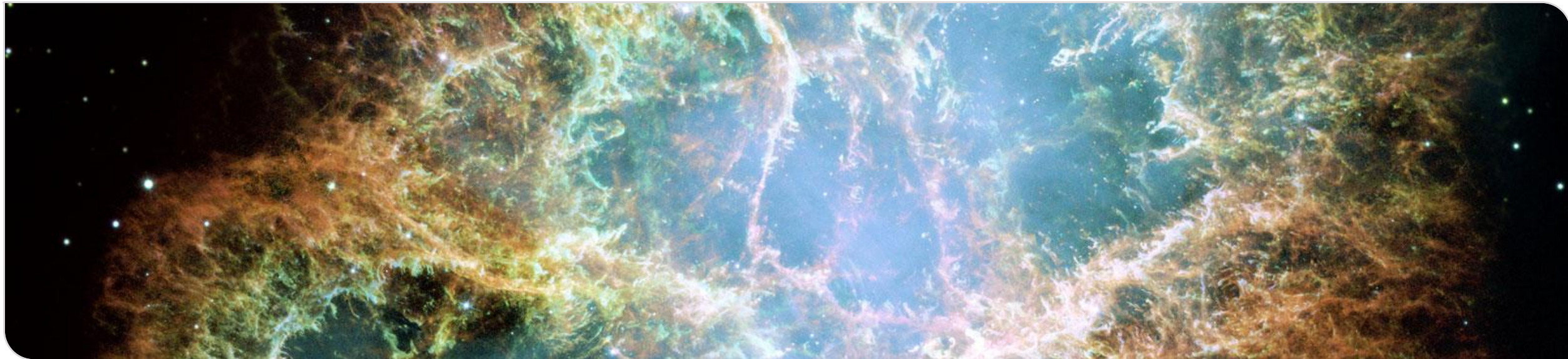


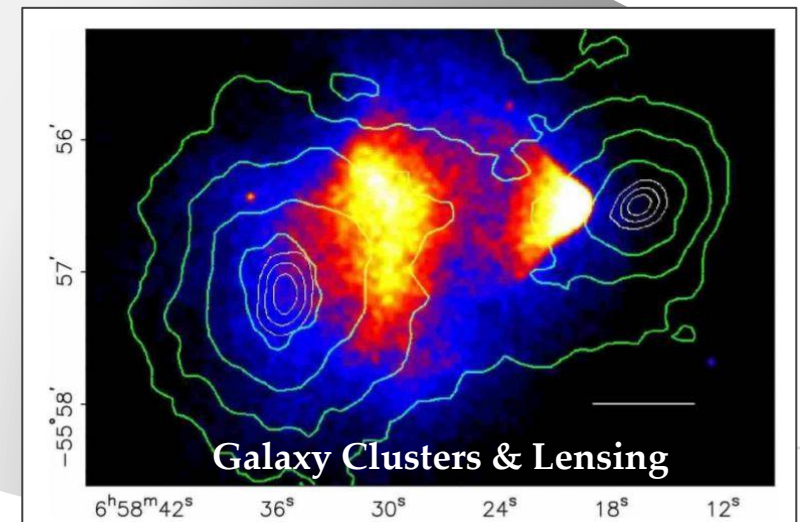
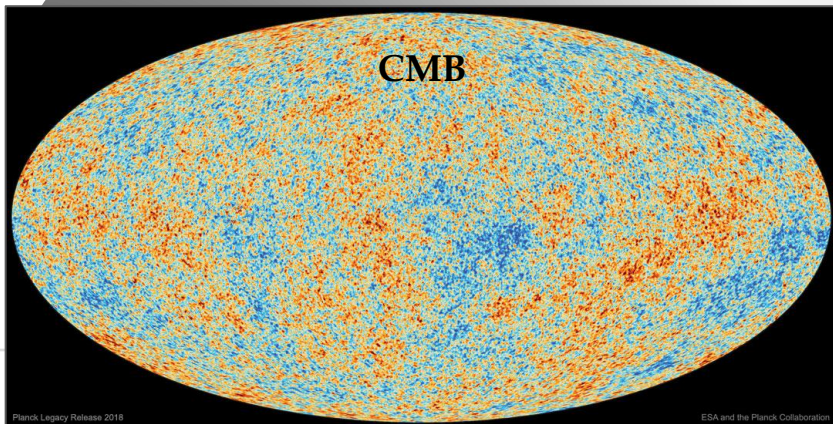
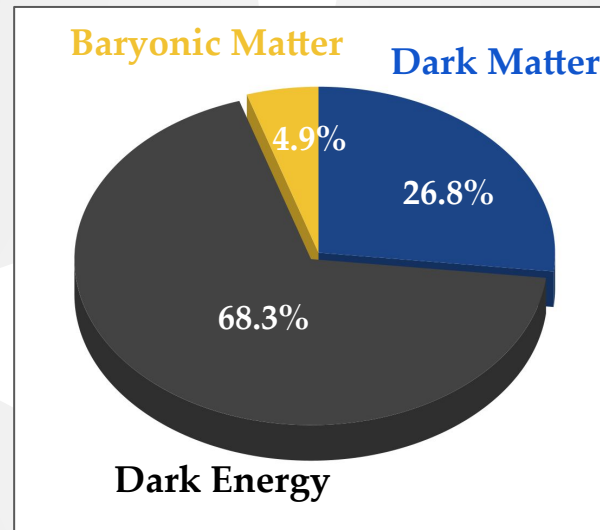
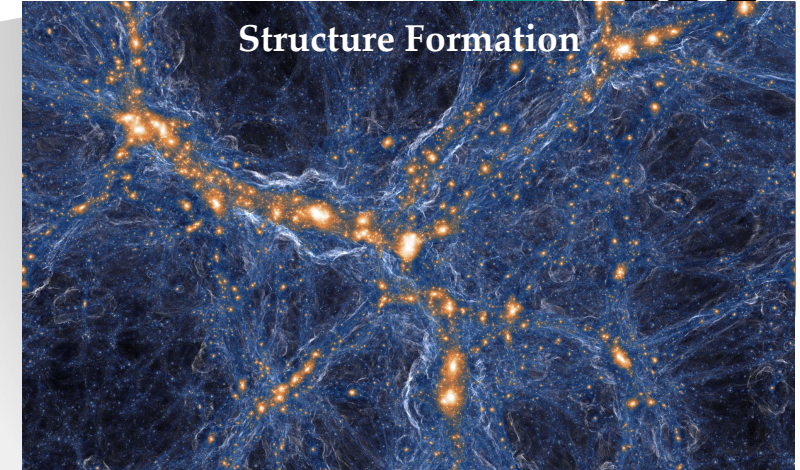
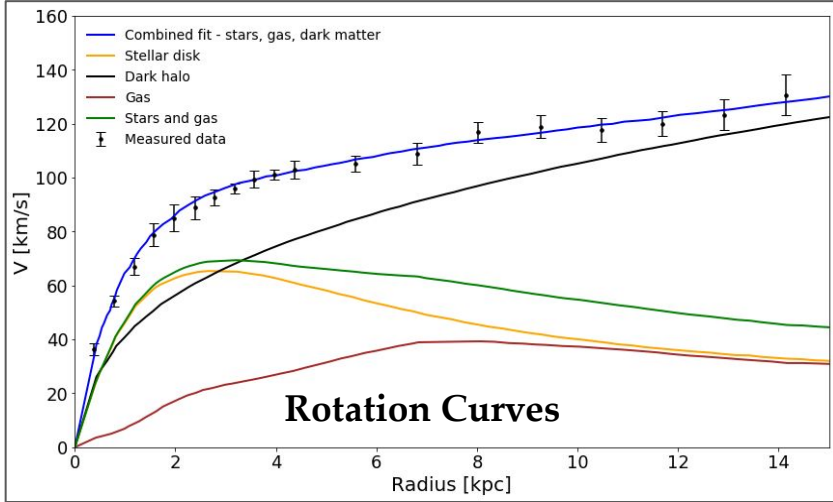
# Mineral Detection of Dark Matter and Neutrinos

*Alexey Elykov*

*Institute for Astroparticle Physics (IAP)*



# Dark Matter



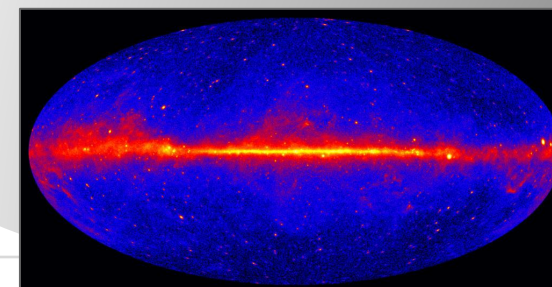
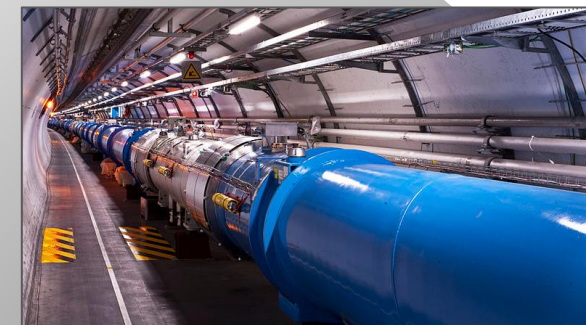
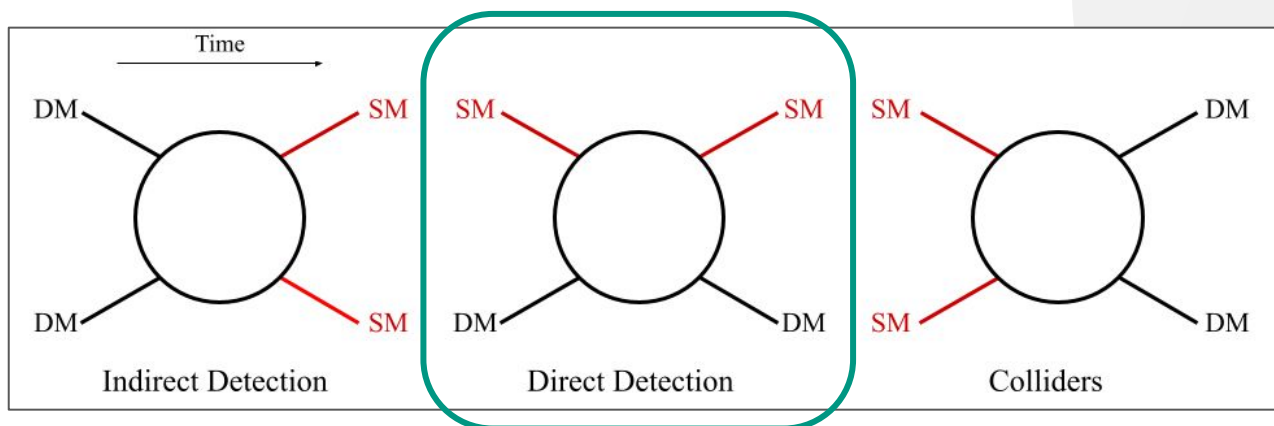
# Dark Matter

**Dark Matter:** ~ 85% of all matter in the Universe, unknown nature

**Dark Matter candidates:**

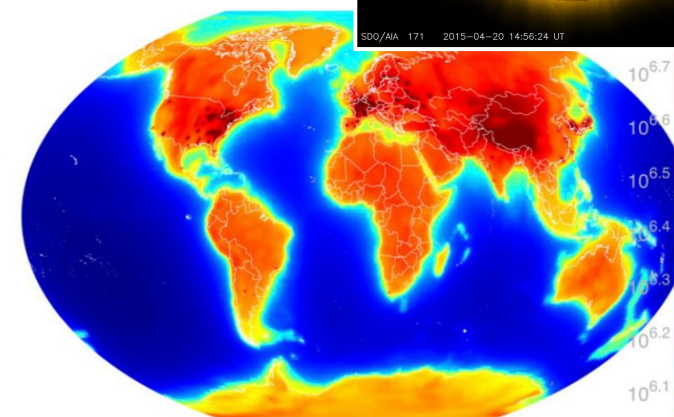
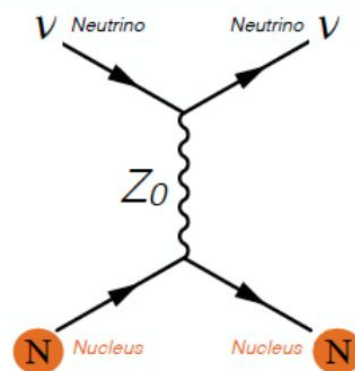
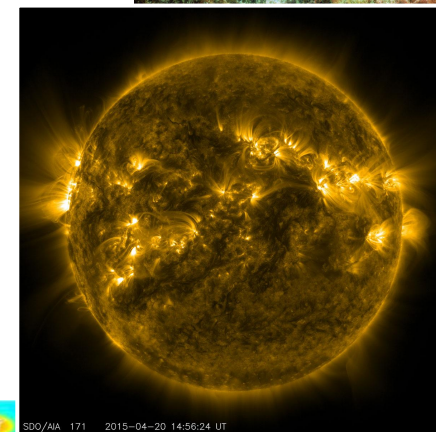
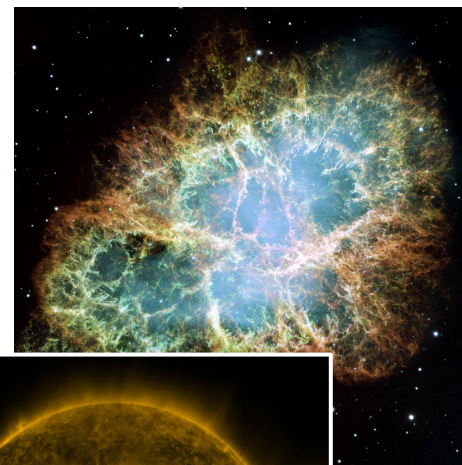
- ❖ **W**eakly **I**nteracting **M**assive **P**articles (WIMPs), mass  $\approx 10$  GeV - few TeV
- ❖ SuperWIMPs, WIMPzillas, “fuzzy” Dark Matter, Axions, ALPs ... etc...

## Paths for Dark Matter detection



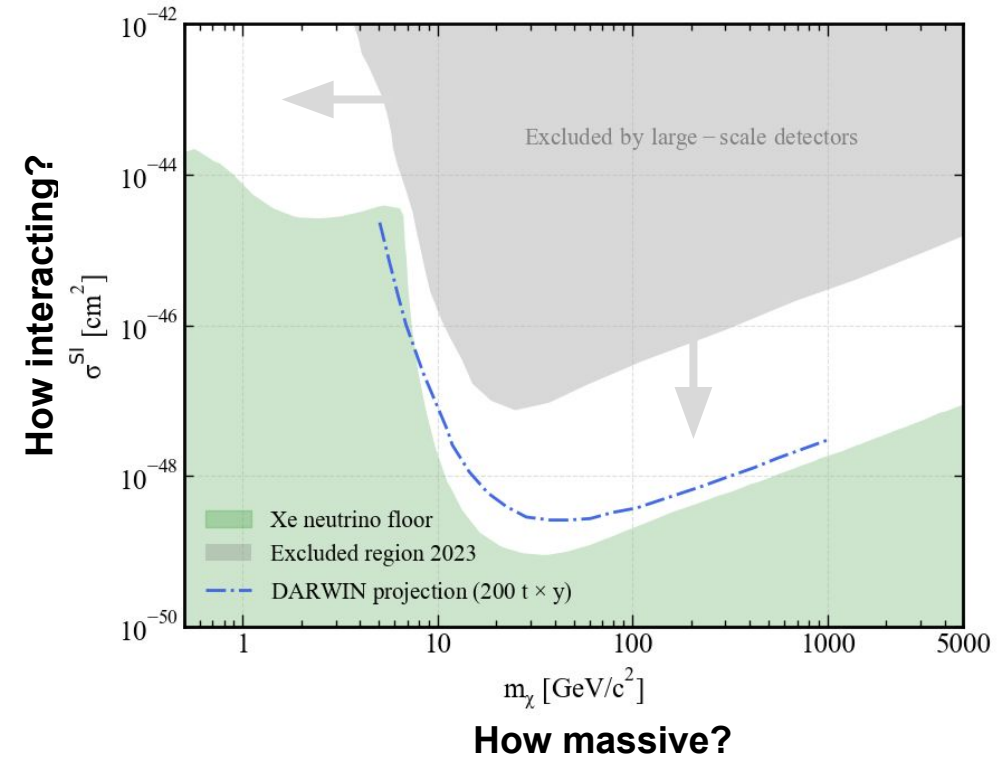
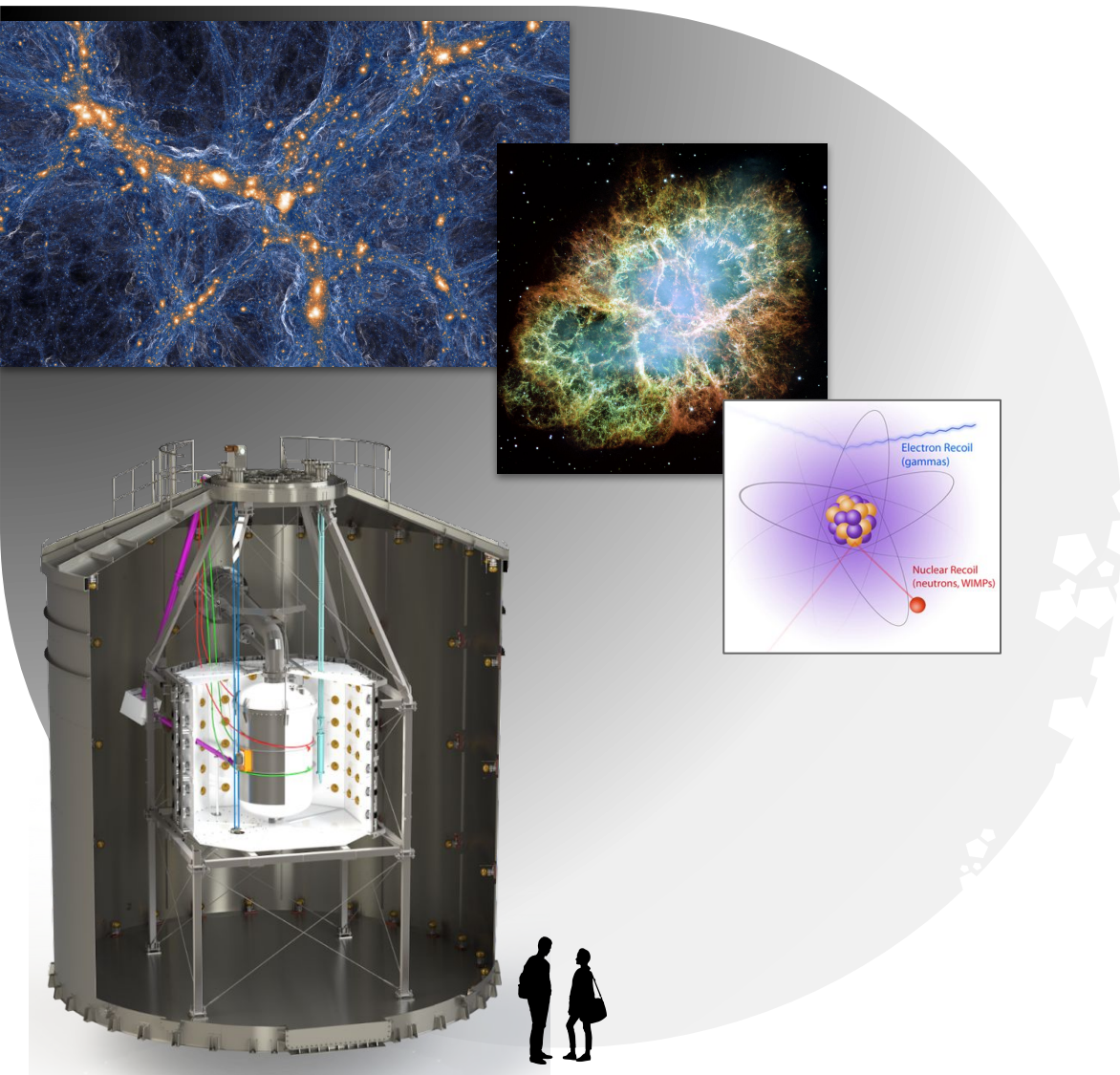
# Neutrinos

- ❖ Properties are still largely unknown
- ❖ Can shed light on fundamental open questions
- ❖ Large range of energies
- ❖ **Astrophysical messengers (history & evolution):**
  - Sun
  - Supernovae
  - Cosmic-rays
  - Galactic & extragalactic
- ❖ Geoneutrinos



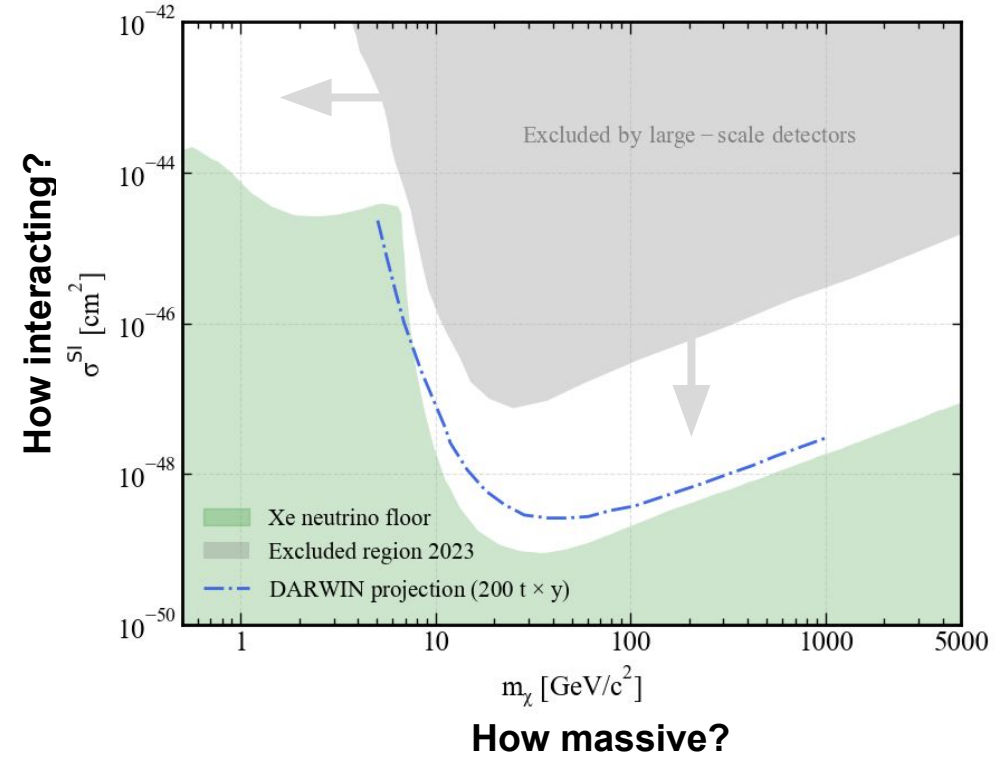
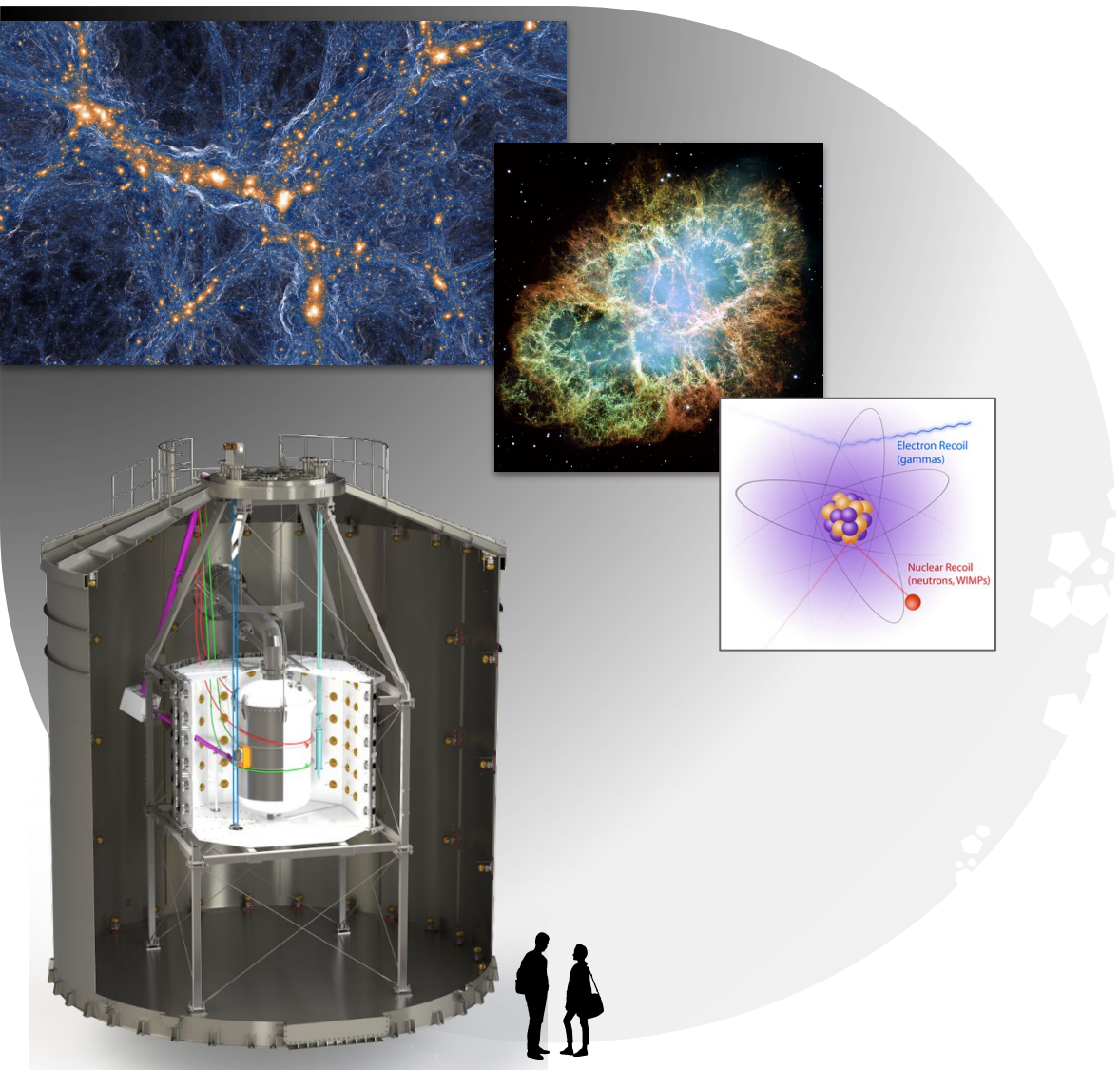
CEvNS

# Dark Matter & Neutrinos - Detection



**Dark Matter still eludes detection & we are still puzzled by the nature of neutrinos!**

# Dark Matter & Neutrinos - Detection



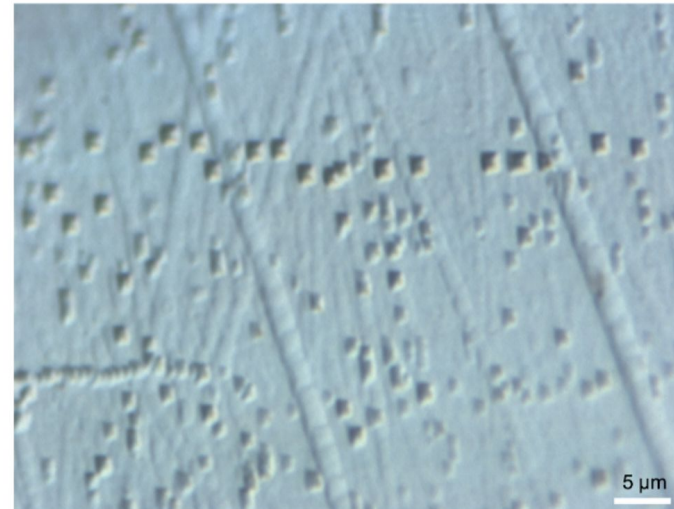
**Think  
outside  
the box**

X	O	X
X	<del>O</del>	X
O	X	<del>O</del>

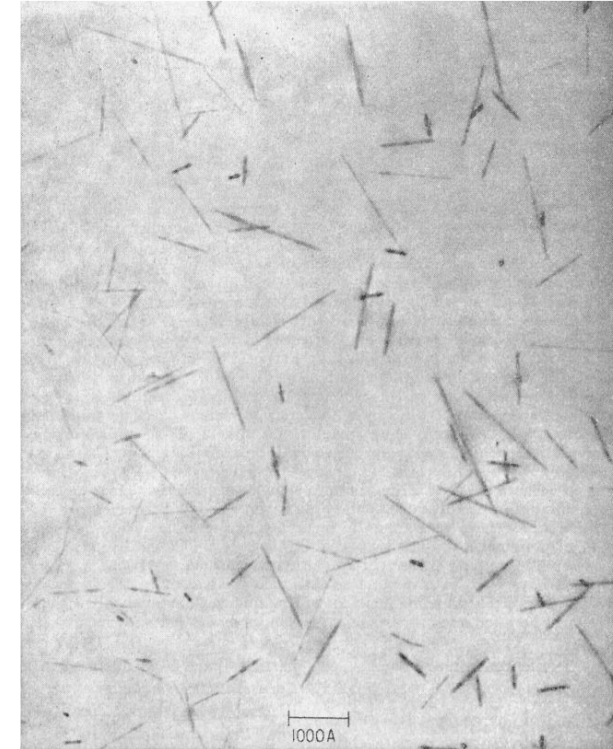


# Solid State Nuclear Track Detectors (SSNTDs)

- ❖ **SSNTDs - natural & synthetic crystals**
  - Geology & geophysics
  - Radiation damage
  - Cosmochemistry
  - Material science
  - Astrophysics
  
- ❖ Ionizing radiation produces damage tracks
  
- ❖ Chemical etching
  
- ❖ Readout with microscopy



*Etch pits in Olivine - courtesy of U. Glasmacher*



*1963 : Fission tracks in synthetic mica as viewed by TEM (DOI: 10.1029/JZ068i016p04847)*

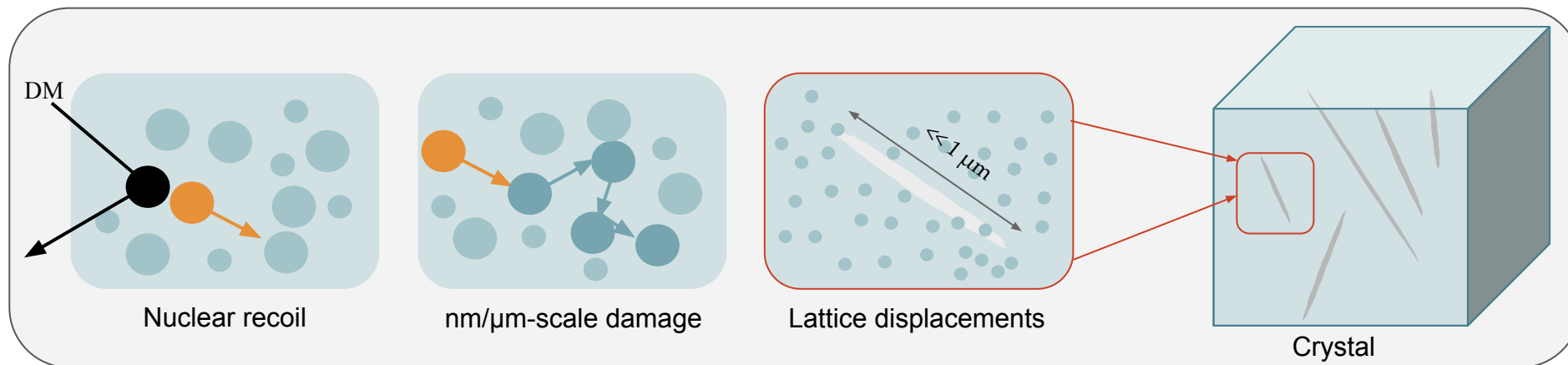
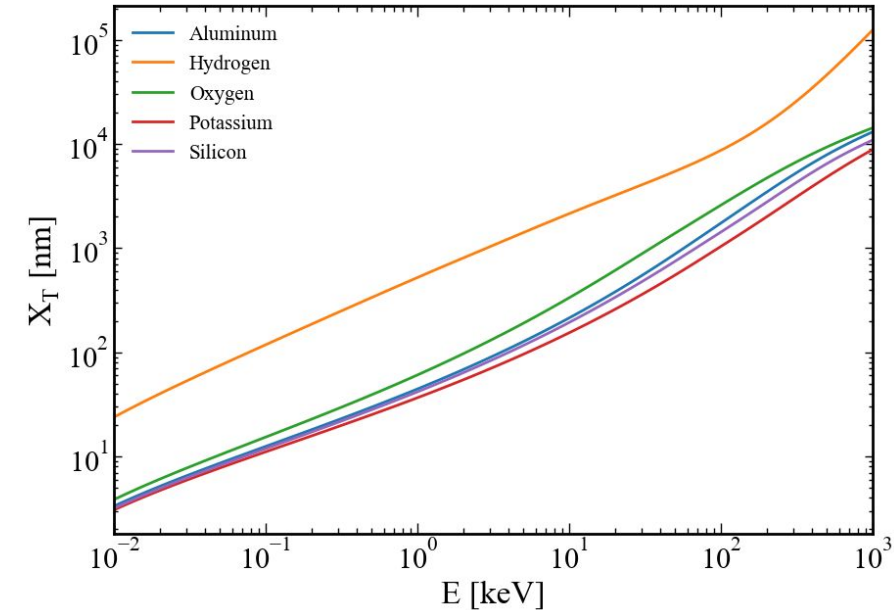
# Track & Damage Features Formation

## ❖ Energy loss in solid materials due to :

- Electronic stopping (off electron clouds)
- Nuclear stopping (off nuclei)

$$x_T(E_R) = \int_0^{E_R} \left| \frac{dE}{dx_T} \right|^{-1} dE$$

- Nuclear recoils down to 0.1 - 1 keV





People tried this in the past!

VOLUME 56, NUMBER 12

PHYSICAL REVIEW LETTERS

24 MARCH 1986

## Search for Supermassive Magnetic Monopoles Using Mica Crystals

P. B. Price and M. H. Salamon

*Department of Physics, University of California, Berkeley, California 94720*  
(Received 18 November 1985)

## Nuclear tracks from Cold Dark Matter interactions in mineral crystals: a computational study

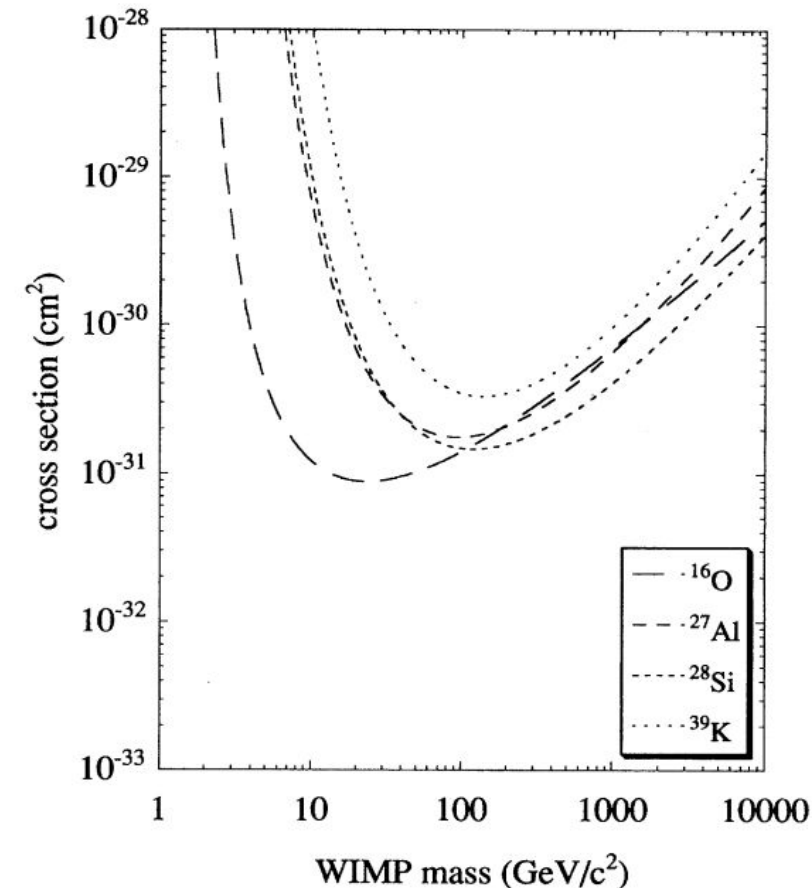
J.I. Collar \*, F.T. Avignone III

*Department of Physics and Astronomy, University of South Carolina, Columbia, SC 29208, USA*  
Received 26 July 1994; revised form received 8 November 1994

## Limits on Dark Matter Using Ancient Mica

D. P. Snowden-Ifft,\* E. S. Freeman, and P. B. Price\*

*Physics Department, University of California at Berkeley, Berkeley, California 94720*  
(Received 20 September 1994)



1995 : Limits on Dark Matter Using Ancient Mica  
(DOI: 10.1103/PhysRevLett.74.4133)

# Mineral-Detectors Now

Digging for dark matter: Spectral analysis and discovery potential of paleo-detectors

Thomas D. P. Edwards, Bradley J. Kavanagh, Christoph Weniger, Sebastian Baum, Andrzej K. Drukier, Katherine Freese, Maciej Górski, and Patrick Stengel  
Phys. Rev. D **99**, 043541 – Published 27 February 2019

Measuring Changes in the Atmospheric Neutrino Rate over Gigayear Timescales

Johnathon R. Jordan, Sebastian Baum, Patrick Stengel, Alfredo Ferrari, Maria Cristina Morone, Paola Sala, and Joshua Spitz  
Phys. Rev. Lett. **125**, 231802 – Published 30 November 2020

Paleodetectors for Galactic supernova neutrinos

Sebastian Baum, Thomas D. P. Edwards, Bradley J. Kavanagh, Patrick Stengel, Andrzej Katherine Freese, Maciej Górski, and Christoph Weniger  
Phys. Rev. D **101**, 103017 – Published 13 May 2020

Rocks, water, and noble liquids: Unfolding the flavor content of neutrinos

Sebastian Baum, Francesco Capozzi, and Shunsaku Horiuchi  
Phys. Rev. D **106**, 123008 – Published 9 December 2022

❖ **Worldwide interest - novel emerging research field**



# Mineral-Detectors Now

Digging for dark matter: Spectral analysis and discovery potential of paleo-detectors

Thomas D. P. Edwards, Bradley J. Kavanagh, Christoph Weniger, Sebastian Baum, Andrzej K. Drukier, Katherine Freese, Maciej Górski, and Patrick Stengel  
Phys. Rev. D **99**, 043541 – Published 27 February 2019

Measuring Changes in the Atmospheric Neutrino Rate over Gigayear Timescales

Johnathon R. Jordan, Sebastian Baum, Patrick Stengel, Alfredo Ferrari, Maria Cristina Morone, Paola Sala, and Joshua Spitz  
Phys. Rev. Lett. **125**, 231802 – Published 30 November 2020

Paleodetectors for Galactic supernova neutrinos

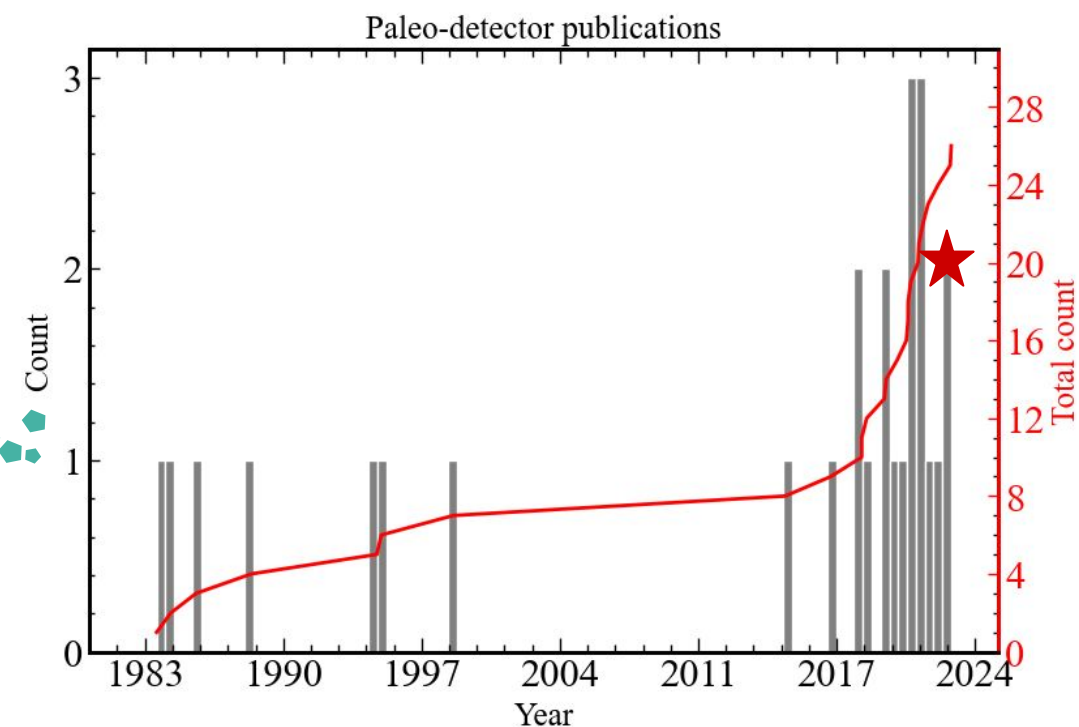
Sebastian Baum, Thomas D. P. Edwards, Bradley J. Kavanagh, Patrick Stengel, Andrzej Katherine Freese, Maciej Górski, and Christoph Weniger  
Phys. Rev. D **101**, 103017 – Published 13 May 2020

Rocks, water, and noble liquids: Unfolding the flavor content of neutrinos

Sebastian Baum, Francesco Capozzi, and Shunsaku Horiuchi  
Phys. Rev. D **106**, 123008 – Published 9 December 2022




- ❖ **Worldwide interest - novel emerging research field**
- ❖ **Growing amount of publications in last ~ 5 yr**




- ❖ **Worldwide interest - novel emerging research field**
- ❖ **White paper in “*Physics of the Dark Universe*” (editor’s invitation)**
  - 67 authors, 46 institutions, 113 pages




- ❖ **MDvDM Jan. 2024 - Virginia Tech, USA**



Physics of the Dark Universe  
Volume 41, August 2023, 101245



## Mineral detection of neutrinos and dark matter. A whitepaper

[Sebastian Baum](#)<sup>1</sup>  , [Patrick Stengel](#)<sup>2</sup> , [Natsue Abe](#)<sup>3</sup>, [Javier F. Acevedo](#)<sup>4</sup>, [Gabriela R. Araujo](#)<sup>5</sup> <sup>a</sup>, [Yoshihiro Asahara](#)<sup>6</sup>, [Frank Avignone](#)<sup>7</sup>, [Levente Balogh](#)<sup>8</sup>, [Laura Baudis](#)<sup>5</sup>, [Yilda Boukhtouchen](#)<sup>9</sup>, [Joseph Bramante](#)<sup>9</sup> <sup>10</sup>, [Pieter Alexander Breur](#)<sup>4</sup>, [Lorenzo Caccianiga](#)<sup>11</sup>, [Francesco Capozzi](#)<sup>12</sup>, [Juan I. Collar](#)<sup>13</sup>, [Reza Ebadi](#)<sup>14</sup> <sup>15</sup>, [Thomas Edwards](#)<sup>16</sup>, [Klaus Eitel](#)<sup>17</sup>, [Alexey Elykov](#)<sup>17</sup>, [Rodney C. Ewing](#)<sup>18</sup>, [Katherine Freese](#)<sup>19</sup> <sup>20</sup>, [Audrey Fung](#)<sup>9</sup>, [Claudio Galelli](#)<sup>21</sup>, [Ulrich A. Glasmacher](#)<sup>22</sup>, [Arianna Gleason](#)<sup>4</sup>, [Noriko Hasebe](#)<sup>23</sup>, [Shigenobu Hirose](#)<sup>24</sup>, [Shunsaku Horiuchi](#)<sup>25</sup> <sup>26</sup>, [Yasushi Hoshino](#)<sup>27</sup>, [Patrick Huber](#)<sup>25</sup> <sup>a</sup>, [Yuki Ido](#)<sup>28</sup>, [Yohei Igami](#)<sup>29</sup>, [Norito Ishikawa](#)<sup>30</sup>,



astro-ph > arXiv:2405.01626

Search...  
Help | Adv

Astrophysics > Cosmology and Nongalactic Astrophysics

[Submitted on 2 May 2024]

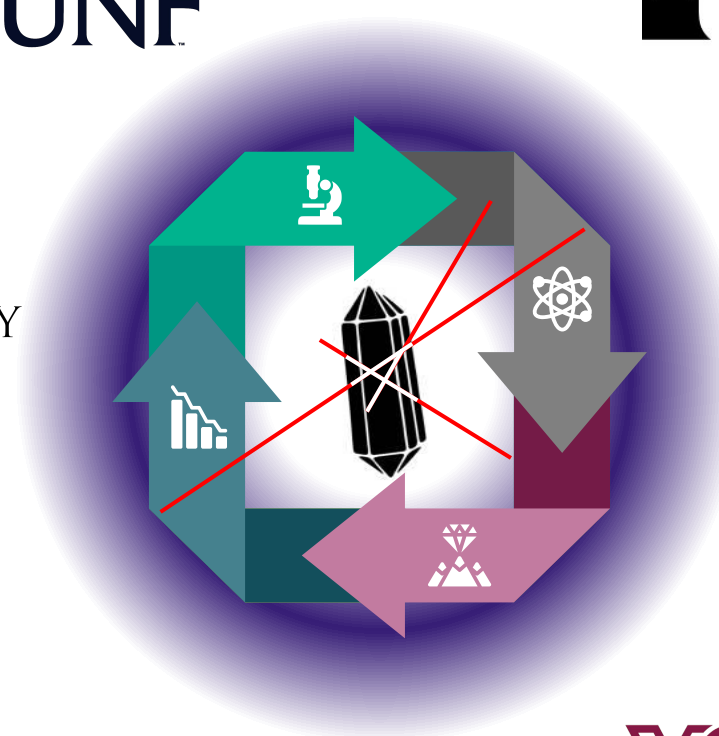
## Mineral Detection of Neutrinos and Dark Matter 2024. Proceedings

[Sebastian Baum](#), [Patrick Huber](#), [Patrick Stengel](#), [Natsue Abe](#), [Daniel G. Ang](#), [Lorenzo Apollonio](#), [Gabriela R. Araujo](#), [Levente Balogh](#), [Pranshu Bhaumik](#), [Yilda Boukhtouchen](#), [Joseph Bramante](#), [Lorenzo Caccianiga](#), [Andrew Calabrese-Day](#), [Qing Chang](#), [Juan I. Collar](#), [Reza Ebadi](#), [Alexey Elykov](#), [Katherine Freese](#), [Audrey Fung](#), [Claudio Galelli](#), [Arianna E. Gleason](#), [Mariano Guerrero Perez](#), [Janina Hakenmüller](#), [Takeshi Hanyu](#), [Noriko Hasebe](#), [Shigenobu Hirose](#), [Shunsaku Horiuchi](#), [Yasushi Hoshino](#), [Yuki Ido](#), [Vsevolod Ivanov](#), [Takashi Kamiyama](#), [Takenori Kato](#), [Yoji Kawamura](#), [Chris Kelso](#), [Giti A. Khodaparast](#), [Emilie M. LaVoie-Ingram](#), [Matthew Leybourne](#), [Xingxin Liu](#), [Thalles Lucas](#), [Brenden A. Magill](#), [Federico M. Mariani](#), [Charlotte Mkhonto](#), [Hans Pieter Mumm](#), [Kohta Murase](#), [Tatsuhiko Naka](#), [Kenji Oguni](#), [Kathryn Ream](#), [Kate Scholberg](#), [Maximilian Shen](#), [Joshua Spitz](#), [Katsuhiko Suzuki](#), [Alexander Takla](#), [Jiashen Tang](#), [Natalia Tapia-Arellano](#), [Pieter Vermeesch](#), [Aaron C. Vincent](#), [Nikita Vladimirov](#), [Ronald Walsworth](#), [David Waters](#), [Greg Wurtz](#), [Seiko Yamasaki](#), [Xianyi Zhang](#)

# Mineral-Detectors Now



UNIVERSITÀ  
DEGLI STUDI  
DI MILANO



Universität  
Zürich<sup>UZH</sup>

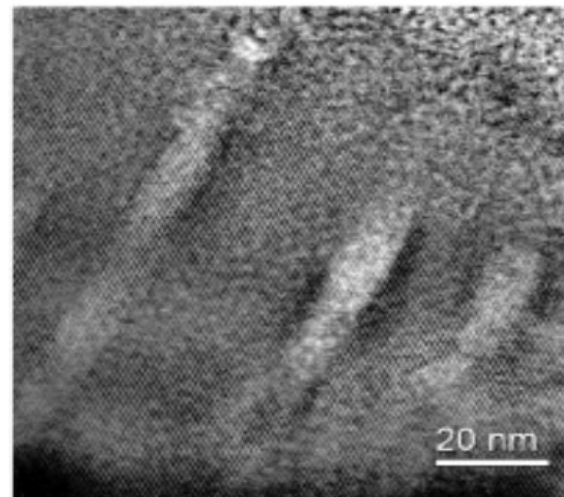
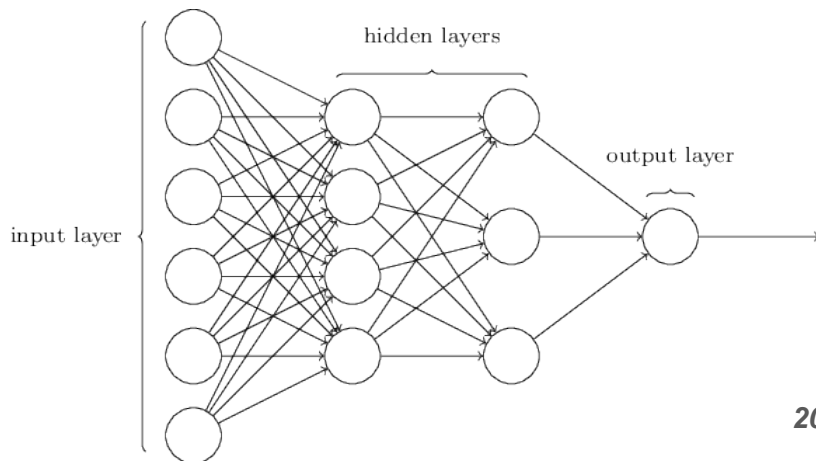


... and more ...

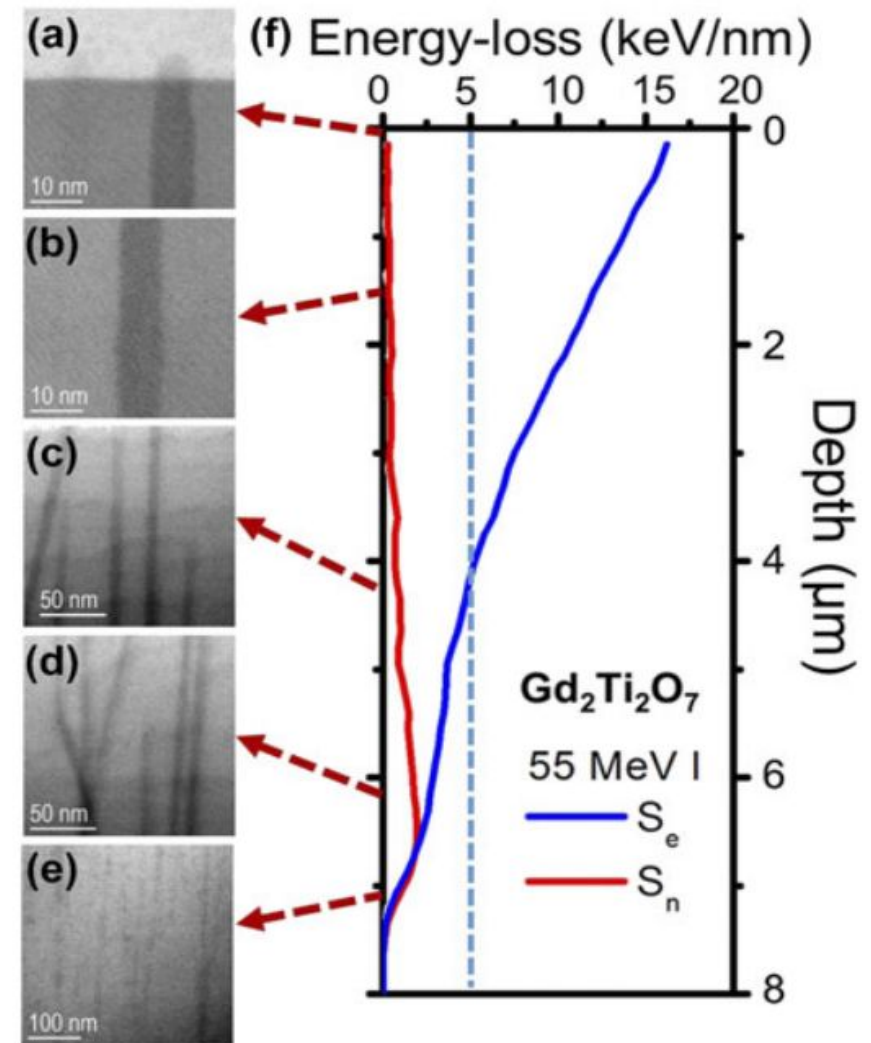
# Mineral-Detectors - Why Now?

## ❖ Renewed interest worldwide

- Unprecedented advances in nm-scale microscopy & manipulation techniques
- Computational advances - simulations, data processing
- Machine learning



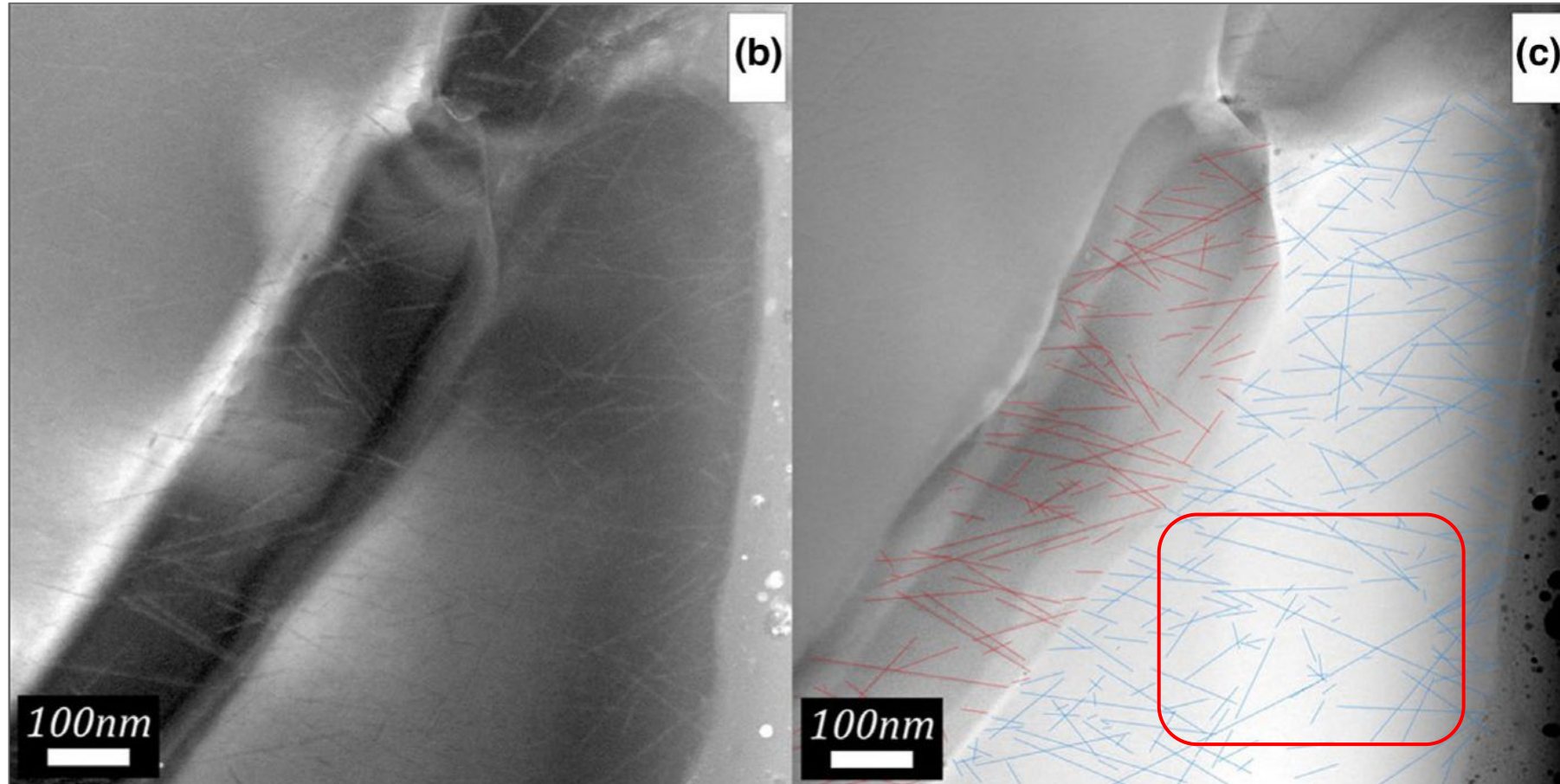
2011 : TEM imaged tracks in apatite from 2.2 GeV Au ions  
(DOI: 10.1103/PhysRevB.83.064116)



2016 : Ion track morphology at different depths in the material. (DOI: 10.1557/jmr.2016.418)

# Mineral-Detectors - Why Now?

Apollo 16 - Lunar sample



2021 : STEM images from a lunar sample. Solar energetic particle induced tracks are present in olivine and plagioclase. (b) STEM images. (c) Tracks are highlighted in red and blue for the olivine and plagioclase grains, respectively. (DOI: 10.1111/maps.13732)

# Ancient Natural Crystals - Paleo-Detectors

## Ancient minerals - look into the past

- ❖ Natural minerals - good SSNTDs
- ❖ **Tracks** - nuclear recoils induced by Dark Matter & Neutrinos
- ❖ Preserve tracks for Myr/Gyr
- ❖ Accessible, relatively cheap
  
- ❖ Small samples but Myr/Gyr exposure
- ❖ Neutrinos - guaranteed signal/background





# Ancient Natural Crystals - Paleo-Detectors

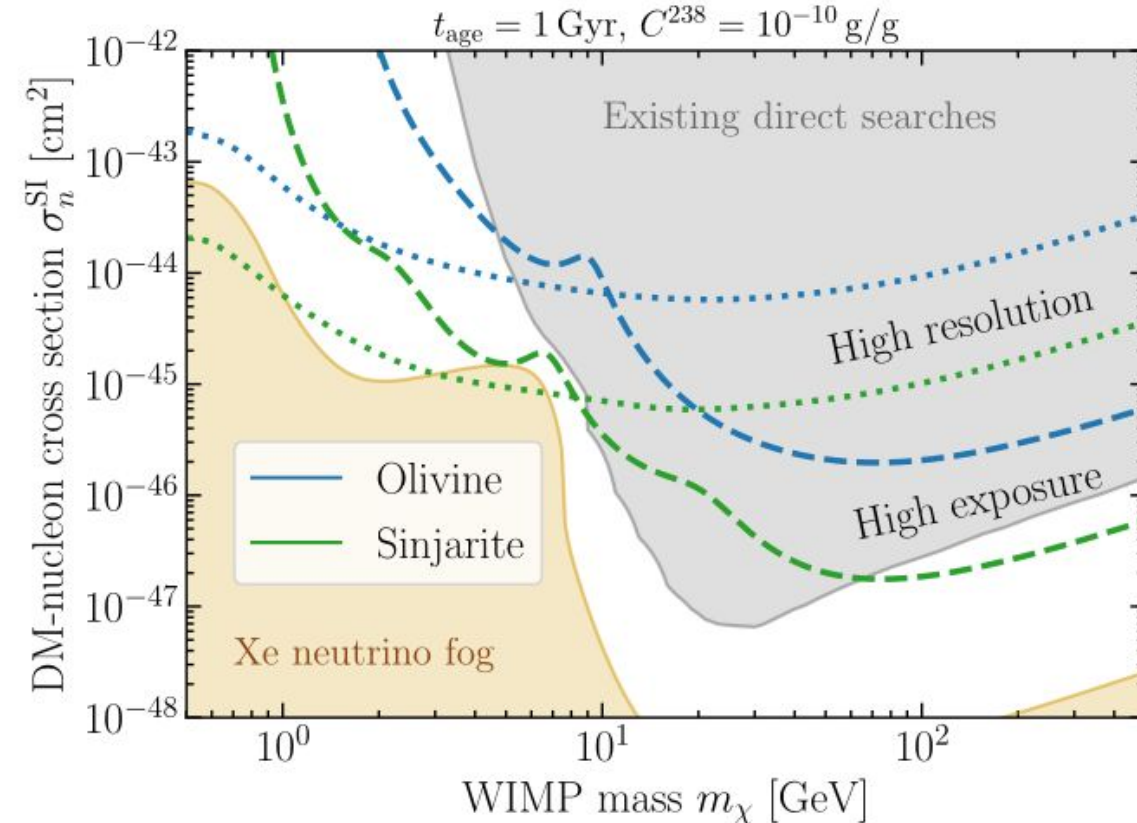
## Ancient minerals - look into the past

- ❖ Natural minerals - good SSNTDs
- ❖ **Tracks** - nuclear recoils induced by Dark Matter & Neutrinos
- ❖ Preserve tracks for Myr/Gyr
- ❖ Accessible, relatively cheap
  
- ❖ **Small samples** but **Myr/Gyr exposure**
- ❖ **Neutrinos** - guaranteed signal/background



# Excellent Dark Matter Discovery Reach

- ❖ Nuclear recoil energy thresholds down to 0.1 - 1 keV
- ❖ Mineral, readout method & resolution dependent
- ❖ Leverage high-exposure or/and high-resolution
  - Probe large range of Dark Matter candidates
- ❖ Competitive & complementary to large-scale detectors



Two scenarios : High resolution ( $\sigma_x = 1 \text{ nm}, M_{\text{sample}} = 10 \text{ mg}$ , dotted lines),  
 High exposure ( $\sigma_x = 15 \text{ nm}, M_{\text{sample}} = 100 \text{ g}$ , dashed lines).  
 The projections were produced using <https://github.com/sbaum90/paleoSens.git>

# Dark Matter Flux Variation

- ❖ Unique ability to study **time varying signals** over Myr to Gyr
  - Complementary to modern large-scale detectors
  - Dark Matter halo substructure e.g. sub-halos, “Dark Disk”
- ❖ **Smooth DM halo**
- ❖ **DM disk** - Earth would pass every ~45 Myr
- ❖ **DM subhalo** - Earth encountered during the past Gyr

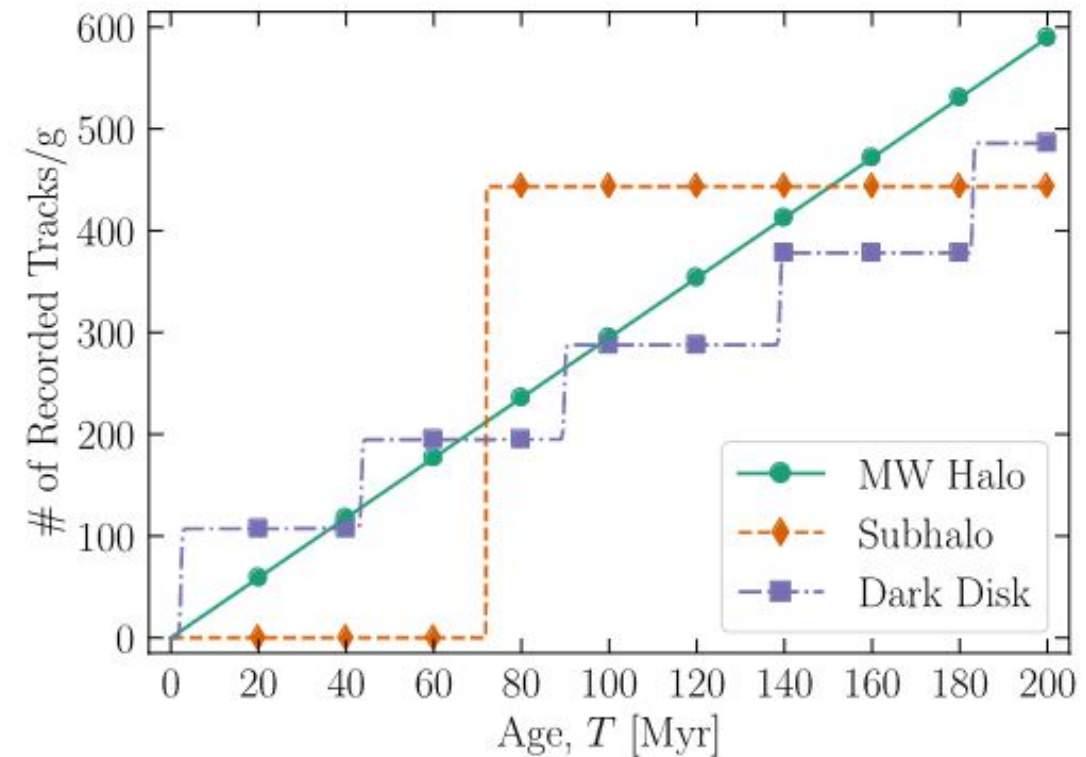


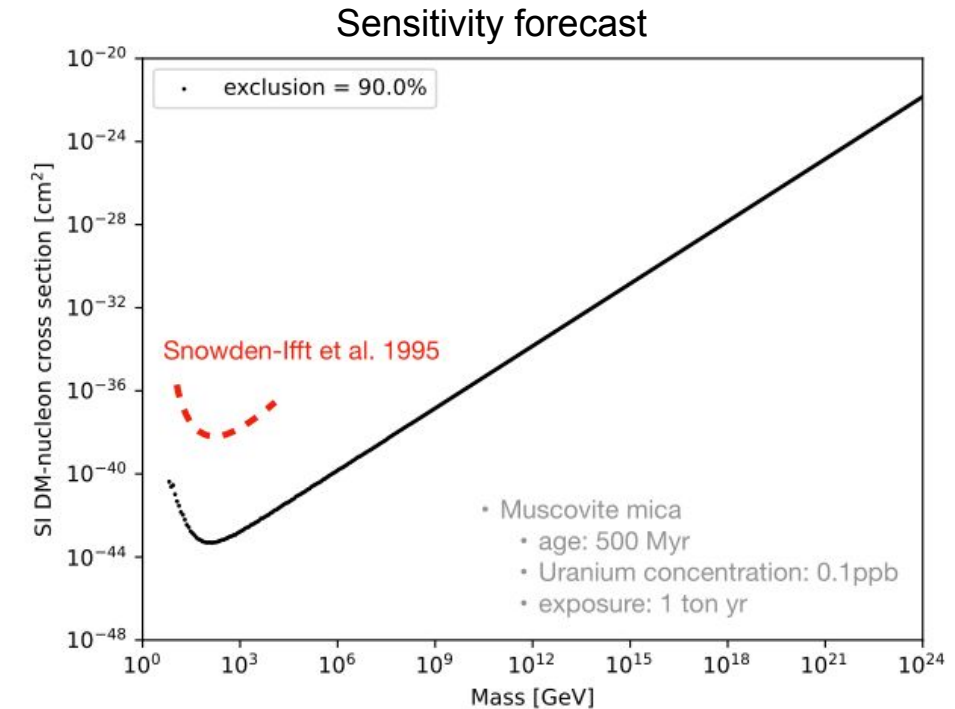
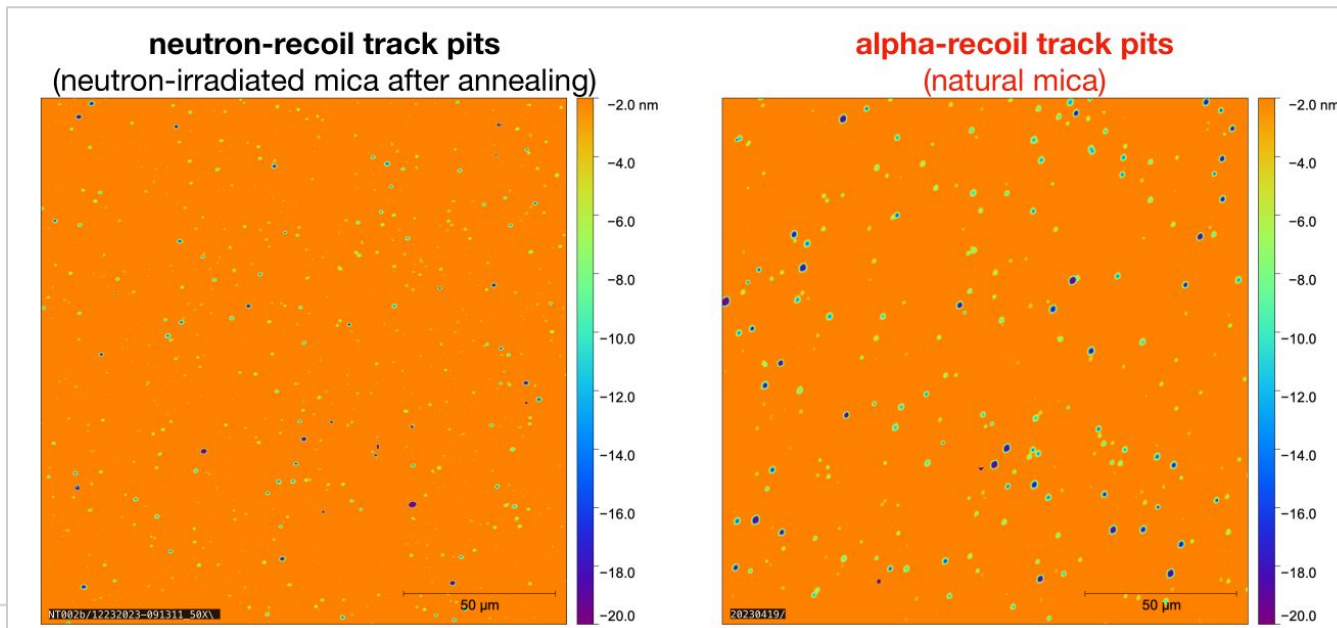
Illustration of the time-dependence of the number of damage tracks which would be recorded in a gram-sized paleo-detector of age  $T$  for three different DM signals.

# Mineral-Detectors around the World



## DMICA: exploring Dark Matter in natural muscovite MICA

- ❖ Employ methodology established by Snowden et al. (1995)
  - Chemical etching
  - Pit depth measurement optical profiler instead of AFM
  - Processed a mica of  $524,765 \mu\text{m}^2$
  - DMICA aims to scan  $\sim 1$  ton-year exposure



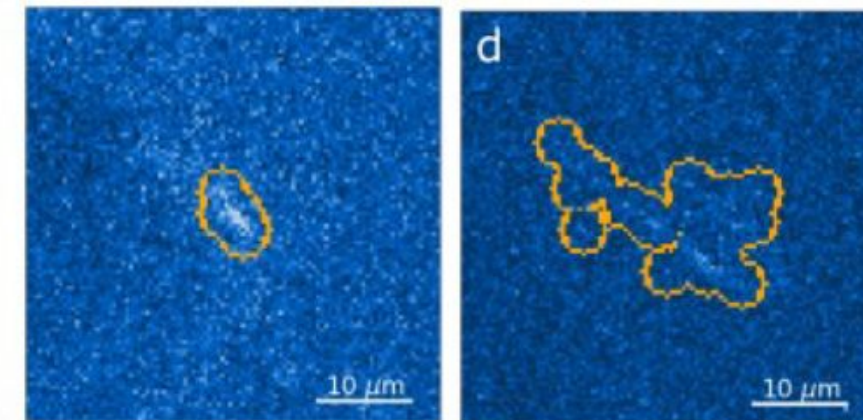
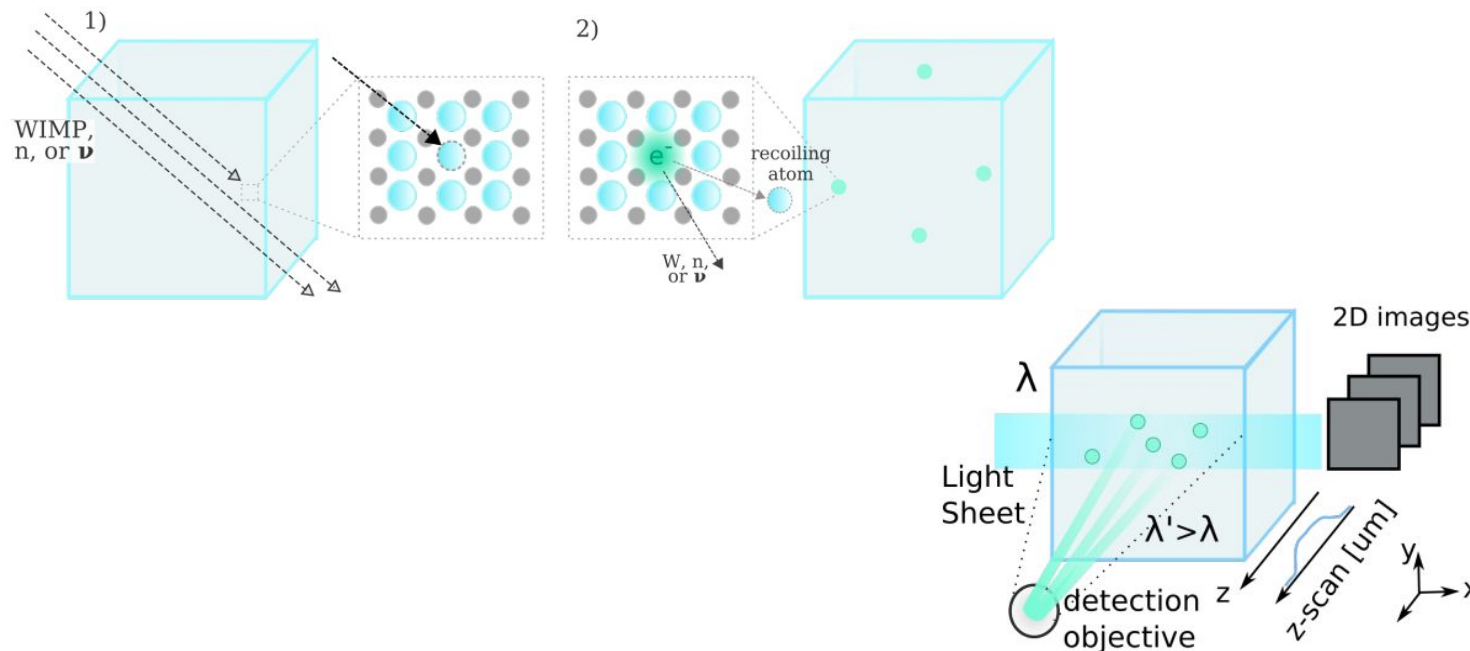
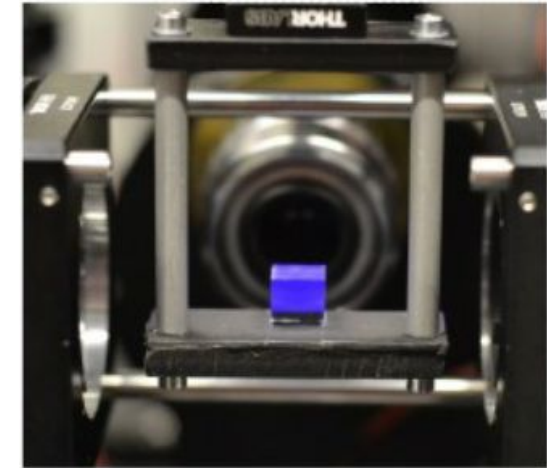
*Shigenobu Hirose et al.*

# Mineral-Detectors around the World



## ❖ Passive low energy nuclear recoil detection with color centers - PALEOCENE

- Large-scale light-sheet microscopy with mesoSPIM (mesospim.org)
- Non-destructive, resolution  $< 10 \mu\text{m}$
- Suitable crystals -  $\text{CaF}_2$ ,  $\text{LiF}$ , etc...



Gabriela R. Araujo et al.

# Mineral-Detectors at KIT



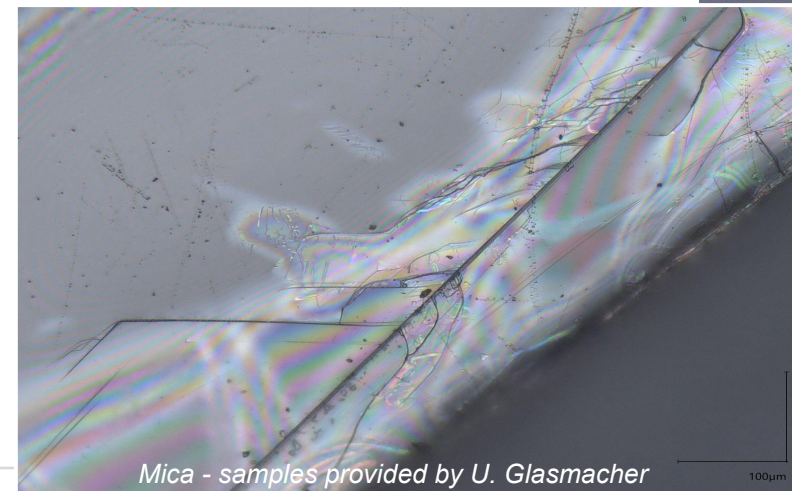
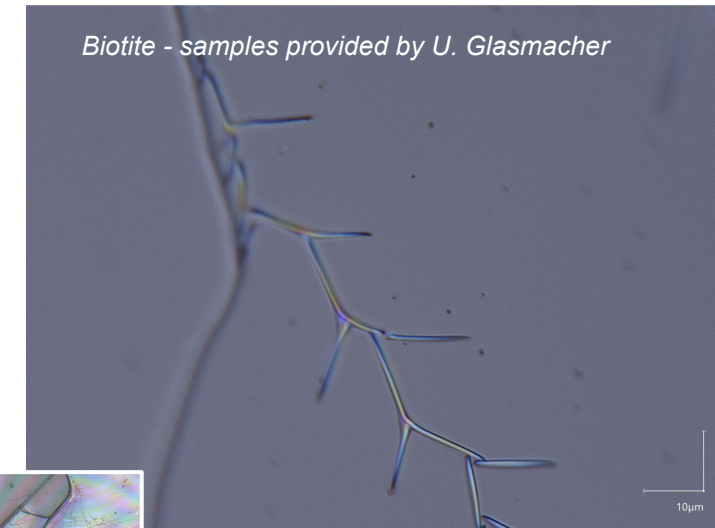
- ❖ Established connections with geologists - U. Glasmacher
- ❖ KCETA - seed funding for preliminary studies & lab setup
- ❖ Unique combination of different microscopy facilities & expertise

## ➤ Transmission Electron Microscopy (TEM)

- Resolution:  $\ll 1$  nm

## ➤ nanoCT (3D)

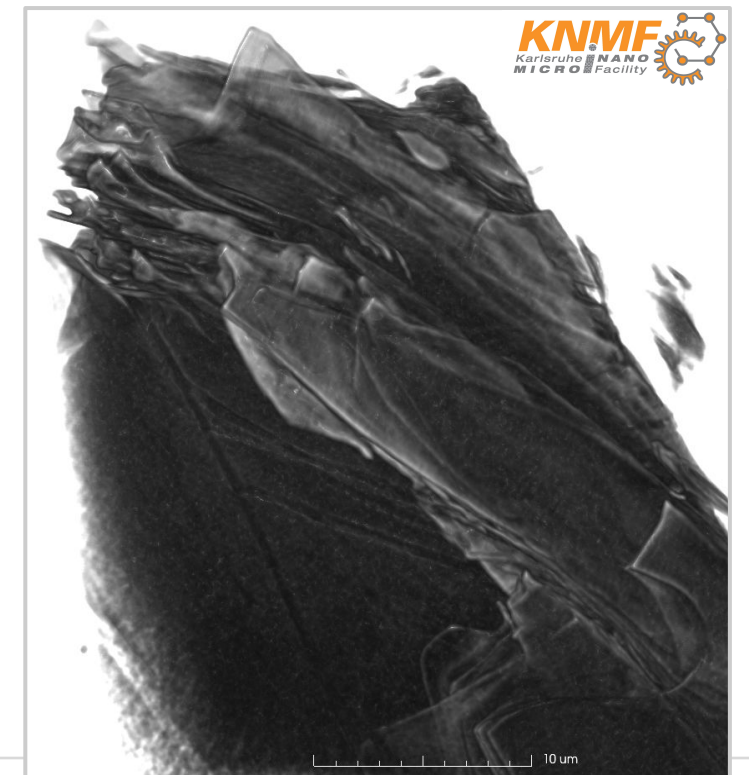
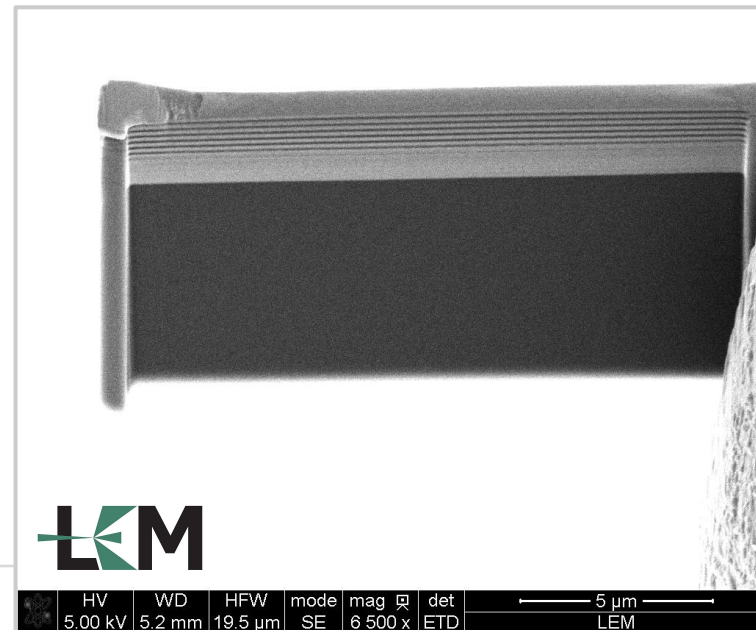
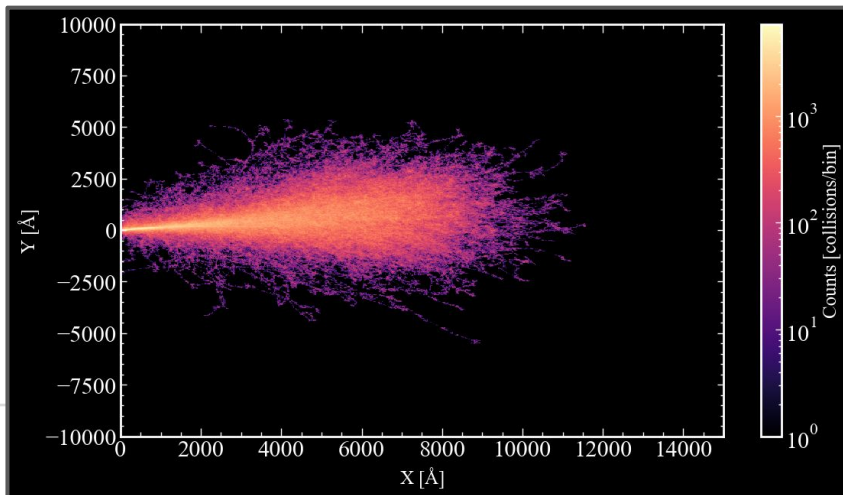
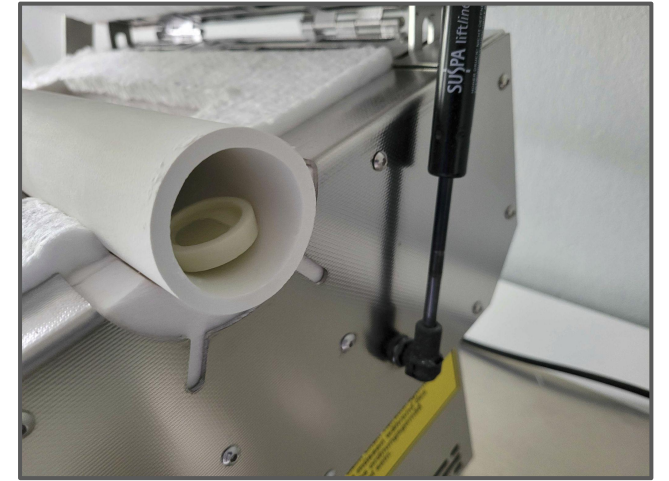
- X-ray energy: 5.4 keV
- FoV: 16  $\mu\text{m}$  (HRES), 65  $\mu\text{m}$  (LRES)
- Resolution: 50 - 100 nm
- Non-destructive



# Mineral-Detectors at KIT

## Main Goals of the Pilot Project

- ❖ Establish technology for track imaging & analysis in selected minerals
- ❖ Establish a realistic list of paleo-detector candidate minerals
- ❖ Deepen cooperation with microscopy & geology experts

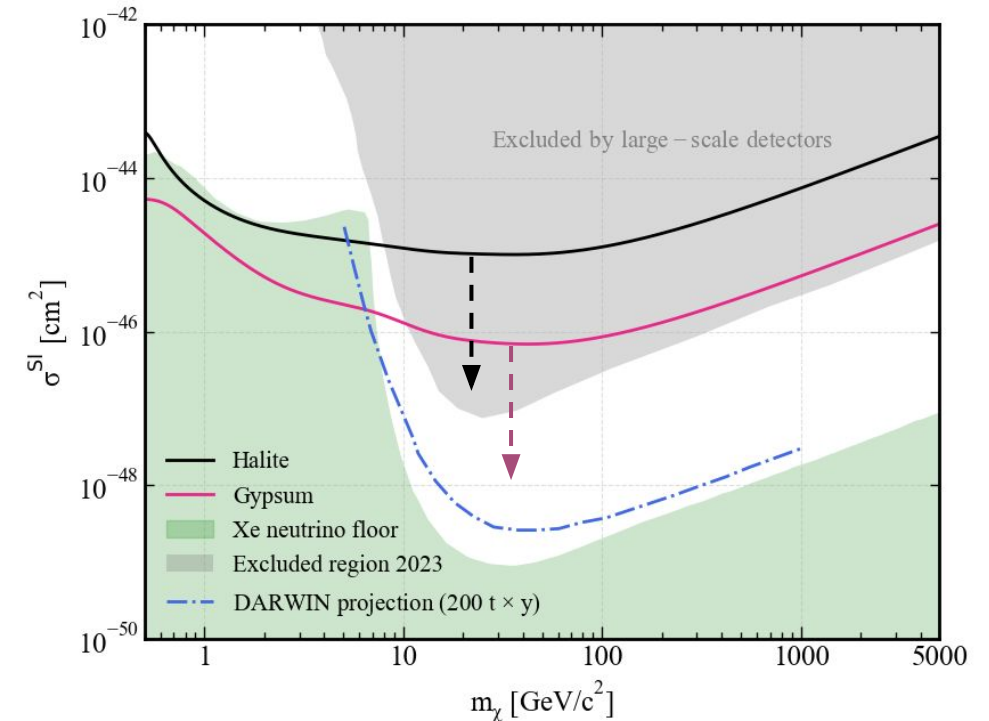


# Challenging Project

- ❖ **Suitable minerals** - not only theoretically
  - Sensitivity, attainable, chemistry, backgrounds, etc...
- ❖ **Geology** - tracks survival over Myr-Gyr?
- ❖ **Readout** & imaging techniques (< 10 nm resolution)
- ❖ **Data acquisition** & processing (~ mg samples)
- ❖ **Data analysis** - ML techniques

... It is worth it!

**Mineral-detectors may compete with large-scale experiments**



Projected WIMP Dark Matter discovery reach. The grey region is excluded by modern experiments while the green region is the so-called neutrino floor (neutrino expectation) for xenon-based detectors.  
Used : <https://github.com/sbaum90/paleoSens>



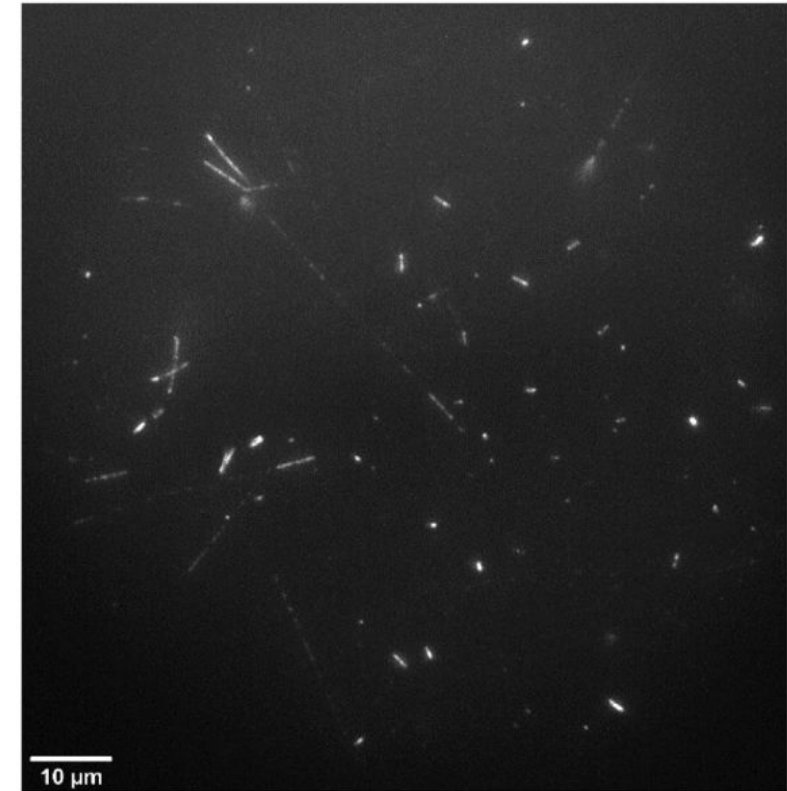
# Mineral-Detectors - Summary

## Breakthrough potential for Dark Matter & Neutrino physics

- ❖ Paleo-detectors - ancient minerals store information about nuclear recoils
  - Myr/Gyr exposure - probe of DM,  $\nu$ , cosmic rays
- ❖ Applications for “mundane” neutron/neutrino detection
- ❖ Nuclear recoils down to 0.1 - 1 keV

## Growing community & interest around the world

- ❖ **Interdisciplinary:** microscopy, geology, physics, ML & more
- ❖ **If you're interested in mineral-based detectors - contact us!**



*The experiment was already conducted by nature,  
we just need to read out the data!*

DOI: 10.1016/j.radmeas.2018.06.022