

17TH MARCEL GROSSMANN MEETING 2024



DETECT SUPERMASSIVE DARK STARS WITH ROMAN SPACE TELESCOPE

Keywords: detectors, early universe, dark matter, first stars

PRESENTER: SAIYANG ZHANG

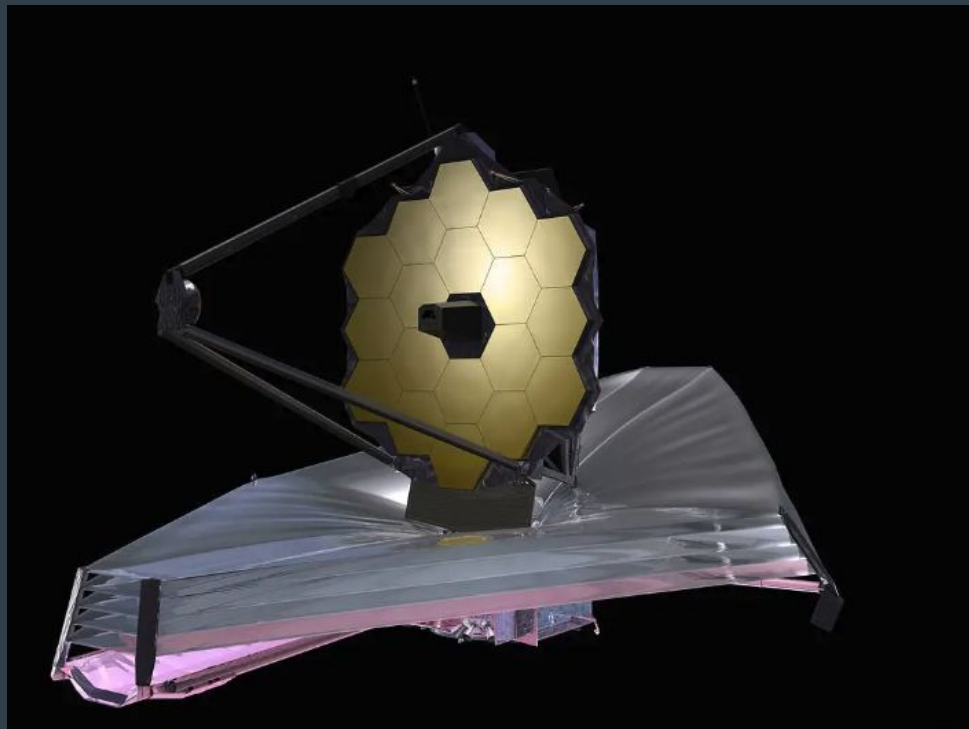
Department of Physics, University of Texas at Austin

COAUTHOR:

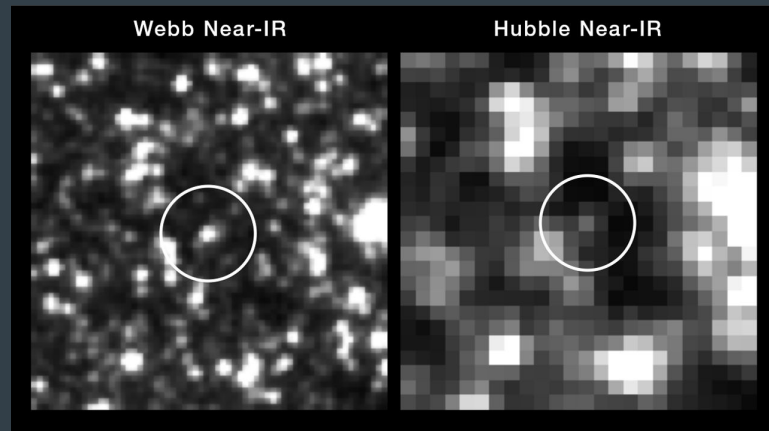
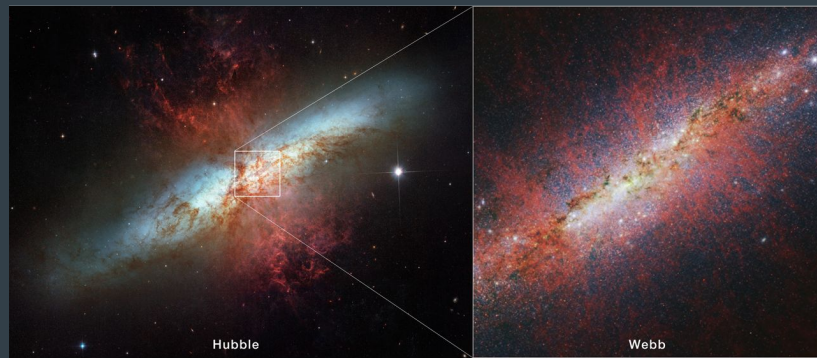
Cosmin Ilie, Department of Physics and Astronomy, Colgate University

Katherine Freese, Department of Physics, University of Texas at Austin

James Webb Space Telescope (JWST)



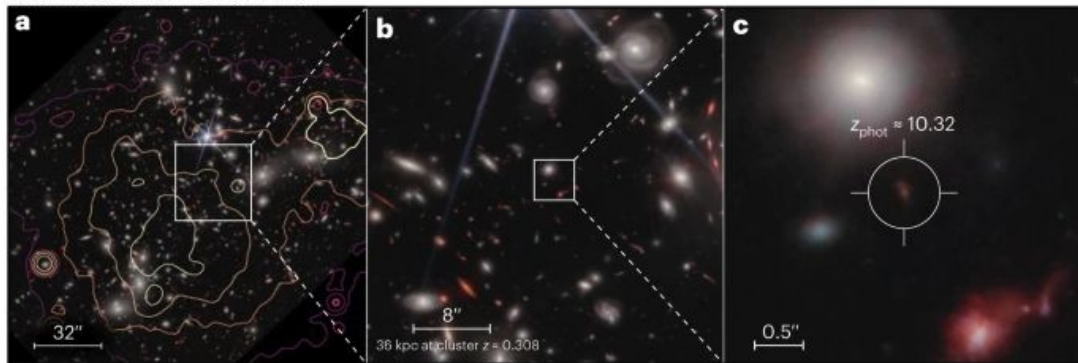
<https://www.nasa.gov/image-article/james-webb-space-telescope-jwst/>



Some Recent Discoveries by JWST

Fig. 2: JWST and Chandra images of UHZ1.

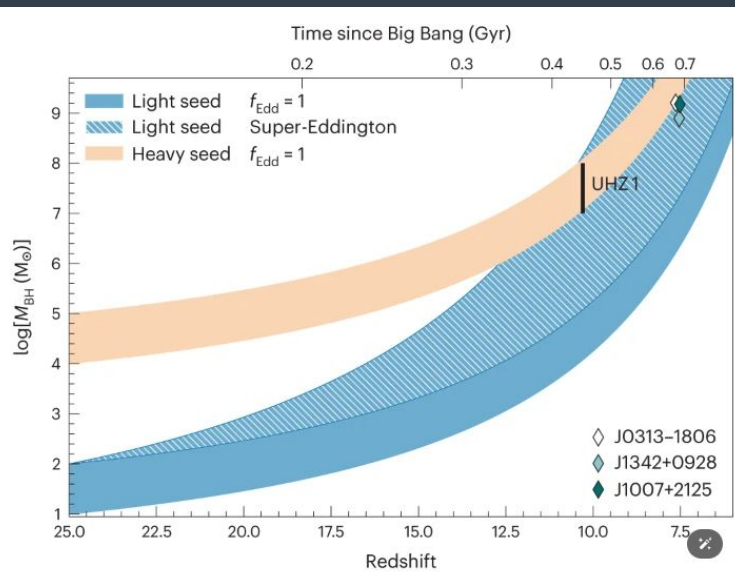
JWST NIRCam zoom-in on UHZ1



JWST NIRCam UHZ1 images



$z=10$ AGN: UHZ-1

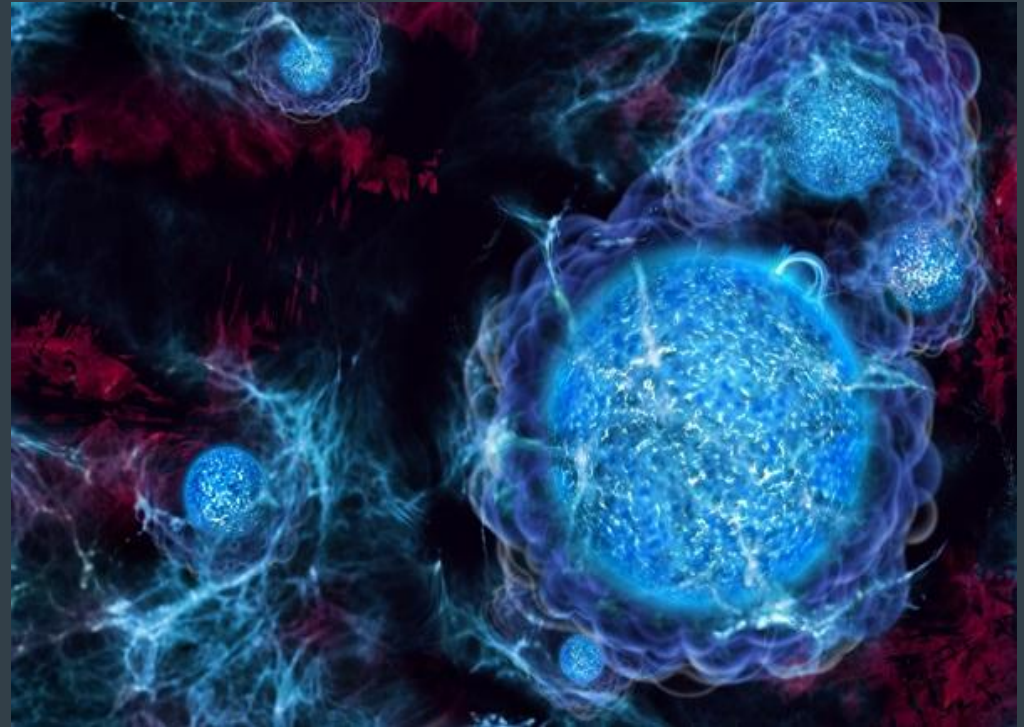


Bodgan A., et al. (2024)

What are the Dark Stars?*

Freese, K. et al. 2008; Freese, K. et al. 2016

- Formed in the early universe ($z \sim 10-50$)
- Powered by dark matter (WIMP) annihilation
- Giant, cool and puffy stellar object
- Two formation scenario: Adiabatic Contraction (AC) and Capture of DM
- Accrete mass from surrounding to grow supermassive
- Potential seeds for supermassive black holes
- Possible Candidates Detected by JWST*



Freese, K. et al. 2010; Ilie, C. et al. 2012; Ilie, C. et al. 2023

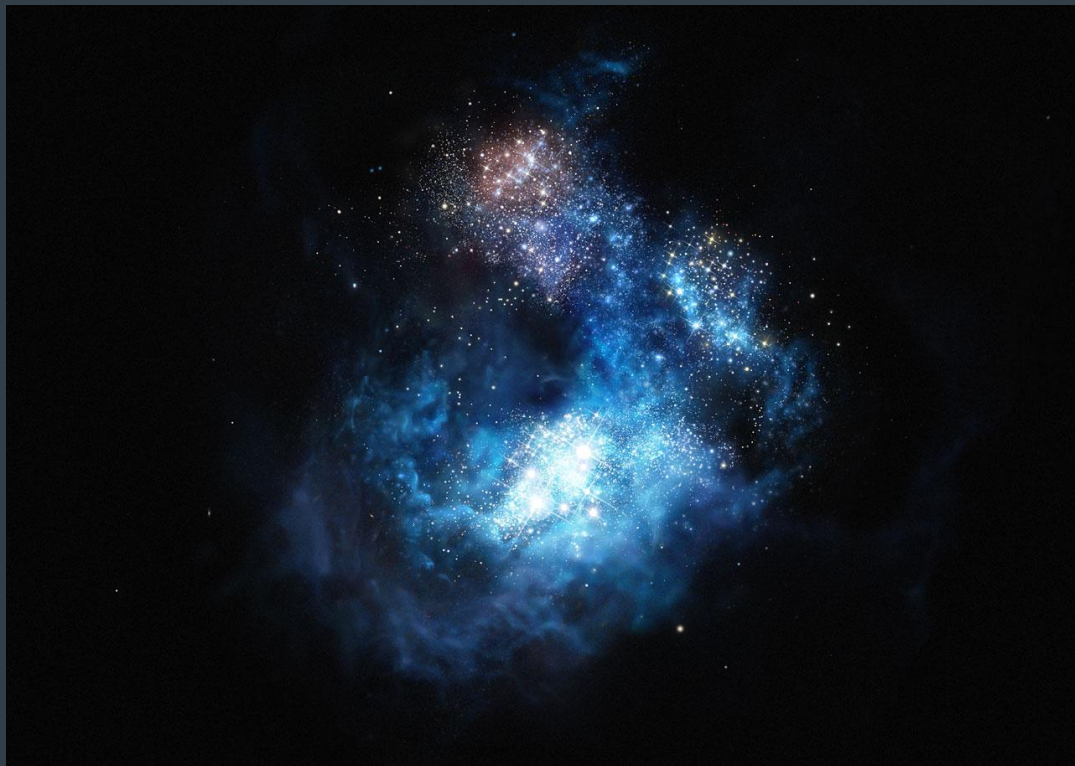
About Pop.III Galaxies

- Formed from low metallicity gas (almost pure H and He)
- Galaxies made mainly of population III stars
- Can grow as large as $\sim 10^7$ solar mass in Lambda-CDM universe*, and even larger in other scenarios

Jaacks, J., et al. 2018

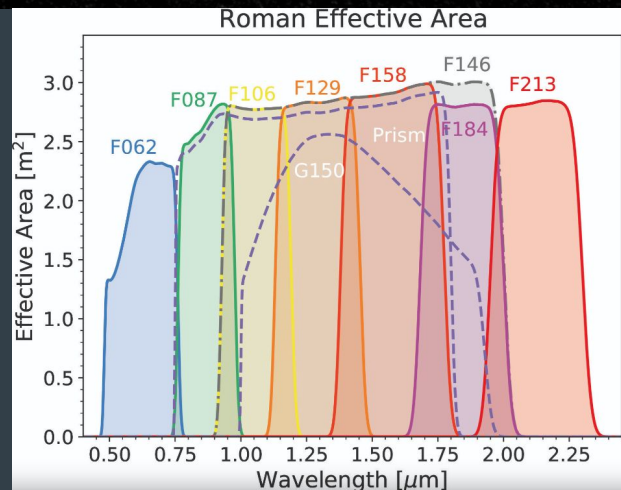
- Evolutionary spectrum modeled by YggDrasil*, and detectable by JWST

Zackrisson, E., et al. 2011



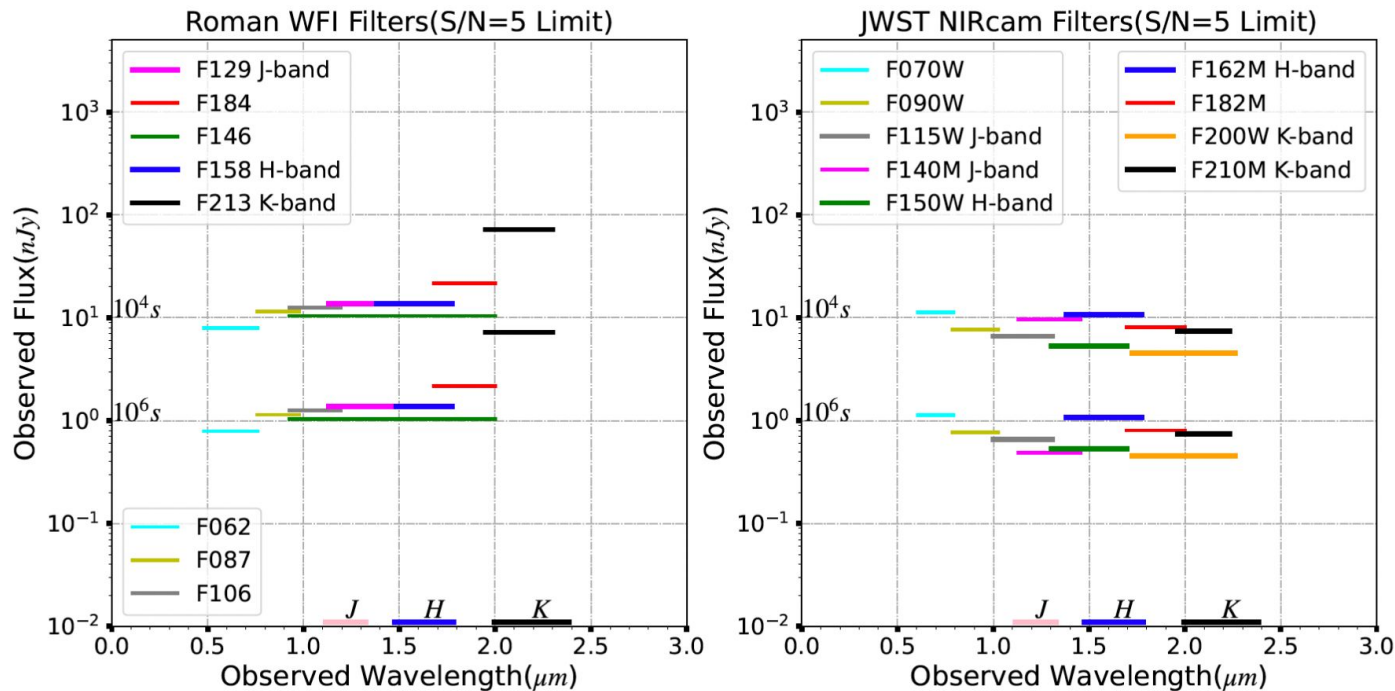
Roman Space Telescope

- Designed to launch in Oct 2026
- 0.28 deg² field of view, 100x larger field of view compared to Hubble Space Telescope
- Designed for Wide Field Survey, which will explore Dark Energy, Dark Matter, Exoplanet and Near-Infrared Science
- Wide Field Imaging covers wavelengths from ~0.5-2.3 microns
- Pandeia Engine for image simulation*



Pontoppidan, Klaus M., et al. 2016

JWST vs. Roman



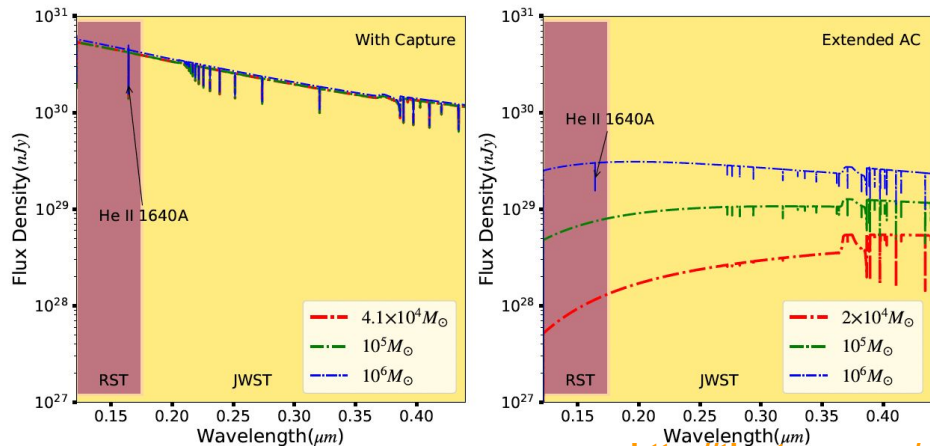
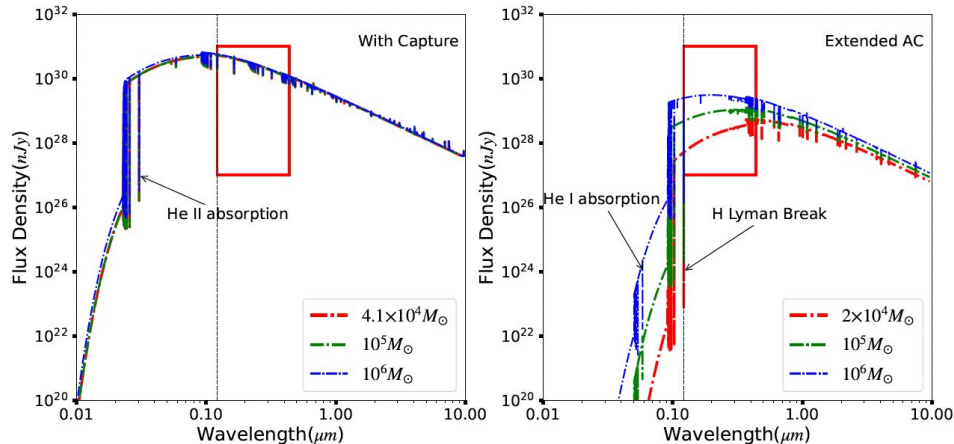
JWST vs. Roman

- Rest Frame Spectrum generated by TLUSTY*

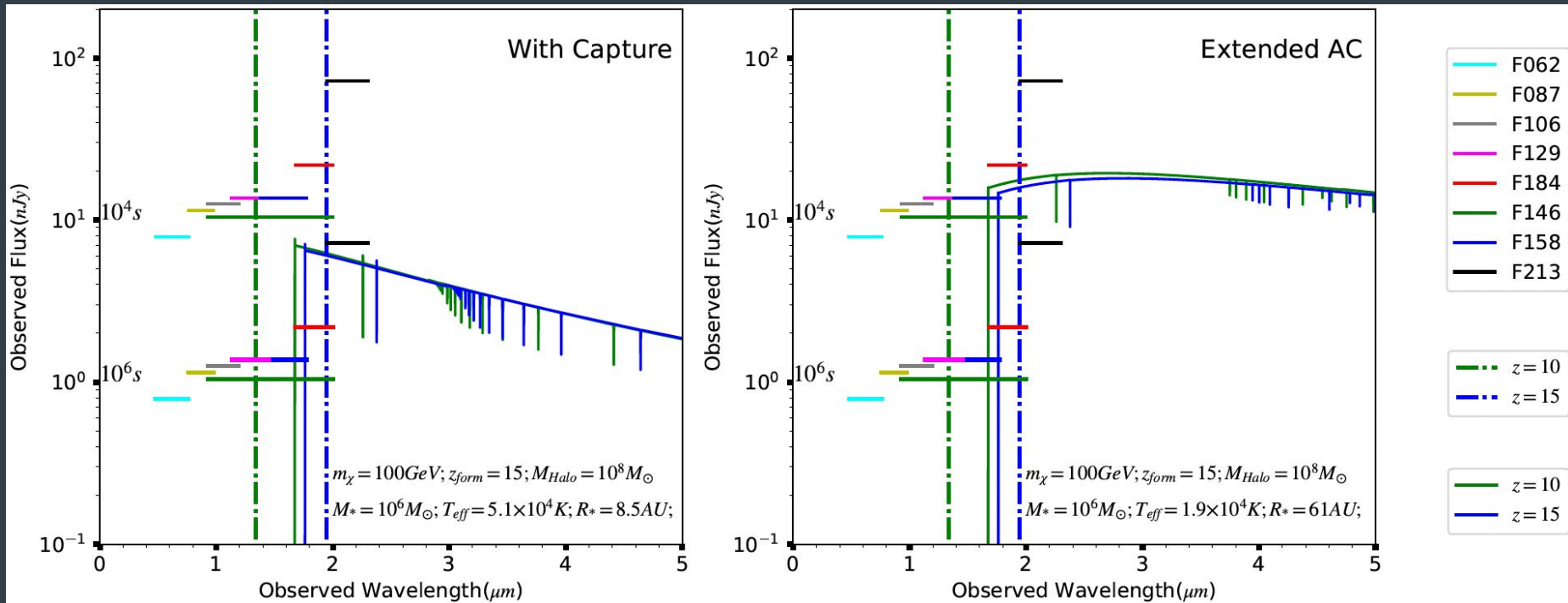
Hubeny, I., 1988

- He II 1640A absorption line as the smoking gun
- JWST covers wider wavelength while having smaller field of view compared to Roman
- Combine JWST and Roman to detect Dark Stars*

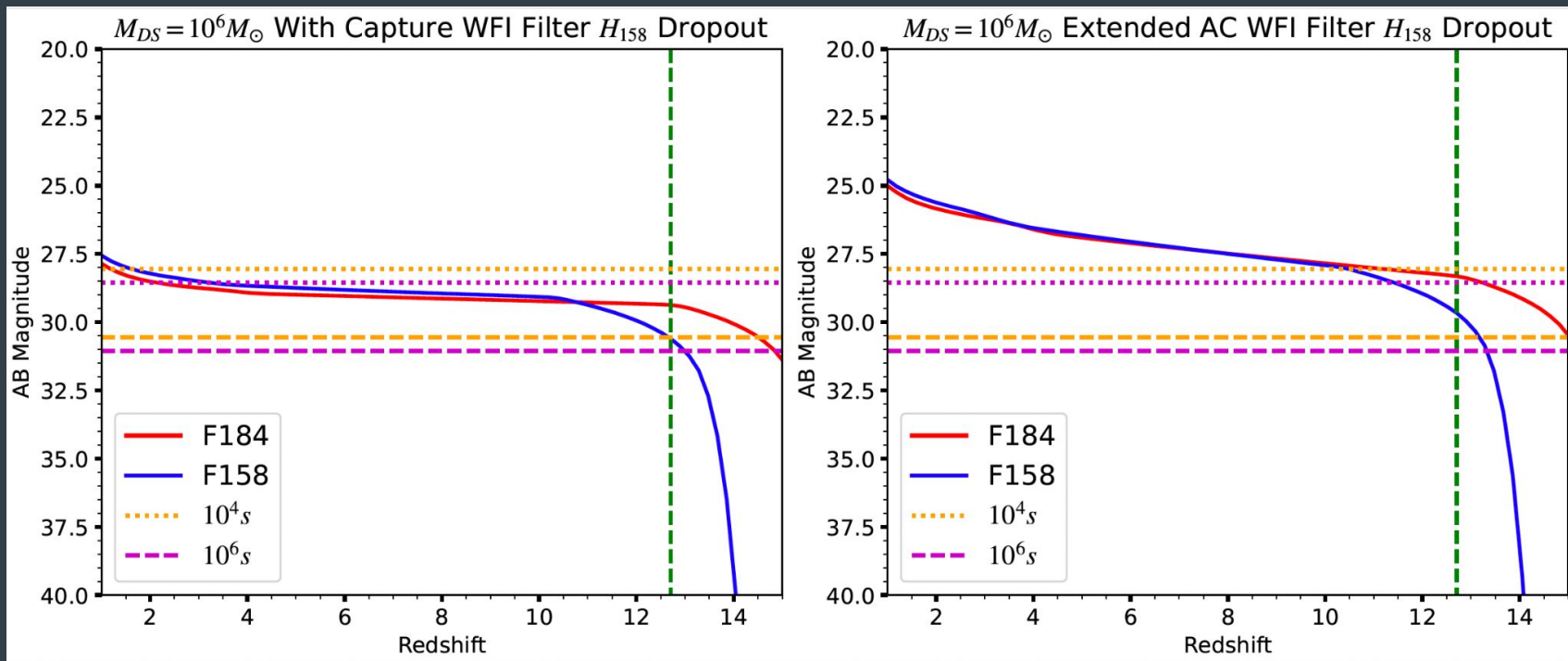
Zhang, S., et al, 2024



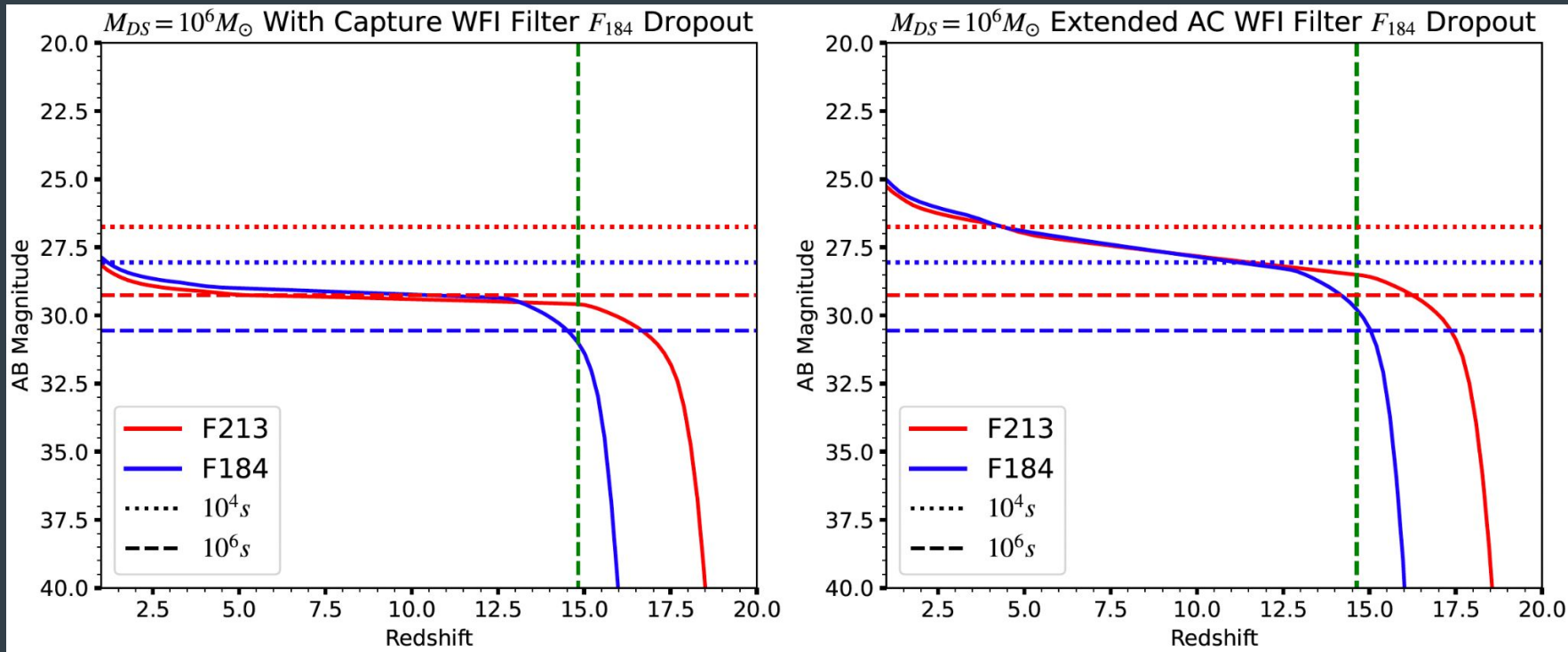
Photometry with Roman



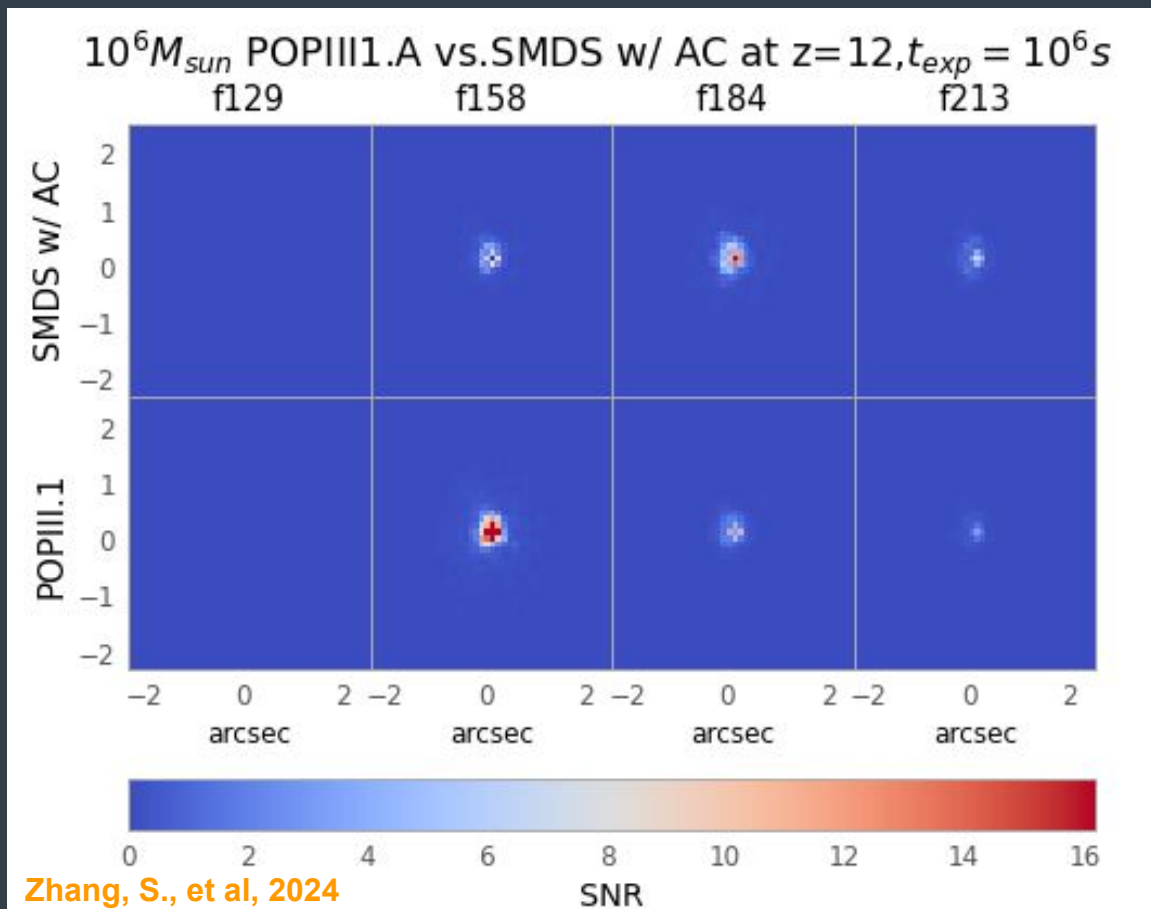
Magnitude Prediction



Magnitude Prediction

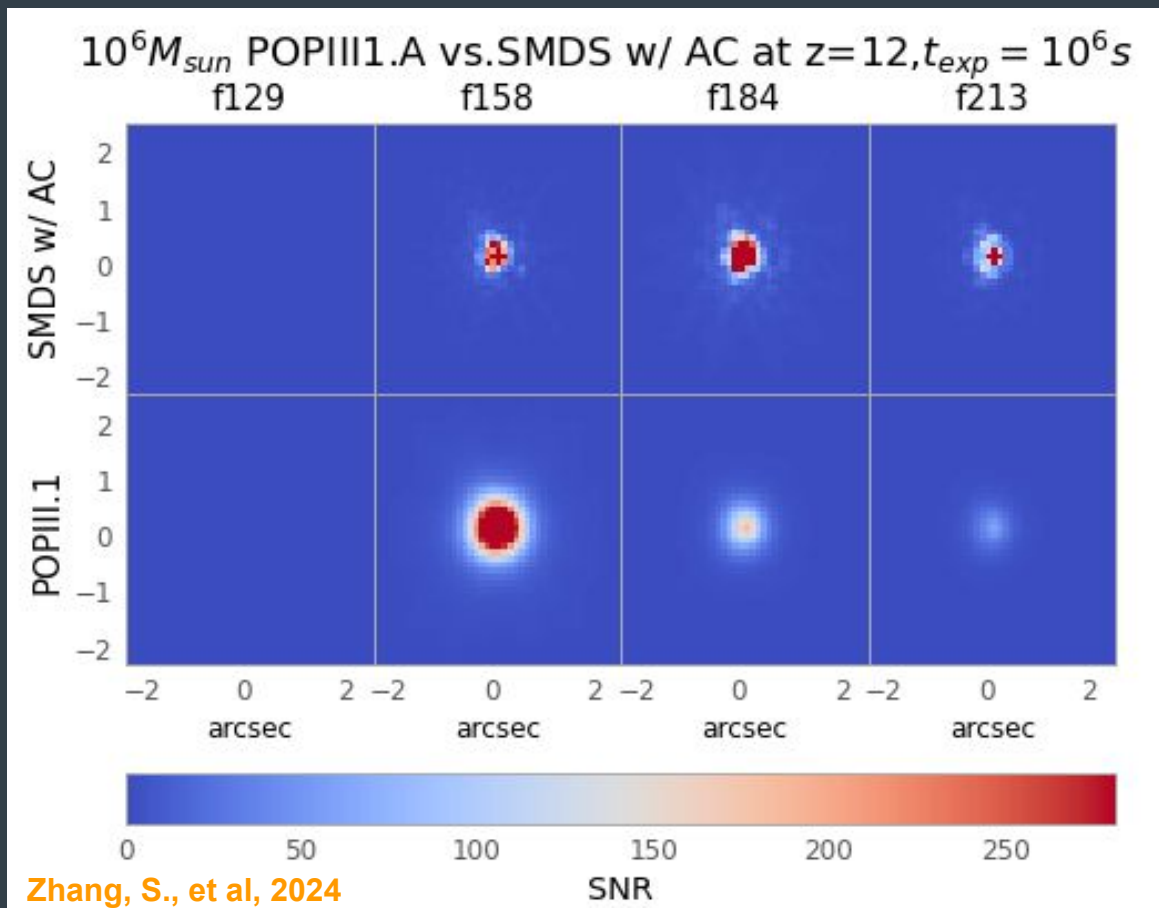


Without the effect of lensing, how the images might look like in the Roman Space Telescope?



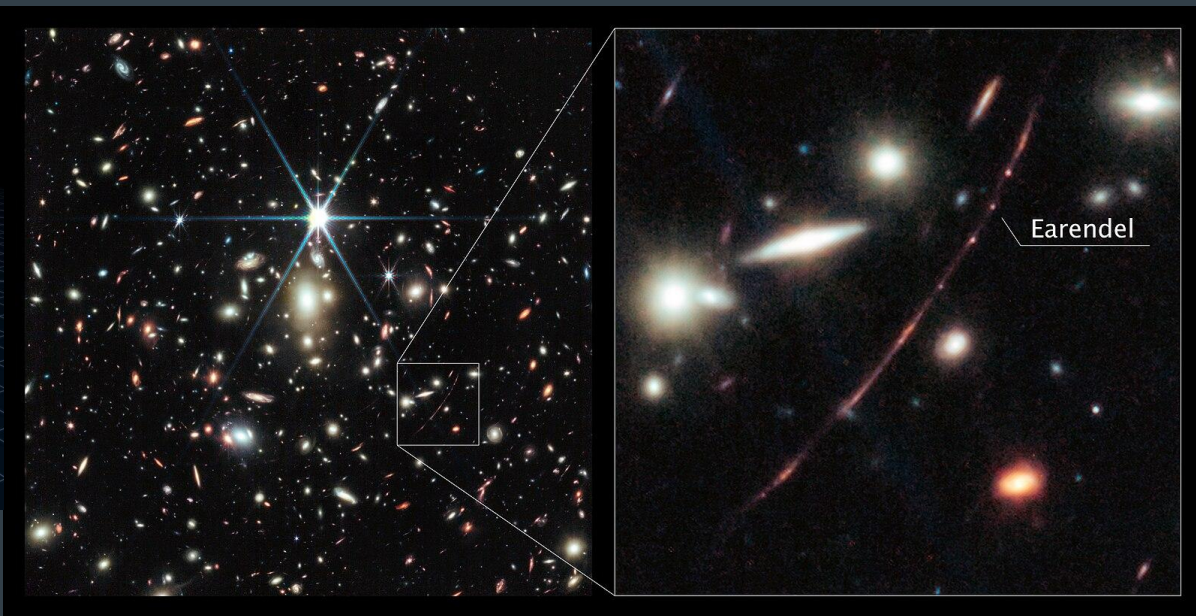
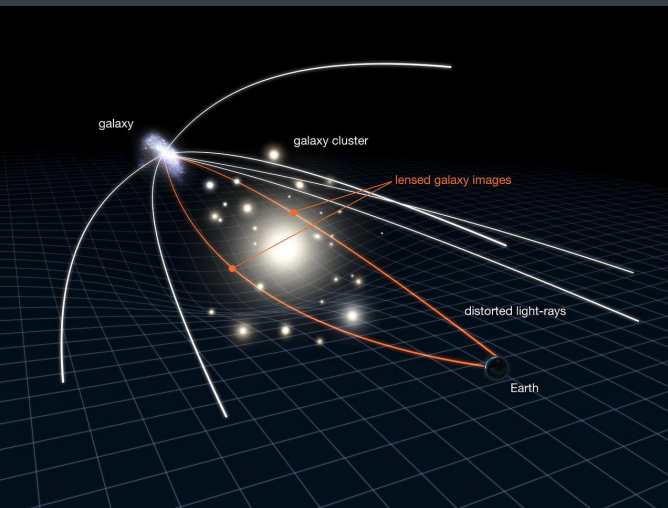
Zhang, S., et al, 2024

When the objects are magnified by 100x through gravitational lensing, then how the image might look like:



Zhang, S., et al, 2024

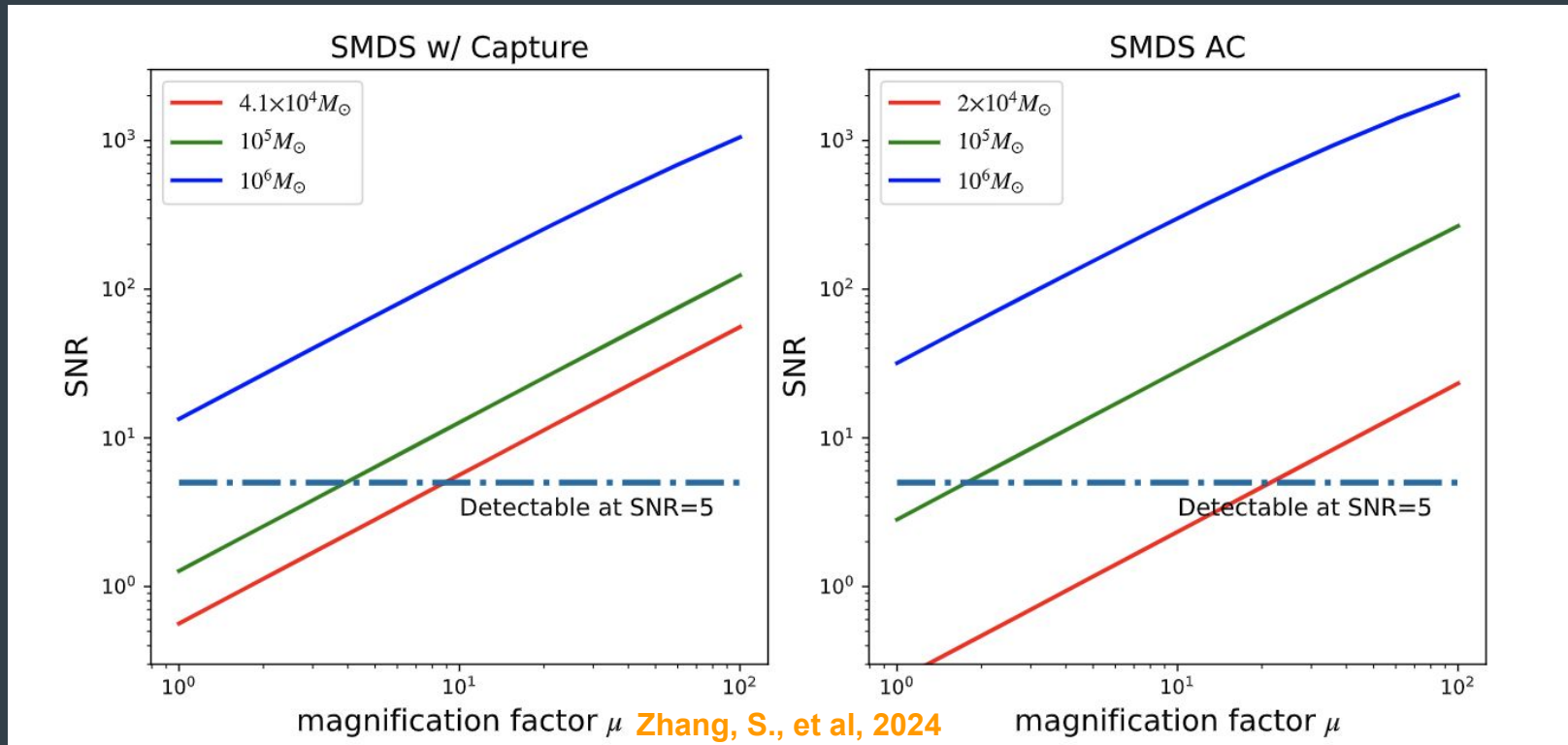
Earendel: an Example of Gravitational Lensing



<https://esahubble.org/images/heic1106c/>

Welch, B., et al 2022

Photometric Detection with Lensing

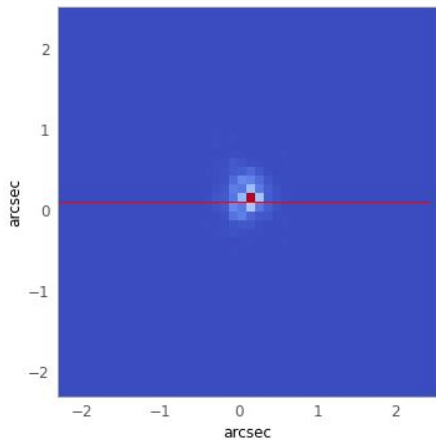
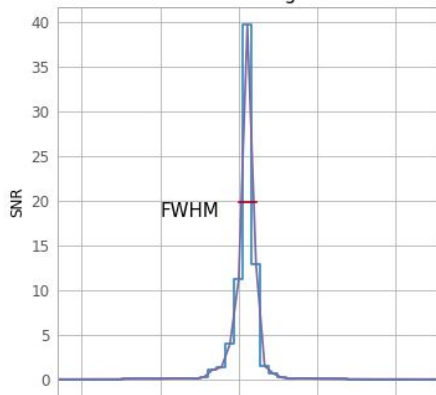


Zhang, S., et al, 2024

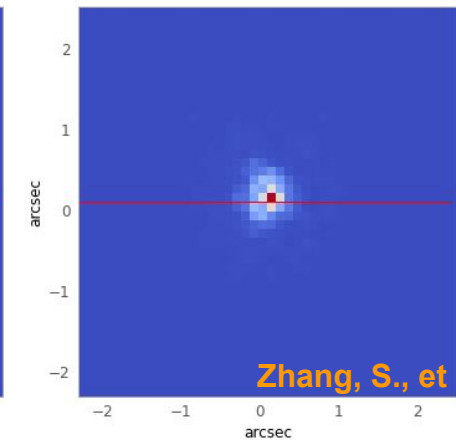
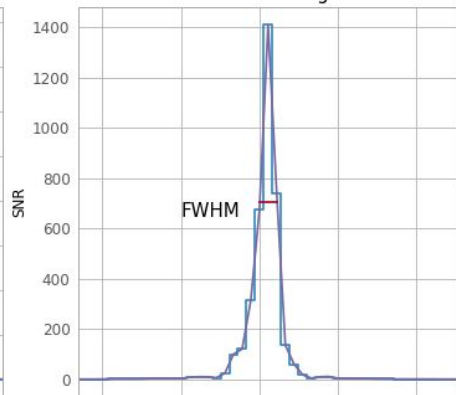
Effective size: SMDS vs. POP III galaxies

Roman PSFf184 and 1e6s total exposure at $z=12$

no lensing

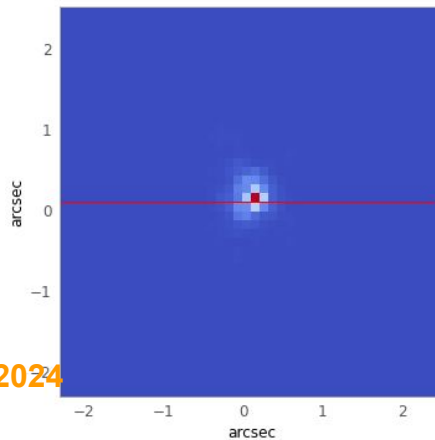
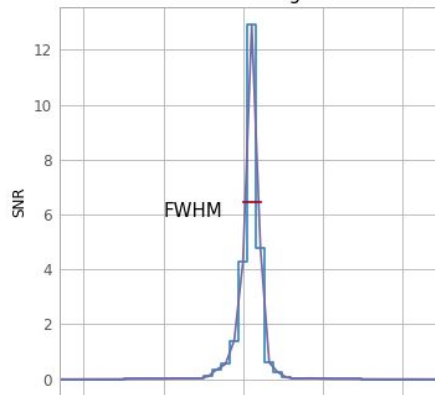


100x lensing

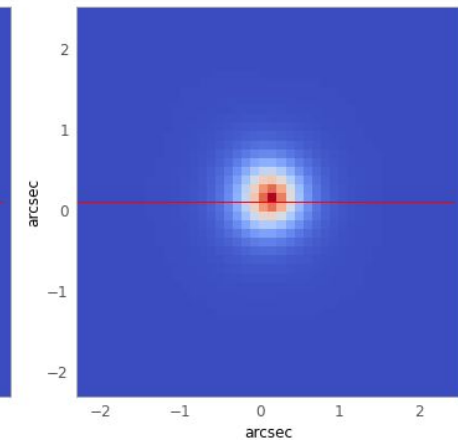
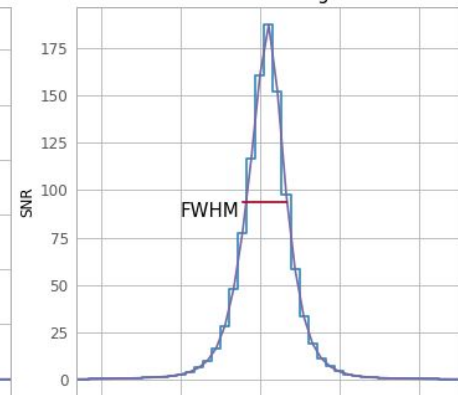


Roman PSFf184 and 1e6s total exposure at $z=12$

no lensing

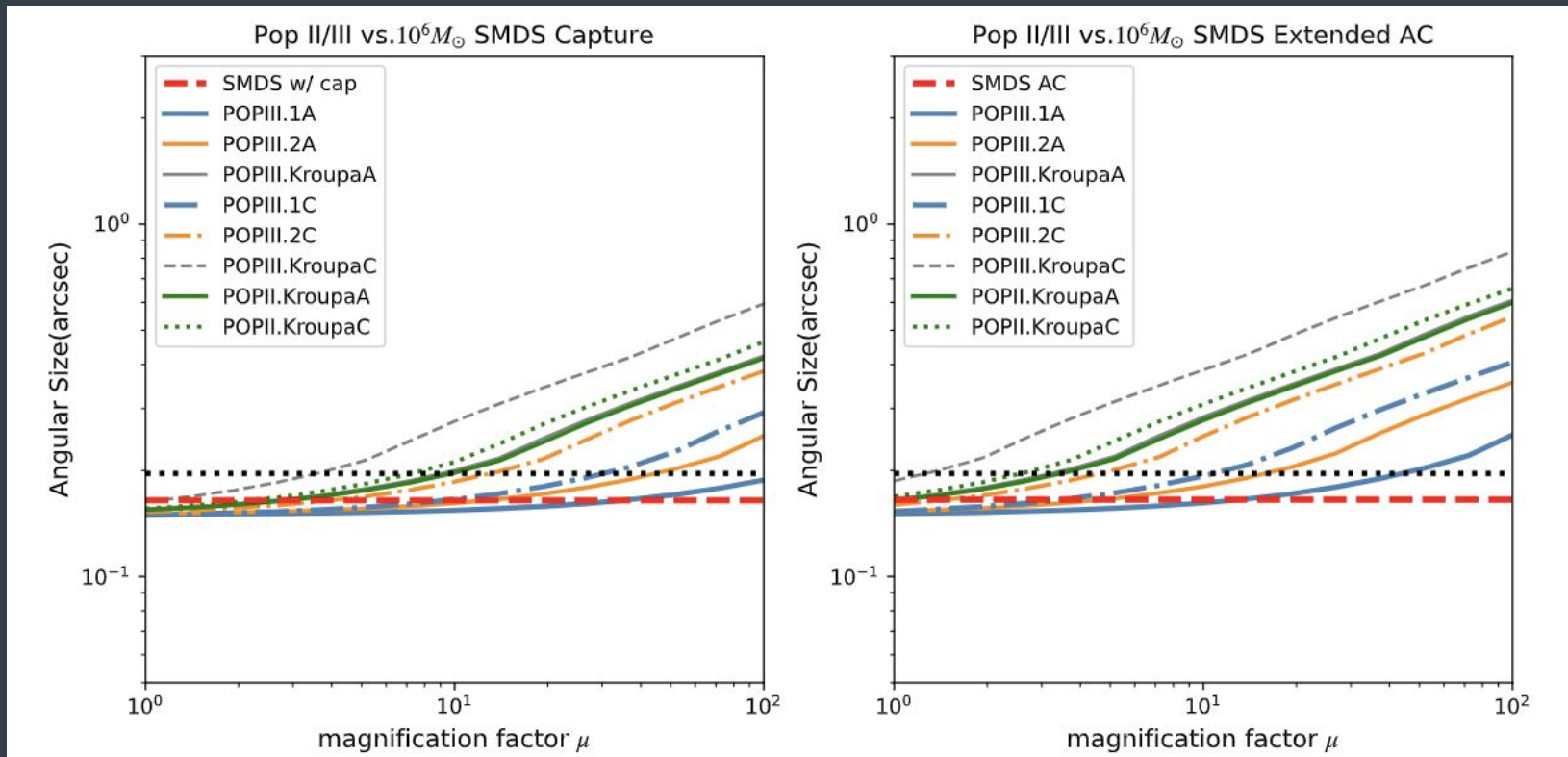


100x lensing



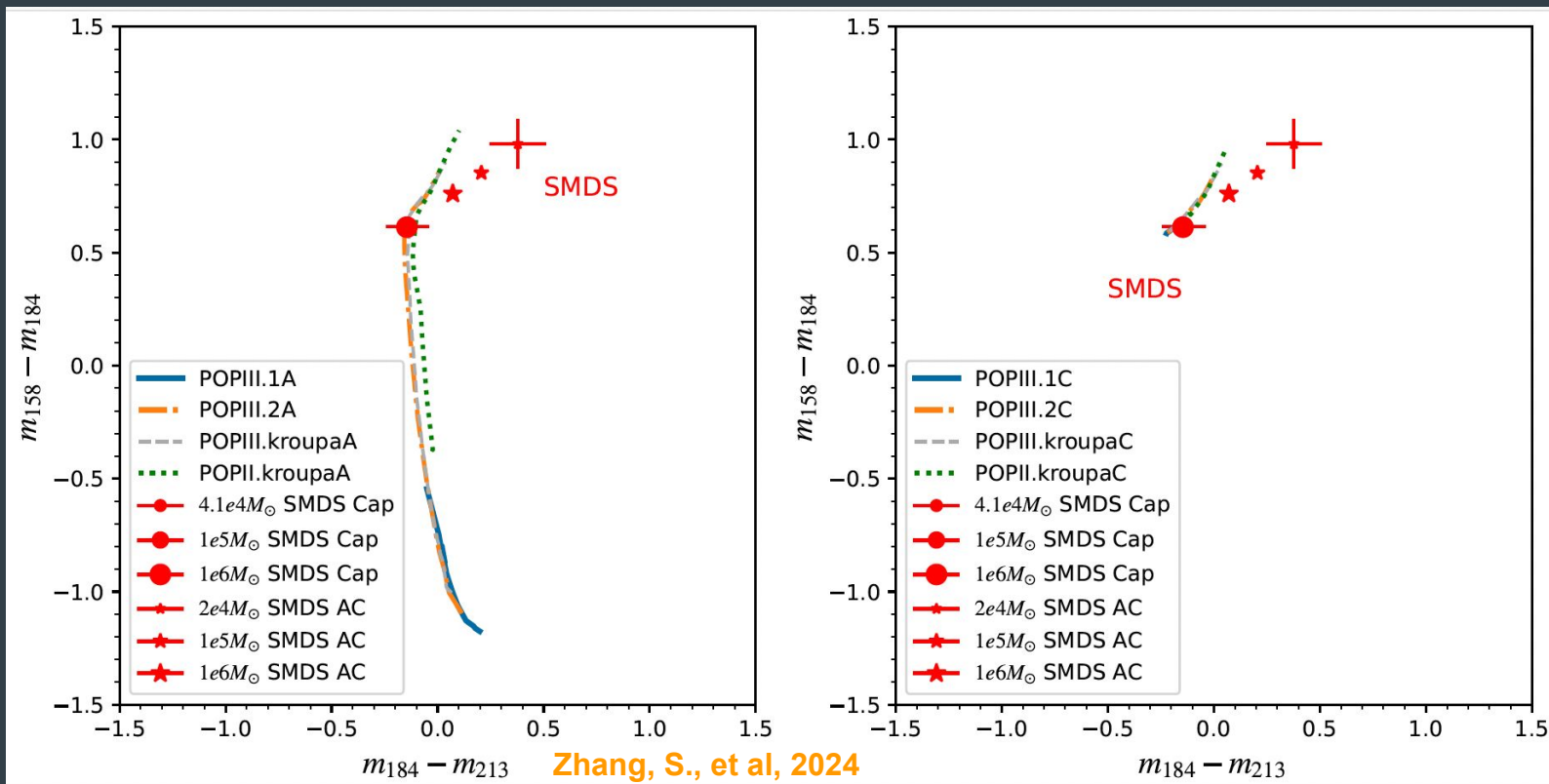
Zhang, S., et al, 2024

Effective size: SMDS vs. POP III galaxies



Zhang, S., et al, 2024

Color Index: SMDS vs. POP III galaxies



Summary of the results

- Super massive dark stars ($\sim 10^6$ solar mass) could be observed by Roman Space Telescope without lensing
- With lensing, it is possible to resolve supermassive dark stars of lower mass ($\sim 10^4$ solar mass) from other luminous objects based on color-color diagram
- With lensing by foreground clusters ($\mu \sim 10$), it is possible to tell SMDS from Pop III galaxies by their effective size
- It is more likely to distinguish super massive dark stars formed via adiabatic contraction from other objects than those formed by dark matter capture

Future Directions

- Find more possible Dark Star observational signatures
- Algorithm to distinguish between Dark Stars from objects (if enough data has been collected)
- Spectroscopic analysis of the SMDS Candidates
- Find more possible dark star candidates from available JWST dataset

Thank you!
Grazie!