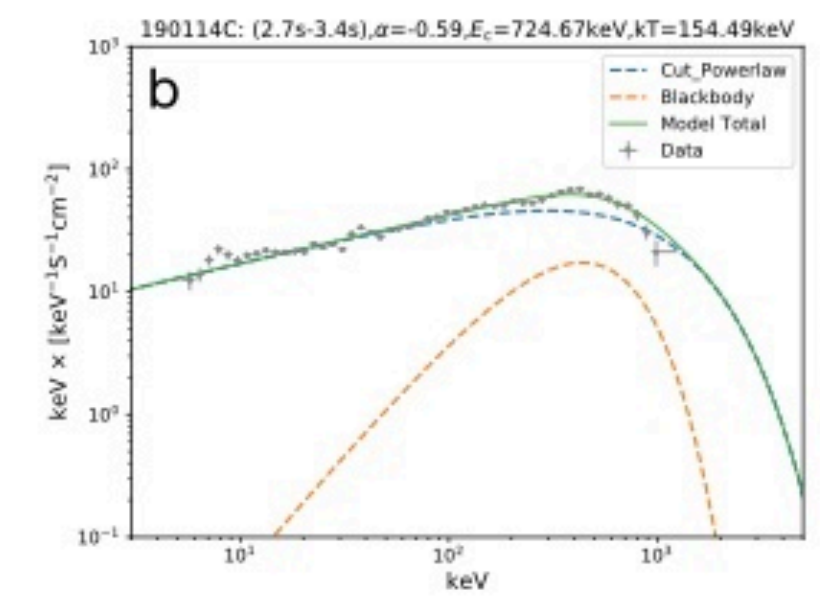
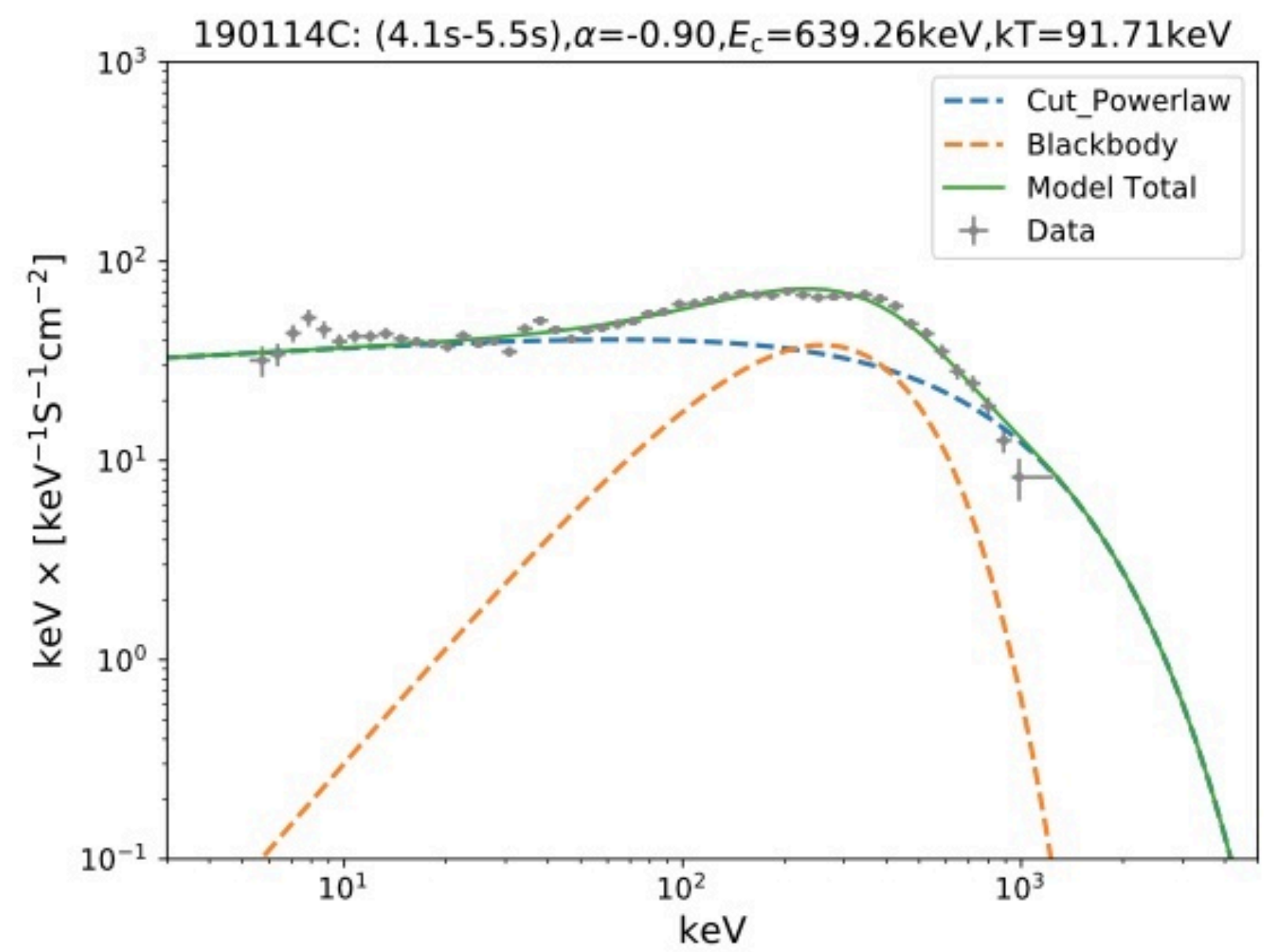
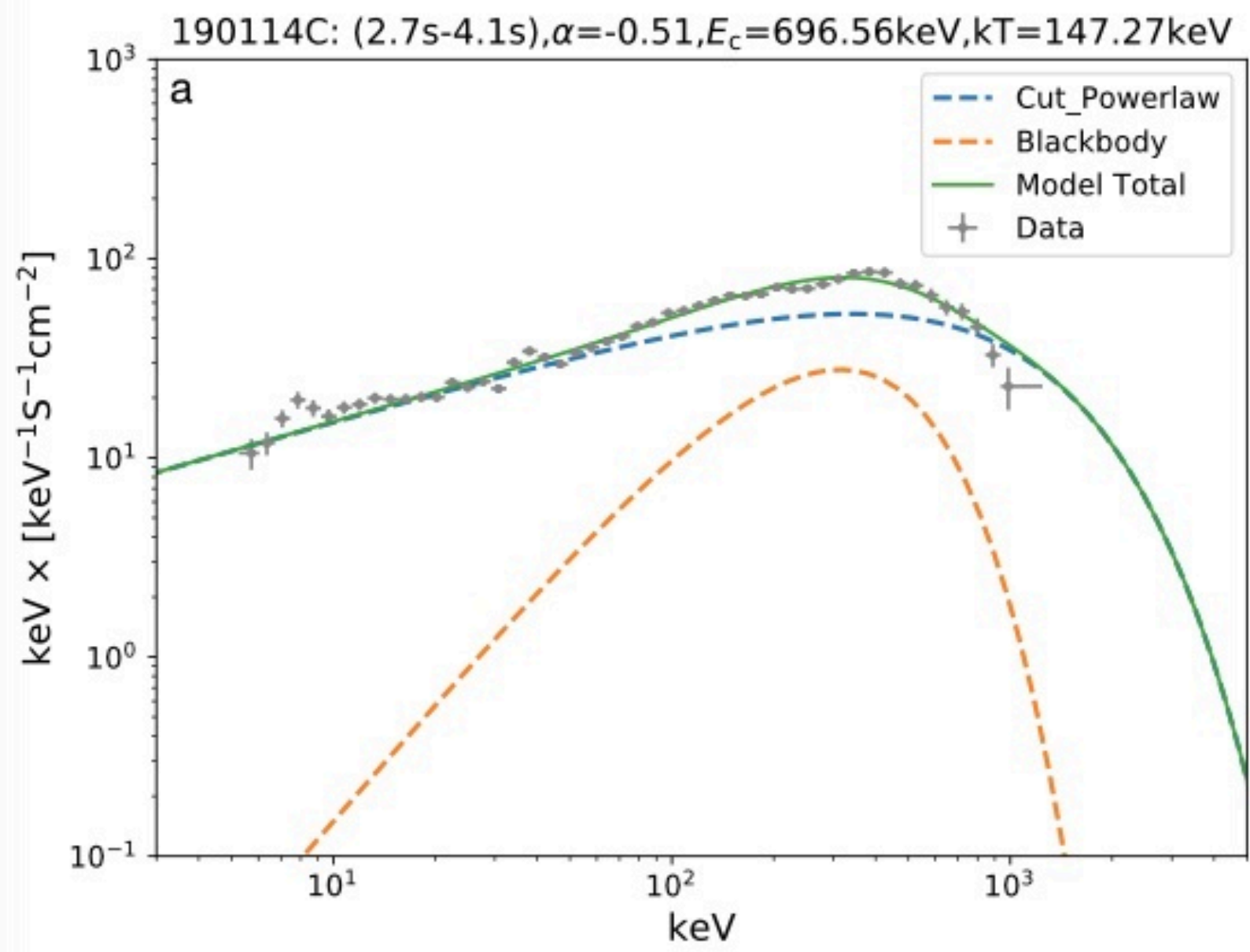


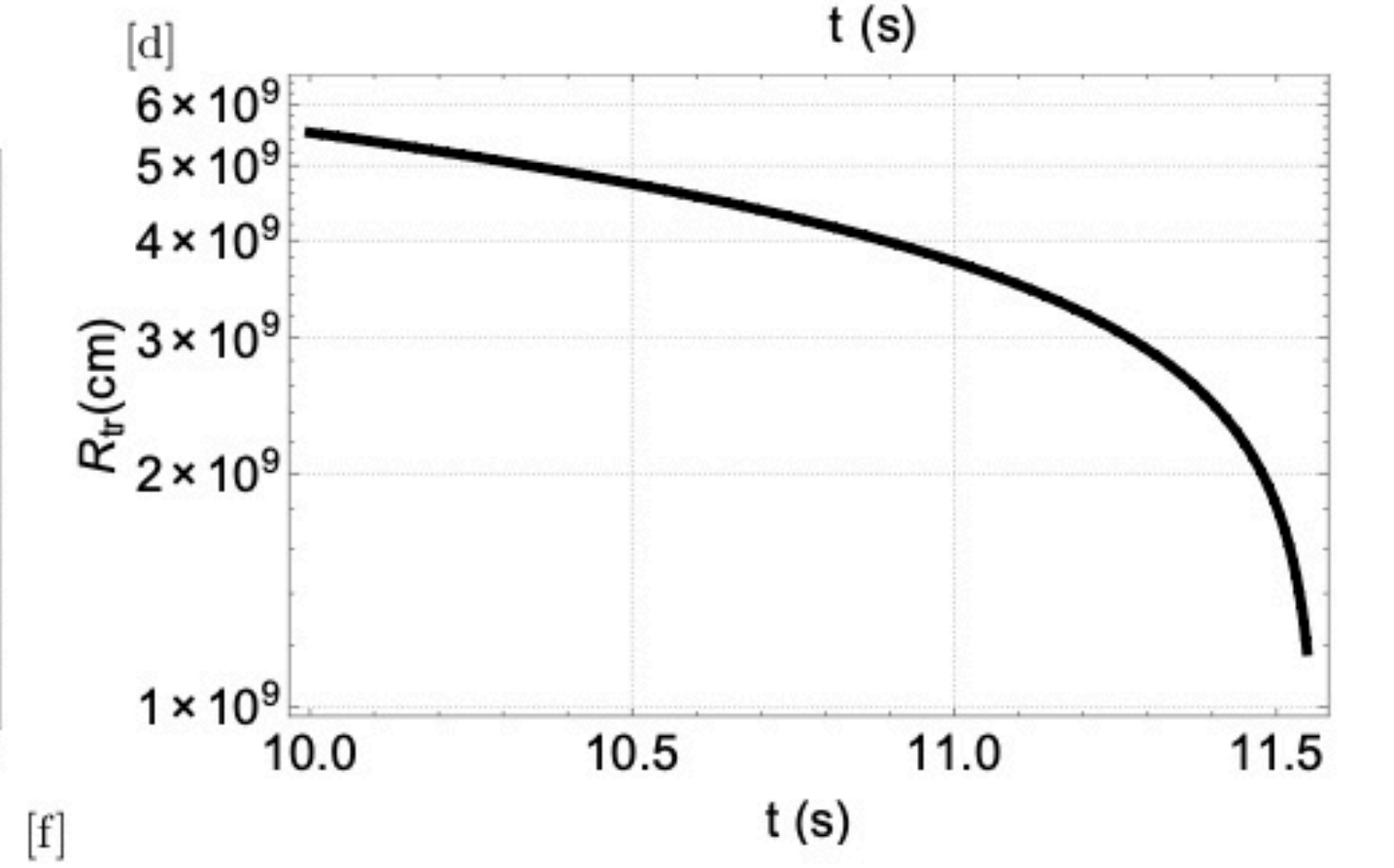
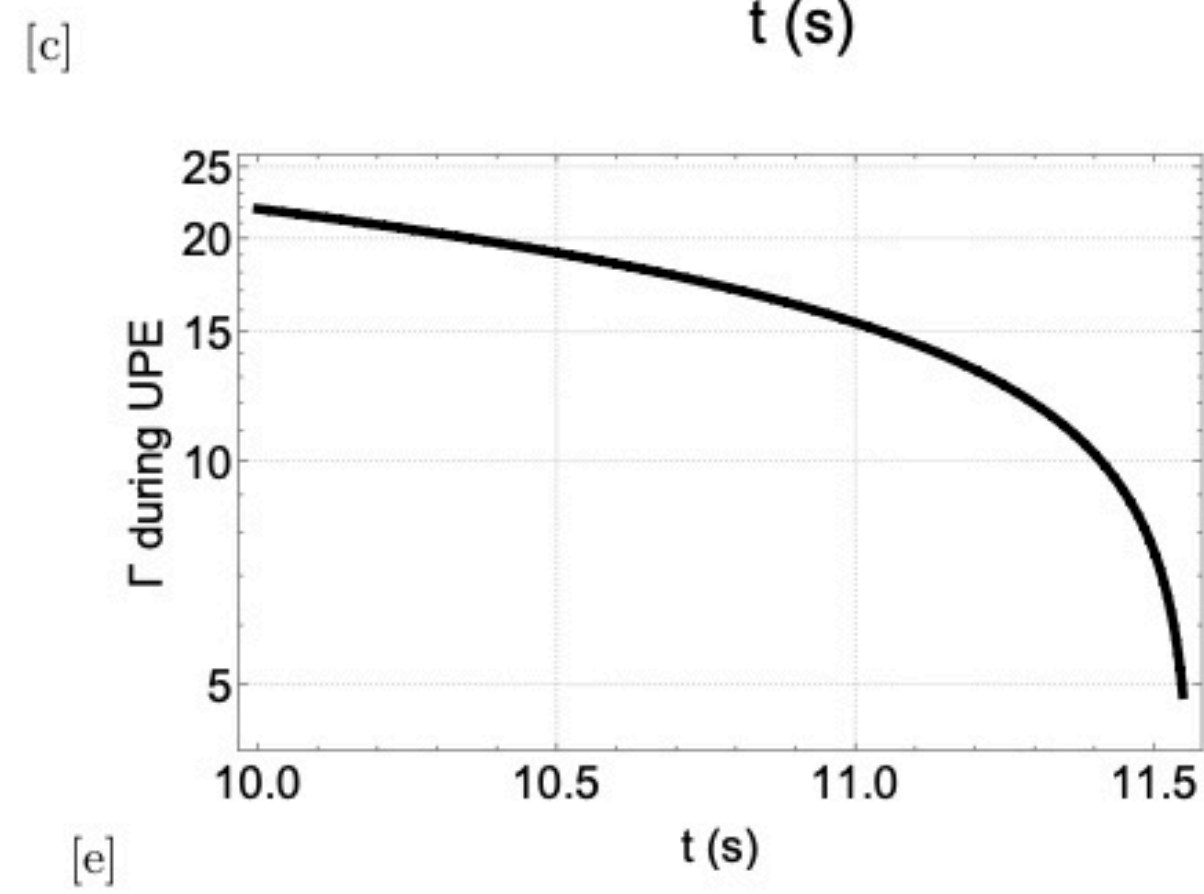
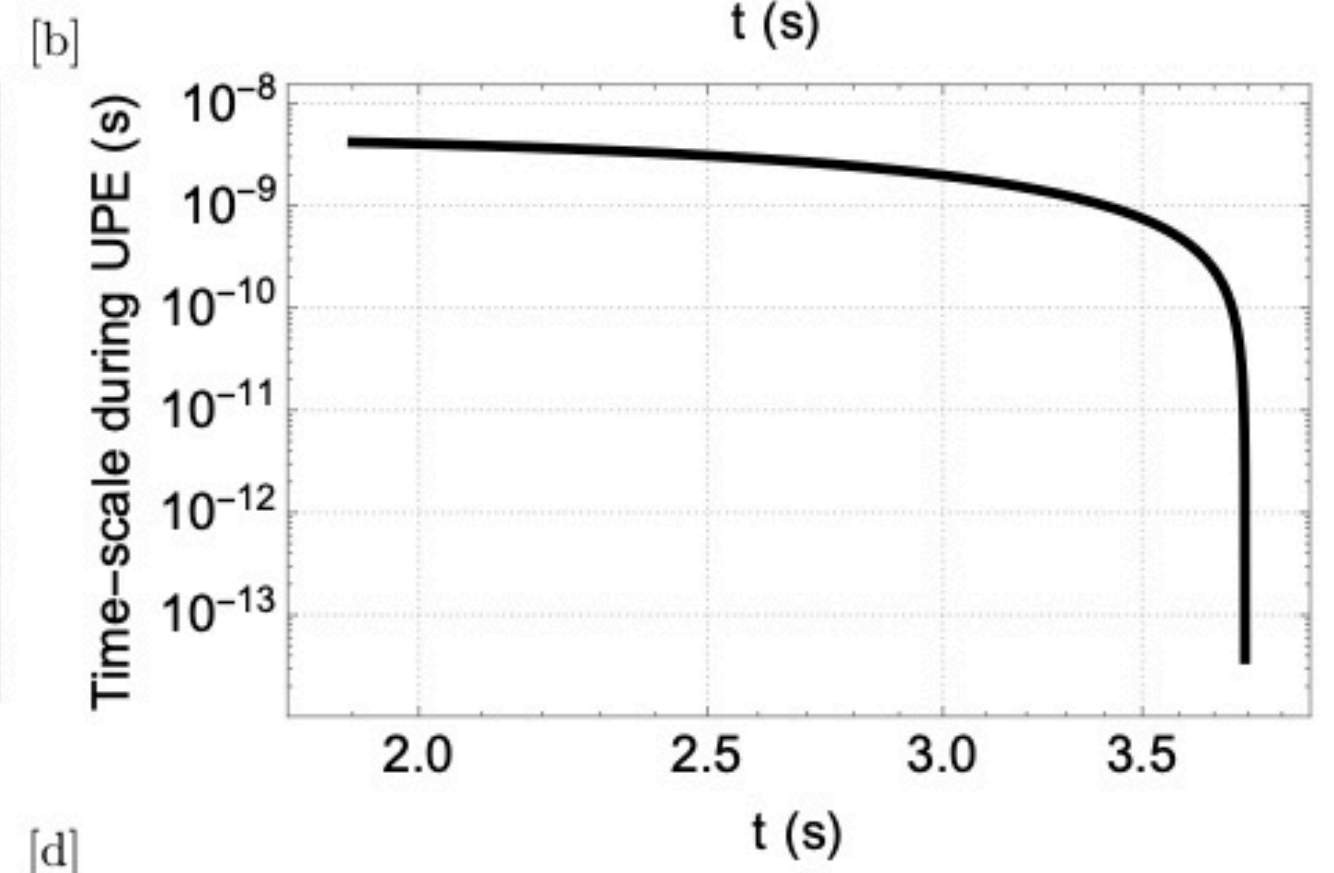
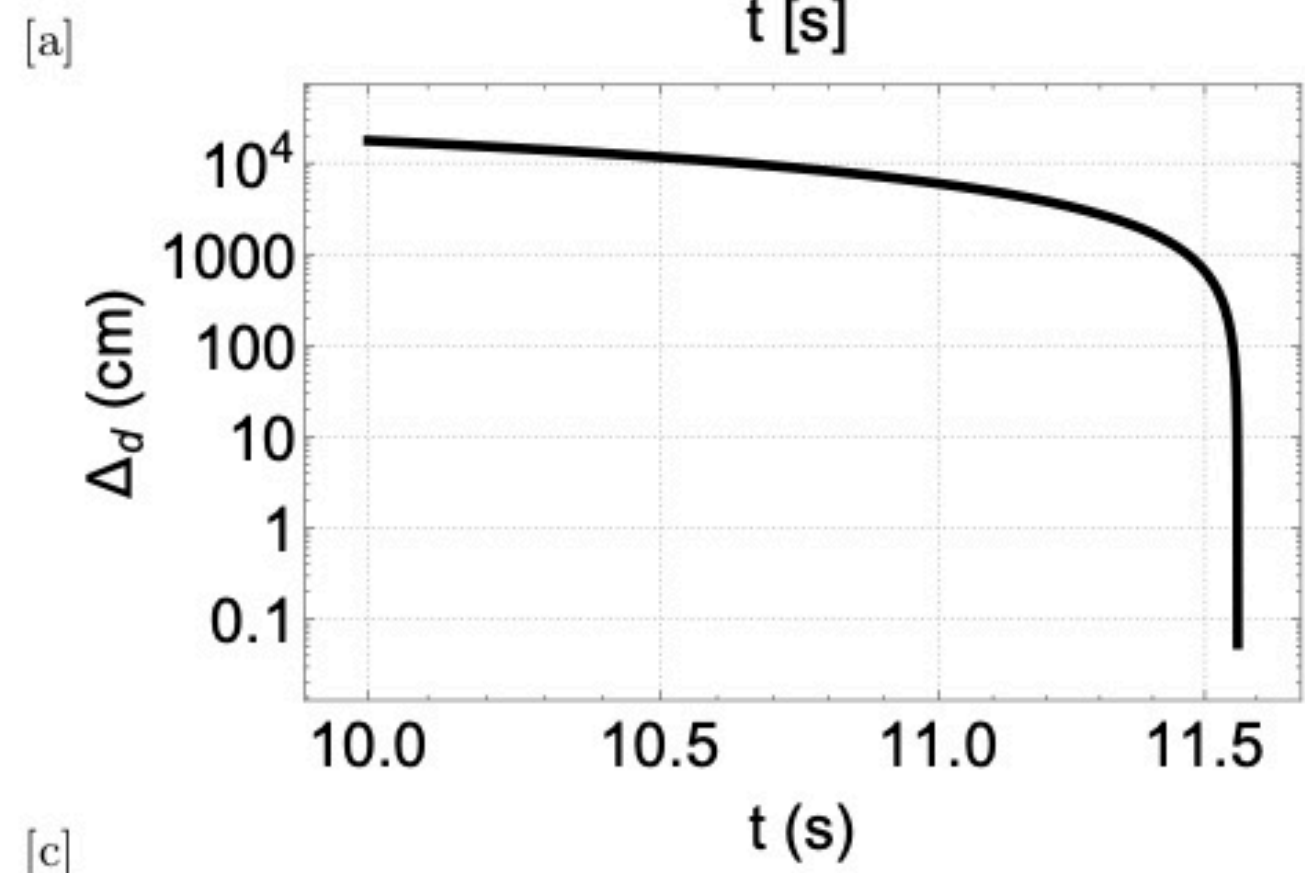
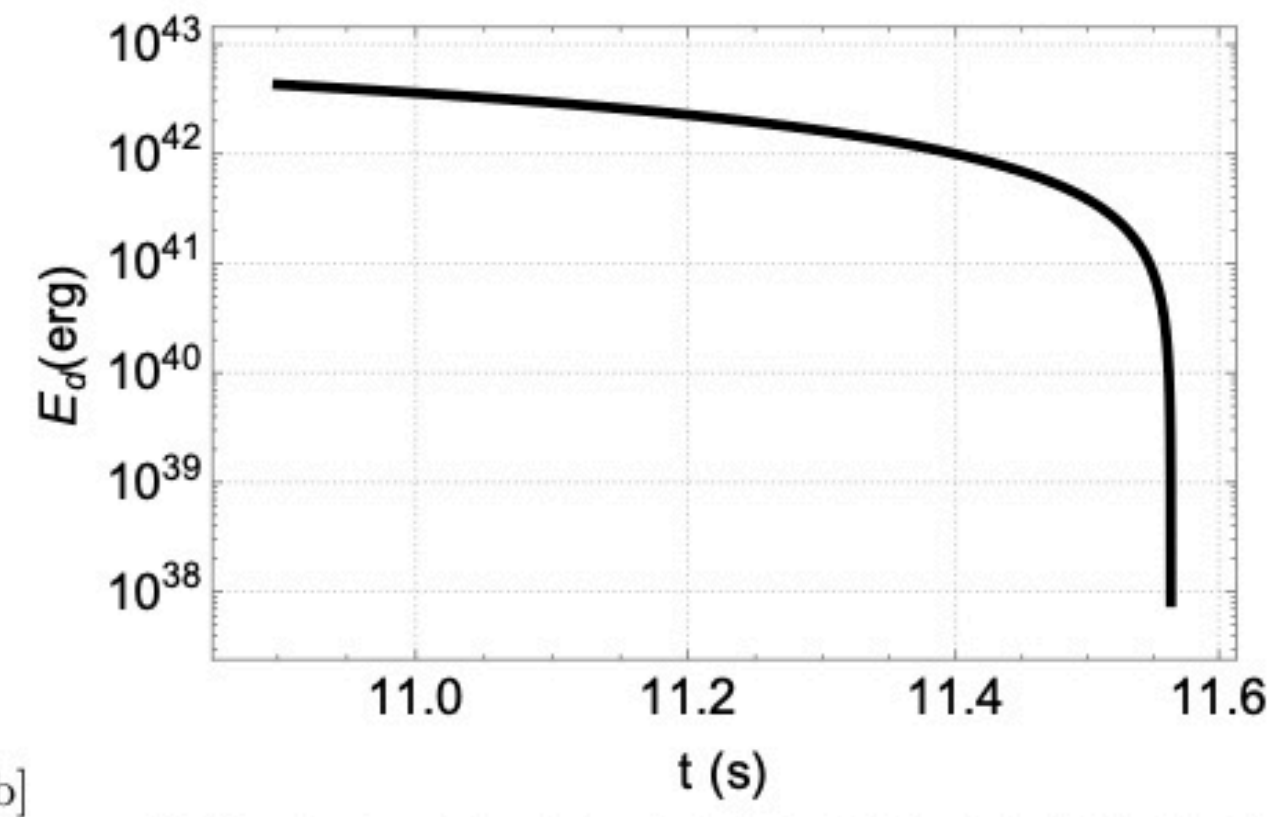
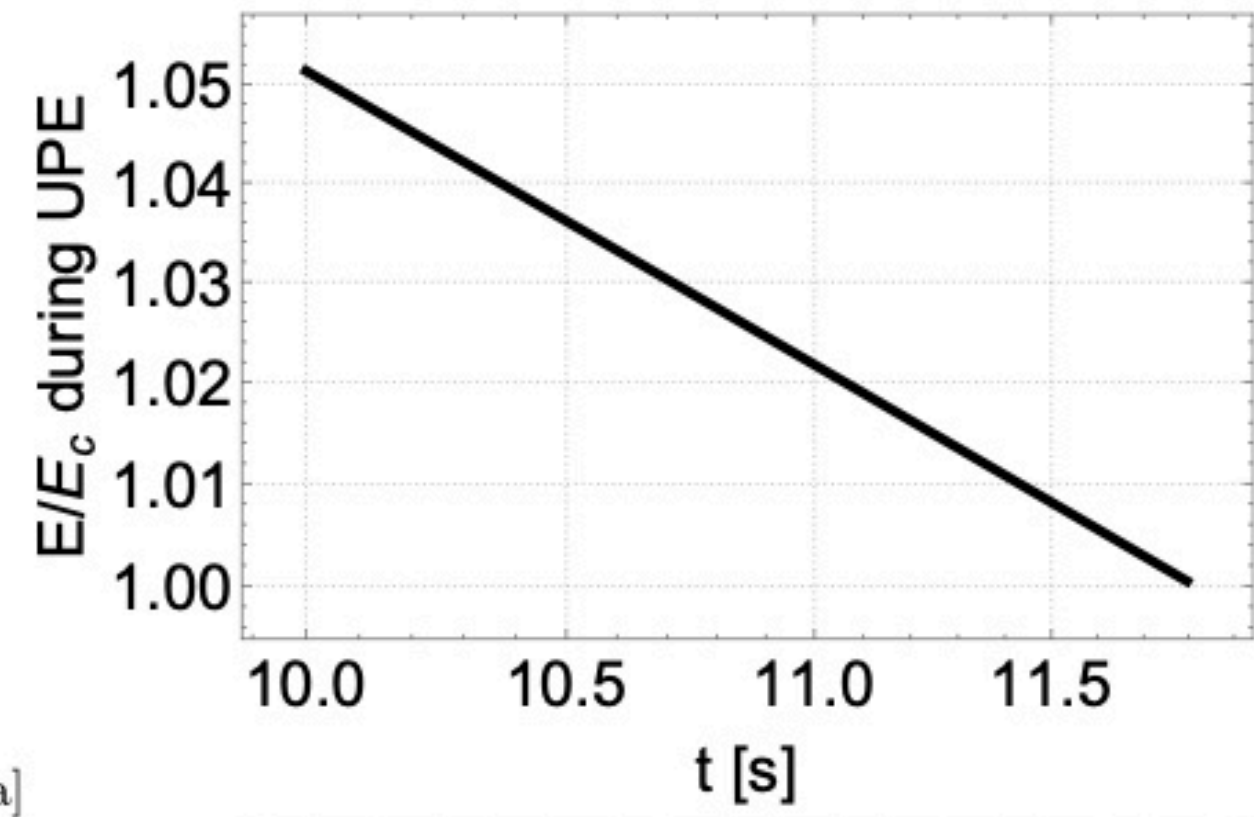
GRB 180720B

Rastegarnia et al in prepsration

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What are we learning on BHs from GRB observations?

The usual assumptions adopted for mathematical convenience in the description of the Kerr BH are:

1. The condition of the stationary space-time.
2. The condition of the vacuum space-time.
3. The condition of asymptotic flatness.

From the observation of GRBs (see e.g. GRB 1807820B and GRB 190114C) we conclude that Kerr BHs in these systems are:

1. are endowed with a magnetic field, B_0 parallel to their rotation axis;
2. They have no electric charge and the gravitomagnetic interaction between the BH angular momentum (J) and the surrounding magnetic field (B_0) induces an electric field which can be viewed as produced by an effective charge $Q_{\text{eff}} = 2J B_0$ (Papapetrou-Wald solution);
3. They are embedded in a low density fully ionized plasma, essential for the electrodynamics of the energy extraction process (Ruffini, et al., 2019, ApJ 883, 191).

What are we learning on BHs from GRB observations?

Thanks to “the extractable energy” in the Christodoulou-Ruffini mass-energy formula, the Kerr BHs (Ruffini, et al., 2019, ApJ 886, 82):

1. radiate in principle for an infinite time; they follow a power-law luminosity in the source rest-frame $L_{\text{GeV}} = A_{\text{GeV}} t^{-\alpha_{\text{GeV}}}$ (Ruffini, et al., 2021, MNRAS 504, 5301);
2. radiate GeV photons in a conical region of 60° around the BH rotation axis;
3. they do not emit continuously but in blackholic quanta (Rueda & Ruffini, 2020, EPJC 80, 300);

All the above applies not only to GRBs but also to AGN scaling the BH from $\sim 10M_\odot$ to 10^9M_\odot .